

# $\rho(1700)$

$$I^G(J^{PC}) = 1^+(1^{--})$$

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## $\rho(1700)$ MASS

### $\eta\rho^0$ AND $\pi^+\pi^-$ MODES

VALUE (MeV)	DOCUMENT ID
<b>1720±20 OUR ESTIMATE</b>	

### $\eta\rho^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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The data in this block is included in the average printed for a previous datablock.

• • • We do not use the following data for averages, fits, limits, etc. • • •

1740±20	ANTONELLI	88	DM2	e <sup>+</sup> e <sup>-</sup> → ηπ <sup>+</sup> π <sup>-</sup>
1701±15	<sup>1</sup> FUKUI	88	SPEC	8.95 π <sup>-</sup> p → ηπ <sup>+</sup> π <sup>-</sup> n

<sup>1</sup> Assuming ρ<sup>+</sup>f<sub>0</sub>(1370) decay mode interferes with a<sub>1</sub>(1260)<sup>+</sup>π background. From a two Breit-Wigner fit.

### ππ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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The data in this block is included in the average printed for a previous datablock.

• • • We do not use the following data for averages, fits, limits, etc. • • •

1780 ±20 <sup>+15</sup> / <sub>-20</sub>	63.5k	<sup>2</sup> ABRAMOWICZ12	ZEUS	ep → eπ <sup>+</sup> π <sup>-</sup> p
1861 ±17		<sup>3</sup> LEES	12G	BABR e <sup>+</sup> e <sup>-</sup> → π <sup>+</sup> π <sup>-</sup> γ
1728 ±17 ±89	5.4M	<sup>4,5</sup> FUJIKAWA	08	BELL τ <sup>-</sup> → π <sup>-</sup> π <sup>0</sup> ν <sub>τ</sub>
1780 <sup>+37</sup> / <sub>-29</sub>		<sup>6</sup> ABELE	97	CBAR p̄n → π <sup>-</sup> π <sup>0</sup> π <sup>0</sup>
1719 ±15		<sup>6</sup> BERTIN	97C	OBLX 0.0 p̄p → π <sup>+</sup> π <sup>-</sup> π <sup>0</sup>
1730 ±30		CLEGG	94	RVUE e <sup>+</sup> e <sup>-</sup> → π <sup>+</sup> π <sup>-</sup>
1768 ±21		BISELLO	89	DM2 e <sup>+</sup> e <sup>-</sup> → π <sup>+</sup> π <sup>-</sup>
1745.7±91.9		DUBNICKA	89	RVUE e <sup>+</sup> e <sup>-</sup> → π <sup>+</sup> π <sup>-</sup>
1546 ±26		GESHKEN...	89	RVUE
1650		<sup>7</sup> ERKAL	85	RVUE 20–70 γp → γπ
1550 ±70		ABE	84B	HYBR 20 γp → π <sup>+</sup> π <sup>-</sup> p
1590 ±20		<sup>8</sup> ASTON	80	OMEG 20–70 γp → p2π
1600 ±10		<sup>9</sup> ATIYA	79B	SPEC 50 γC → C2π
1598 <sup>+24</sup> / <sub>-22</sub>		BECKER	79	ASPK 17 π <sup>-</sup> p polarized
1659 ±25		<sup>7</sup> LANG	79	RVUE
1575		<sup>7</sup> MARTIN	78C	RVUE 17 π <sup>-</sup> p → π <sup>+</sup> π <sup>-</sup> n
1610 ±30		<sup>7</sup> FROGGATT	77	RVUE 17 π <sup>-</sup> p → π <sup>+</sup> π <sup>-</sup> n
1590 ±20		<sup>10</sup> HYAMS	73	ASPK 17 π <sup>-</sup> p → π <sup>+</sup> π <sup>-</sup> n

<sup>2</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho$ - $\omega$  interference.

<sup>3</sup> Using the GOUNARIS 68 parametrization of the pion form factor leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.

<sup>4</sup>  $|F_\pi(0)|^2$  fixed to 1.

<sup>5</sup> From the GOUNARIS 68 parametrization of the pion form factor.

<sup>6</sup> T-matrix pole.

<sup>7</sup> From phase shift analysis of HYAMS 73 data.

<sup>8</sup> Simple relativistic Breit-Wigner fit with constant width.

