

Magnetic Monopole Searches

Isolated supermassive monopole candidate events have not been confirmed. The most sensitive experiments obtain negative results.

Best cosmic-ray supermassive monopole flux limit:

 $< \ 1.4 \times 10^{-16} \ {\rm cm}^{-2} {\rm sr}^{-1} {\rm s}^{-1} ~~ {\rm for} ~ 1.1 \times 10^{-4} < \beta < 1$

Supersymmetric Particle Searches

Presently all supersymmetric mass bounds are model dependent. This table contains a selection of bounds indicating the range of possibilities. For a more extensive set of cases consult the detailed listings.

The limits are based on the Minimal Supersymmetric Standard Model (MSSM) with additional assumptions as follows:

1) $\tilde{\chi}_1^0$ is lightest supersymmetric particle; 2) *R*-parity is conserved;

See the Particle Listings for a Note giving details of supersymmetry.

$$\begin{split} \widetilde{\chi}_{i}^{0} &- \text{neutralinos (mixtures of } \widetilde{\gamma}, \ \widetilde{Z}^{0}, \text{ and } \widetilde{H}_{i}^{0}) \\ \text{Mass } m_{\widetilde{\chi}_{1}^{0}} > 0 \text{ GeV, } \text{CL} = 95\% \\ \text{[general MSSM, non-universal gaugino masses]} \\ \text{Mass } m_{\widetilde{\chi}_{1}^{0}} > 46 \text{ GeV, } \text{CL} = 95\% \\ \text{[all } \tan\beta, \text{ all } m_{0}, \text{ all } m_{\widetilde{\chi}_{2}^{0}} - m_{\widetilde{\chi}_{1}^{0}}] \\ \text{Mass } m_{\widetilde{\chi}_{2}^{0}} > 62.4 \text{ GeV, } \text{CL} = 95\% \\ \text{[1<}\tan\beta < 40, \text{ all } m_{0}, \text{ all } m_{\widetilde{\chi}_{2}^{0}} - m_{\widetilde{\chi}_{1}^{0}}] \\ \text{Mass } m_{\widetilde{\chi}_{2}^{0}} > 345 \text{ GeV, } \text{CL} = 95\% \\ \text{[} \widetilde{\chi}_{1}^{\pm} \ \widetilde{\chi}_{2}^{0} \rightarrow \ W \ \widetilde{\chi}_{1}^{0} Z \ \widetilde{\chi}_{1}^{0}, \text{ simplified model}, \ m_{\widetilde{\chi}_{1}^{\pm}} = m_{\widetilde{\chi}_{2}^{0}}, \ m_{\widetilde{\chi}_{1}^{0}} = 0 \text{ GeV}] \end{split}$$

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Mass $m_{\widetilde{\chi}^0_3}~>~99.9$ GeV, CL =95%[1<taneta <40, all m_0 , all $m_{\widetilde{\chi}^0_2}$ - $m_{\widetilde{\chi}^0_1}$] Mass $m_{\widetilde{\chi}^0_4}~>~116$ GeV, CL =~95%[1<taneta <40, all m_0 , all $m_{\widetilde{\chi}^0_2} - m_{\widetilde{\chi}^0_1}$] $\widetilde{\chi}_i^{\pm}$ — charginos (mixtures of \widetilde{W}^{\pm} and \widetilde{H}_i^{\pm}) Mass $m_{\widetilde{\chi}^{\pm}_1}~>~$ 94 GeV, CL =~95% $[\tan\beta < 40, \ m_{\widetilde{\chi}_1^\pm} - m_{\widetilde{\chi}_1^0} > 3 \ \text{GeV, all} \ m_0]$ Mass $m_{\tilde{\chi}_1^{\pm}}$ > 345 GeV, CL = 95%[simplified model, $m_{\widetilde{\chi}_1^\pm} = m_{\widetilde{\chi}_2^0}$, $m_{\widetilde{\chi}_1^0} = 0$ GeV] $\widetilde{\nu}$ — sneutrino Mass m > 94 GeV, CL = 95%[CMSSM, $1 \leq aneta \leq 40$, $m_{\widetilde{e}_R} - m_{\widetilde{\chi}_1^0} > 10$ GeV] \tilde{e} — scalar electron (selectron) Mass $m(\widetilde{e}_L) > 107$ GeV, CL = 95% [all $m_{\widetilde{e}_R} - m_{\widetilde{\chi}_1^0}$] Mass $m(\widetilde{e}_R) > 97.5$ GeV, CL = 95% $[\Delta m > 11 \text{ GeV}, |\mu| > 100 \text{ GeV}, \tan\beta = 1.5]$ $\widetilde{\mu}$ — scalar muon (smuon) Mass m > 94 GeV, CL = 95%[CMSSM, $1 \leq aneta \leq$ 40, $m_{\widetilde{\mu}_R} - m_{\widetilde{\chi}^0_2} > 10 \ {
m GeV}$] $\widetilde{\tau}$ — scalar tau (stau) Mass m > 81.9 GeV, CL = 95% $[m_{\widetilde{ au}_R} - m_{\widetilde{\chi}^0_1} > 15$ GeV, all $heta_ au$, B $(\widetilde{ au} o \ au \, \widetilde{\chi}^0_1) = 100\%$] \tilde{q} – squarks of the first two quark generations The first of these limits is within CMSSM with cascade decays, evaluated assuming a fixed value of the parameters μ and tan β . The first two limits assume two-generations of mass degenerate squarks (\widetilde{q}_L and \widetilde{q}_R) and gaugino mass parameters that are constrained by the unification condition at the grand unification scale. The third limit assumes a simplified model with a 100% branching ratio for the prompt decay $\widetilde{q} \rightarrow q \widetilde{\chi}_1^0$. Mass m > 1450 GeV, CL = 95%[CMSSM, $\tan\beta = 30$, $A_0 = -2\max(m_0, m_{1/2})$, $\mu > 0$] Mass m > 850 GeV, CL = 95%