<sup>9</sup> An additional 40 MeV uncertainty in both the mass and width is present due to the choice of the background shape.

<sup>10</sup> Included in BECKER 79 analysis.

## $\pi\omega$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1550 to 1620	<sup>11</sup> ACHASOV	00I	SND $e^+e^- \rightarrow \pi^0\pi^0\gamma$
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1580 to 1710	<sup>12</sup> ACHASOV	00I	SND $e^+e^- \rightarrow \pi^0\pi^0\gamma$
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1710 $\pm$ 90	ACHASOV	97	RVUE $e^+e^- \rightarrow \omega\pi^0$
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<sup>11</sup> Taking into account both  $\rho(1450)$  and  $\rho(1700)$  contributions. Using the data of ACHASOV 00I on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .  $\rho(1450)$  mass and width fixed at 1400 MeV and 500 MeV respectively.

<sup>12</sup> Taking into account the  $\rho(1700)$  contribution only. Using the data of ACHASOV 00I on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .

## $K\bar{K}$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1740.8 $\pm$ 22.2	27k	<sup>13</sup> ABELE	99D	CBAR $\pm$	0.0 $\bar{p}p \rightarrow K^+K^-\pi^0$
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1582 $\pm$ 36	1600	CLELAND	82B	SPEC $\pm$	50 $\pi p \rightarrow K_S^0 K^\pm p$
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<sup>13</sup> K-matrix pole. Isospin not determined, could be  $\omega(1650)$  or  $\phi(1680)$ .

## $2(\pi^+\pi^-)$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1851 $^{+27}_{-24}$		ACHASOV	97	RVUE $e^+e^- \rightarrow 2(\pi^+\pi^-)$
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1570 $\pm$ 20		<sup>14</sup> CORDIER	82	DM1 $e^+e^- \rightarrow 2(\pi^+\pi^-)$
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1520 $\pm$ 30		<sup>15</sup> ASTON	81E	OMEG $20-70 \gamma p \rightarrow p4\pi$
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1654 $\pm$ 25		<sup>16</sup> DIBIANCA	81	DBC $\pi^+d \rightarrow pp2(\pi^+\pi^-)$
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1666 $\pm$ 39		<sup>14</sup> BACCI	80	FRAG $e^+e^- \rightarrow 2(\pi^+\pi^-)$
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1780	34	KILLIAN	80	SPEC $11 e^-p \rightarrow 2(\pi^+\pi^-)$
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1500		<sup>17</sup> ATIYA	79B	SPEC $50 \gamma C \rightarrow C4\pi^\pm$
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1570 $\pm$ 60	65	<sup>18</sup> ALEXANDER	75	HBC $7.5 \gamma p \rightarrow p4\pi$
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1550 $\pm$ 60		<sup>15</sup> CONVERSI	74	OSPK $e^+e^- \rightarrow 2(\pi^+\pi^-)$
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1550 $\pm$ 50	160	SCHACHT	74	STRC $5.5-9 \gamma p \rightarrow p4\pi$
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1450 $\pm$ 100	340	SCHACHT	74	STRC $9-18 \gamma p \rightarrow p4\pi$
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1430 $\pm$ 50	400	BINGHAM	72B	HBC $9.3 \gamma p \rightarrow p4\pi$
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<sup>14</sup> Simple relativistic Breit-Wigner fit with model dependent width.

<sup>15</sup> Simple relativistic Breit-Wigner fit with constant width.

<sup>16</sup> One peak fit result.

<sup>17</sup> Parameters roughly estimated, not from a fit.

<sup>18</sup> Skew mass distribution compensated by Ross-Stodolsky factor.

### $\pi^+\pi^-\pi^0\pi^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
$1660 \pm 30$	ATKINSON	85B	OMEG 20-70 $\gamma p$

### $3(\pi^+\pi^-)$ AND $2(\pi^+\pi^-\pi^0)$ MODES

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
$1730 \pm 34$	<sup>19</sup> FRABETTI	04	E687 $\gamma p \rightarrow 3\pi^+ 3\pi^- p$
$1783 \pm 15$	CLEGG	90	RVUE $e^+e^- \rightarrow 3(\pi^+\pi^-)2(\pi^+\pi^-\pi^0)$
<sup>19</sup> From a fit with two resonances with the JACOB 72 continuum.			