Mass m > 520 GeV, CL = 95%

 $[\widetilde{q}
ightarrow q \, \widetilde{\chi}_1^0$, simplified model, single light squark, $m_{\widetilde{\chi}_1^0}=0]$

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\tilde{b} — scalar bottom (sbottom)

 \tilde{g} — gluino

The first limit assumes a simplified model with a 100% branching ratio for the prompt 3 body decay, independent of the squark mass. The second of these limits is within the CMSSM (for $m_{\widetilde{g}} \gtrsim 5 \text{ GeV}$), and includes the effects of cascade decays, evaluated assuming a fixed value of the parameters μ and $\tan\beta$. The limit assumes GUT relations between gaugino masses and the gauge couplings. The third limit is based on a combination of searches. Mass m > 1225 GeV, CL = 95% [$\widetilde{g} \rightarrow q \overline{q} \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} = 0$] Mass m > 1150 GeV, CL = 95%

 $\begin{array}{l} [{\sf CMSSM},\, {\sf tan}\beta{=}30,\, A_0{=}{-}2{\sf max}(m_0,m_{1/2}),\, \mu>0] \\ {\sf Mass}\,\,m>\,\,1150\,\,{\sf GeV},\,\,{\sf CL}\,=95\% \\ [{\sf general}\,\,{\sf RPC}\,\,\widetilde{g}\,\,{\sf decays},\,m_{\widetilde{\chi}^0_1}\,\,<\,100\,\,{\sf GeV}] \end{array}$

Technicolor

The limits for technicolor (and top-color) particles are quite varied depending on assumptions. See the Technicolor section of the full *Review* (the data listings).

Quark and Lepton Compositeness, Searches for

Scale Limits Λ for Contact Interactions (the lowest dimensional interactions with four fermions)

If the Lagrangian has the form

$$\pm \frac{g^2}{2\Lambda^2} \,\overline{\psi}_L \gamma_\mu \psi_L \overline{\psi}_L \gamma^\mu \psi_L$$

(with $g^2/4\pi$ set equal to 1), then we define $\Lambda \equiv \Lambda_{LL}^{\pm}$. For the full definitions and for other forms, see the Note in the Listings on Searches for Quark and Lepton Compositeness in the full *Review* and the original literature.