### $\rho(1700)$ WIDTH

#### $\eta\rho^0$ AND $\pi^+\pi^-$ MODES

VALUE (MeV)	DOCUMENT ID
$250 \pm 100$	OUR ESTIMATE

#### $\eta\rho^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
The data in this block is included in the average printed for a previous datablock.			

● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
$150 \pm 30$	ANTONELLI	88	DM2 $e^+e^- \rightarrow \eta\pi^+\pi^-$
$282 \pm 44$	<sup>20</sup> FUKUI	88	SPEC $8.95 \pi^- p \rightarrow \eta\pi^+\pi^- n$
<sup>20</sup> Assuming $\rho^+ f_0(1370)$ decay mode interferes with $a_1(1260)^+\pi$ background. From a two Breit-Wigner fit.			

#### $\pi\pi$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
The data in this block is included in the average printed for a previous datablock.				

● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$310 \pm 30$	$+25$ $-35$	63.5k	<sup>21</sup> ABRAMOWICZ12	ZEUS $ep \rightarrow e\pi^+\pi^-p$
$316 \pm 26$			<sup>22</sup> LEES	12G BABR $e^+e^- \rightarrow \pi^+\pi^-\gamma$
$164 \pm 21$	$+89$ $-26$	5.4M	<sup>23,24</sup> FUJIKAWA	08 BELL $\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
$275 \pm 45$			<sup>25</sup> ABELE	97 CBAR $\bar{p}n \rightarrow \pi^-\pi^0\pi^0$
$310 \pm 40$			<sup>25</sup> BERTIN	97C OBLX $0.0 \bar{p}p \rightarrow \pi^+\pi^-\pi^0$
$400 \pm 100$			CLEGG	94 RVUE $e^+e^- \rightarrow \pi^+\pi^-$
$224 \pm 22$			BISELLO	89 DM2 $e^+e^- \rightarrow \pi^+\pi^-$
$242.5 \pm 163.0$			DUBNICKA	89 RVUE $e^+e^- \rightarrow \pi^+\pi^-$
$620 \pm 60$			GESHKEN...	89 RVUE
$<315$			<sup>26</sup> ERKAL	85 RVUE 20-70 $\gamma p \rightarrow \gamma\pi$
$280 \pm 30$ $-80$			ABE	84B HYBR 20 $\gamma p \rightarrow \pi^+\pi^-p$

230 ± 80	27 ASTON	80 OMEG	20–70 $\gamma p \rightarrow p 2\pi$
283 ± 14	28 ATIYA	79B SPEC	50 $\gamma C \rightarrow C 2\pi$
175 + 98 – 53	BECKER	79 ASPK	17 $\pi^- p$ polarized
232 ± 34	26 LANG	79 RVUE	
340	26 MARTIN	78C RVUE	17 $\pi^- p \rightarrow \pi^+ \pi^- n$
300 ± 100	26 FROGGATT	77 RVUE	17 $\pi^- p \rightarrow \pi^+ \pi^- n$
180 ± 50	29 HYAMS	73 ASPK	17 $\pi^- p \rightarrow \pi^+ \pi^- n$

<sup>21</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho$ - $\omega$  interference.

<sup>22</sup> Using the GOUNARIS 68 parametrization of the pion form factor leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.

<sup>23</sup>  $|F_\pi(0)|^2$  fixed to 1.

<sup>24</sup> From the GOUNARIS 68 parametrization of the pion form factor.

<sup>25</sup> T-matrix pole.

<sup>26</sup> From phase shift analysis of HYAMS 73 data.

<sup>27</sup> Simple relativistic Breit-Wigner fit with constant width.

<sup>28</sup> An additional 40 MeV uncertainty in both the mass and width is present due to the choice of the background shape.

<sup>29</sup> Included in BECKER 79 analysis.

## $K\bar{K}$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

187.2 ± 26.7	27k	<sup>30</sup> ABELE	99D CBAR	±	0.0 $\bar{p} p \rightarrow K^+ K^- \pi^0$
265 ± 120	1600	CLELAND	82B SPEC	±	50 $\pi p \rightarrow K_S^0 K^\pm p$

<sup>30</sup> K-matrix pole. Isospin not determined, could be  $\omega(1650)$  or  $\phi(1680)$ .

## $2(\pi^+\pi^-)$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

510 ± 40		<sup>31</sup> CORDIER	82 DM1	$e^+ e^- \rightarrow 2(\pi^+ \pi^-)$
400 ± 50		<sup>32</sup> ASTON	81E OMEG	20–70 $\gamma p \rightarrow p 4\pi$
400 ± 146		<sup>33</sup> DIBIANCA	81 DBC	$\pi^+ d \rightarrow p p 2(\pi^+ \pi^-)$
700 ± 160		<sup>31</sup> BACCI	80 FRAG	$e^+ e^- \rightarrow 2(\pi^+ \pi^-)$
100	34	KILLIAN	80 SPEC	11 $e^- p \rightarrow 2(\pi^+ \pi^-)$
600		<sup>34</sup> ATIYA	79B SPEC	50 $\gamma C \rightarrow C 4\pi^\pm$
340 ± 160	65	<sup>35</sup> ALEXANDER	75 HBC	7.5 $\gamma p \rightarrow p 4\pi$
360 ± 100		<sup>32</sup> CONVERSI	74 OSPK	$e^+ e^- \rightarrow 2(\pi^+ \pi^-)$
400 ± 120	160	<sup>36</sup> SCHACHT	74 STRC	5.5–9 $\gamma p \rightarrow p 4\pi$
850 ± 200	340	<sup>36</sup> SCHACHT	74 STRC	9–18 $\gamma p \rightarrow p 4\pi$
650 ± 100	400	BINGHAM	72B HBC	9.3 $\gamma p \rightarrow p 4\pi$

<sup>31</sup> Simple relativistic Breit-Wigner fit with model-dependent width.

<sup>32</sup> Simple relativistic Breit-Wigner fit with constant width.

<sup>33</sup> One peak fit result.

<sup>34</sup> Parameters roughly estimated, not from a fit.

<sup>35</sup> Skew mass distribution compensated by Ross-Stodolsky factor.

<sup>36</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

### $\pi^+\pi^-\pi^0\pi^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$300 \pm 50$	ATKINSON	85B	OMEG $20-70 \gamma p$

### $\omega\pi^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
350 to 580	<sup>37</sup> ACHASOV	00i	SND $e^+e^- \rightarrow \pi^0\pi^0\gamma$
490 to 1040	<sup>38</sup> ACHASOV	00i	SND $e^+e^- \rightarrow \pi^0\pi^0\gamma$

<sup>37</sup>Taking into account both  $\rho(1450)$  and  $\rho(1700)$  contributions. Using the data of ACHASOV 00i on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .  $\rho(1450)$  mass and width fixed at 1400 MeV and 500 MeV respectively.

<sup>38</sup>Taking into account the  $\rho(1700)$  contribution only. Using the data of ACHASOV 00i on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .

### $3(\pi^+\pi^-)$ AND $2(\pi^+\pi^-\pi^0)$ MODES

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$315 \pm 100$	<sup>39</sup> FRABETTI	04	E687 $\gamma p \rightarrow 3\pi^+3\pi^-\rho$
$285 \pm 20$	CLEGG	90	RVUE $e^+e^- \rightarrow 3(\pi^+\pi^-)2(\pi^+\pi^-\pi^0)$

<sup>39</sup>From a fit with two resonances with the JACOB 72 continuum.

### $\rho(1700)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $4\pi$	
$\Gamma_2$ $2(\pi^+\pi^-)$	large
$\Gamma_3$ $\rho\pi\pi$	dominant
$\Gamma_4$ $\rho^0\pi^+\pi^-$	large
$\Gamma_5$ $\rho^0\pi^0\pi^0$	
$\Gamma_6$ $\rho^\pm\pi^\mp\pi^0$	large
$\Gamma_7$ $a_1(1260)\pi$	seen
$\Gamma_8$ $h_1(1170)\pi$	seen
$\Gamma_9$ $\pi(1300)\pi$	seen
$\Gamma_{10}$ $\rho\rho$	seen
$\Gamma_{11}$ $\pi^+\pi^-$	seen
$\Gamma_{12}$ $\pi\pi$	seen
$\Gamma_{13}$ $K\bar{K}^*(892) + \text{c.c.}$	seen
$\Gamma_{14}$ $\eta\rho$	seen
$\Gamma_{15}$ $a_2(1320)\pi$	not seen
$\Gamma_{16}$ $K\bar{K}$	seen
$\Gamma_{17}$ $e^+e^-$	seen
$\Gamma_{18}$ $\pi^0\omega$	seen

## $\rho(1700) \Gamma(i)\Gamma(e^+e^-)/\Gamma(\text{total})$

This combination of a partial width with the partial width into  $e^+e^-$  and with the total width is obtained from the cross-section into channel<sub>l</sub> in  $e^+e^-$  annihilation.