$\Lambda^+_{LL}(eeee)$	>~ 8.3 TeV, CL $=~$ 95%
$\Lambda_{LL}^{-}(eeee)$	>~ 10.3 TeV, CL = 95%
$\Lambda^+_{LL}(ee\mu\mu)$	>~ 8.5 TeV, CL $=~$ 95%
$\Lambda^{LL}(ee\mu\mu)$	>~ 9.5 TeV, CL $=~$ 95%
$\Lambda^+_{LL}(ee au au)$	>~ 7.9 TeV, CL $=~$ 95%
$\Lambda_{LL}^{-}(ee au au)$	>~ 7.2 TeV, CL $=~$ 95%
$\Lambda^+_{LL}(\ell\ell\ell\ell)$	>~ 9.1 TeV, CL $=~$ 95%
$\Lambda^{-}_{LL}(\ell\ell\ell\ell)$	>~10.3 TeV, CL $= 95%$
$\Lambda^+_{LL}(eeuu)$	>~ 23.3 TeV, CL = 95%
$\Lambda_{LL}^{-}(eeuu)$	>~ 12.5 TeV, CL $=95%$
$\Lambda_{LL}^+(eedd)$	>~11.1 TeV, CL = 95%
$\Lambda_{LL}^{-}(eedd)$	>~ 26.4 TeV, CL = 95%
$\Lambda_{LL}^+(eecc)$	>~ 9.4 TeV, CL $=~$ 95%
$\Lambda_{LL}^{-}(eecc)$	>~ 5.6 TeV, CL $=~$ 95%
$\Lambda_{LL}^+(eebb)$	>~ 9.4 TeV, CL $=~$ 95%
$\Lambda_{LL}^{-}(eebb)$	>~ 10.2 TeV, CL = 95%
$\Lambda^+_{LL}(\mu\mu q q)$	>~ 12.5 TeV, CL = 95%
$\Lambda_{LL}^-(\mu\mu q q)$	>~16.7 TeV, CL $=95%$
$\Lambda(\ell\nu\ell\nu)$	>~3.10 TeV, CL = 90% $>~2.81$ TeV, CL = 95%
$\Lambda(e\nu q q)$	> 2.81 TeV, CL = 95%
$\Lambda_{LL}^+(q q q q)$	>~ 9.0 TeV, CL $=~$ 95%
$\Lambda_{LL}^{-}(q q q q)$	>~ 12.0 TeV, CL $=~$ 95%
$\Lambda_{LL}^+(\nu\nu q q)$	>~ 5.0 TeV, CL $=~$ 95%
$\Lambda^{LL}(\nu\nu q q)$	>~5.4 TeV, CL $=~95%$

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Excited Leptons

The limits from $\ell^{*+}\ell^{*-}$ do not depend on λ (where λ is the $\ell\ell^{*}$ transition coupling). The λ -dependent limits assume chiral coupling. e^{\pm} — excited electron Mass m > 103.2 GeV, CL = 95% (from e^*e^*) Mass $m > 3.000 \times 10^3$ GeV, CL = 95% (from ee^*) Mass m > 356 GeV, CL = 95% (if $\lambda_{\gamma} = 1$) $\mu^{*\pm}$ — excited muon Mass m > 103.2 GeV, CL = 95% (from $\mu^* \mu^*$) Mass $m > 3.000 \times 10^3$ GeV, CL = 95% (from $\mu \mu^*$) $\tau^{*\pm}$ — excited tau Mass m > 103.2 GeV, CL = 95% (from $\tau^* \tau^*$) Mass $m > 2.500 \times 10^3$ GeV, CL = 95% (from $\tau \tau^*$) ν^* — excited neutrino Mass $m > 1.600 \times 10^3$ GeV, CL = 95% (from $\nu^* \nu^*$) Mass m > 213 GeV, CL = 95% (from $\nu^* X$) a^* — excited quark Mass m > 338 GeV, CL = 95% (from $q^* q^*$) Mass $m > 4.060 \times 10^3$ GeV, CL = 95% (from q^*X) Color Sextet and Octet Particles Color Sextet Quarks (q_6) Mass m > 84 GeV, CL = 95% (Stable q_6) Color Octet Charged Leptons (ℓ_8) Mass m > 86 GeV, CL = 95% (Stable ℓ_8) Color Octet Neutrinos (ν_8) Mass m > 110 GeV, CL = 90% $(\nu_8 \rightarrow \nu g)$

Extra Dimensions

Please refer to the Extra Dimensions section of the full *Review* for a discussion of the model-dependence of these bounds, and further constraints.

Constraints on the radius of the extra dimensions, for the case of two-flat dimensions of equal radii

 $R < 30 \ \mu$ m, CL = 95% (direct tests of Newton's law) $R < 15 \ \mu$ m, CL = 95% ($pp \rightarrow jG$) R < 0.16–916 nm (astrophysics; limits depend on technique and assumptions)

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Constraints on the fundamental gravity scale

 $M_{TT} > 6.3$ TeV, CL = 95% ($pp \rightarrow dijet$, angular distribution) $M_c > 4.16$ TeV, CL = 95% ($pp \rightarrow \ell \overline{\ell}$)

Constraints on the Kaluza-Klein graviton in warped extra dimensions

 $M_G~>~2.73$ TeV, CL =95% ($p\,p
ightarrow~e^+\,e^-$, $~\mu^+\,\mu^-$)

Constraints on the Kaluza-Klein gluon in warped extra dimensions

 $M_{g_{KK}} > 2.5 \text{ TeV}, \text{ CL} = 95\% \quad (g_{KK} \rightarrow t \overline{t})$