### $\Gamma(2(\pi^+\pi^-)) \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_2\Gamma_{17}/\Gamma$

<u>VALUE (keV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
2.6 ± 0.2	DELCOURT	81B DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
2.83 ± 0.42	BACCI	80 FRAG	$e^+e^- \rightarrow 2(\pi^+\pi^-)$

### $\Gamma(\pi^+\pi^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_{11}\Gamma_{17}/\Gamma$

<u>VALUE (keV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
0.13	<sup>40</sup> DIEKMAN	88 RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
0.029 <sup>+0.016</sup> <sub>-0.012</sub>	KURDADZE	83 OLYA	0.64–1.4 $e^+e^- \rightarrow \pi^+\pi^-$

<sup>40</sup> Using total width = 220 MeV.

### $\Gamma(K\bar{K}^*(892)+\text{c.c.}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_{13}\Gamma_{17}/\Gamma$

<u>VALUE (keV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
0.305 ± 0.071	<sup>41</sup> BIZOT	80 DM1	$e^+e^-$

<sup>41</sup> Model dependent.

### $\Gamma(\eta\rho) \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_{14}\Gamma_{17}/\Gamma$

<u>VALUE (eV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
7 ± 3	ANTONELLI	88 DM2	$e^+e^- \rightarrow \eta\pi^+\pi^-$

### $\Gamma(K\bar{K}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_{16}\Gamma_{17}/\Gamma$

<u>VALUE (keV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
0.035 ± 0.029	<sup>42</sup> BIZOT	80 DM1	$e^+e^-$

<sup>42</sup> Model dependent.

### $\Gamma(\rho\pi\pi) \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_3\Gamma_{17}/\Gamma$

<u>VALUE (keV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
3.510 ± 0.090	<sup>43</sup> BIZOT	80 DM1	$e^+e^-$

<sup>43</sup> Model dependent.

## $\rho(1700)$ BRANCHING RATIOS

### $\Gamma(\rho\pi\pi)/\Gamma(4\pi)$ $\Gamma_3/\Gamma_1$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
0.28 ± 0.06	<sup>44</sup> ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
<sup>44</sup> $\omega\pi$ not included.			

### $\Gamma(\rho^0\pi^+\pi^-)/\Gamma(2(\pi^+\pi^-))$ $\Gamma_4/\Gamma_2$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
~ 1.0		DELCOURT	81B DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
0.7 ± 0.1	500	SCHACHT	74 STRC	5.5–18 $\gamma p \rightarrow p4\pi$
0.80		<sup>45</sup> BINGHAM	72B HBC	9.3 $\gamma p \rightarrow p4\pi$
<sup>45</sup> The $\pi\pi$ system is in $S$ -wave.				

### $\Gamma(\rho^0\pi^0\pi^0)/\Gamma(\rho^\pm\pi^\mp\pi^0)$ $\Gamma_5/\Gamma_6$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
< 0.10	ATKINSON	85B OMEG		20–70 $\gamma p$
< 0.15	ATKINSON	82 OMEG 0		20–70 $\gamma p \rightarrow p4\pi$

### $\Gamma(a_1(1260)\pi)/\Gamma(4\pi)$ $\Gamma_7/\Gamma_1$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
0.16 ± 0.05	<sup>46</sup> ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
<sup>46</sup> $\omega\pi$ not included.			

### $\Gamma(h_1(1170)\pi)/\Gamma(4\pi)$ $\Gamma_8/\Gamma_1$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
0.17 ± 0.06	<sup>47</sup> ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
<sup>47</sup> $\omega\pi$ not included.			

### $\Gamma(\pi(1300)\pi)/\Gamma(4\pi)$ $\Gamma_9/\Gamma_1$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
0.30 ± 0.10	<sup>48</sup> ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
<sup>48</sup> $\omega\pi$ not included.			

### $\Gamma(\rho\rho)/\Gamma(4\pi)$ $\Gamma_{10}/\Gamma_1$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
0.09 ± 0.03	<sup>49</sup> ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
<sup>49</sup> $\omega\pi$ not included.			

$\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{11}/\Gamma$

VALUE DOCUMENT ID TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.287^{+0.043}_{-0.042}$	BECKER	79	ASPK	17 $\pi^- p$ polarized
0.15 to 0.30	<sup>50</sup> MARTIN	78C	RVUE	17 $\pi^- p \rightarrow \pi^+\pi^- n$
<0.20	<sup>51</sup> COSTA...	77B	RVUE	$e^+e^- \rightarrow 2\pi, 4\pi$
$0.30 \pm 0.05$	<sup>50</sup> FROGGATT	77	RVUE	17 $\pi^- p \rightarrow \pi^+\pi^- n$
<0.15	<sup>52</sup> EISENBERG	73	HBC	5 $\pi^+ p \rightarrow \Delta^{++} 2\pi$
$0.25 \pm 0.05$	<sup>53</sup> HYAMS	73	ASPK	17 $\pi^- p \rightarrow \pi^+\pi^- n$

<sup>50</sup> From phase shift analysis of HYAMS 73 data.

<sup>51</sup> Estimate using unitarity, time reversal invariance, Breit-Wigner.

<sup>52</sup> Estimated using one-pion-exchange model.

<sup>53</sup> Included in BECKER 79 analysis.

$\Gamma(\pi^+\pi^-)/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{11}/\Gamma_2$

VALUE DOCUMENT ID TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.13 \pm 0.05$	ASTON	80	OMEG	20–70 $\gamma p \rightarrow p 2\pi$
<0.14	<sup>54</sup> DAVIER	73	STRC	6–18 $\gamma p \rightarrow p 4\pi$
<0.2	<sup>55</sup> BINGHAM	72B	HBC	9.3 $\gamma p \rightarrow p 2\pi$

<sup>54</sup> Upper limit is estimate.

<sup>55</sup>  $2\sigma$  upper limit.

$\Gamma(\pi\pi)/\Gamma(4\pi)$   $\Gamma_{12}/\Gamma_1$

VALUE DOCUMENT ID TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.16 \pm 0.04$	<sup>56,57</sup> ABELE	01B	CBAR	0.0 $\bar{p} n \rightarrow 5\pi$
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<sup>56</sup> Using ABELE 97.

<sup>57</sup>  $\omega\pi$  not included.

$\Gamma(K\bar{K}^*(892)+\text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{13}/\Gamma$

VALUE DOCUMENT ID TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

possibly seen	COAN	04	CLEO	$\tau^- \rightarrow K^- \pi^- K^+ \nu_\tau$
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$\Gamma(K\bar{K}^*(892)+\text{c.c.})/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{13}/\Gamma_2$

VALUE DOCUMENT ID TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.15 \pm 0.03$	<sup>58</sup> DELCOURT	81B	DM1	$e^+e^- \rightarrow \bar{K} K \pi$
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<sup>58</sup> Assuming  $\rho(1700)$  and  $\omega$  radial excitations to be degenerate in mass.

$\Gamma(\eta\rho)/\Gamma_{\text{total}}$   $\Gamma_{14}/\Gamma$

VALUE CL% DOCUMENT ID TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

possibly seen		AKHMETSHIN 00D	CMD2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
<0.04		DONNACHIE 87B	RVUE	
<0.02	58	ATKINSON 86B	OMEG	20–70 $\gamma p$

$\Gamma(\eta\rho)/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{14}/\Gamma_2$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.123±0.027	DELCOURT 82	DM1	$e^+e^- \rightarrow \pi^+\pi^- MM$
~ 0.1	ASTON 80	OMEG	20–70 $\gamma p$

$\Gamma(\pi^+\pi^- \text{ neutrals})/\Gamma(2(\pi^+\pi^-))$   $(\Gamma_5+\Gamma_6+0.714\Gamma_{14})/\Gamma_2$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.6±0.4	<sup>59</sup> BALLAM 74	HBC	9.3 $\gamma p$
<sup>59</sup> Upper limit. Background not subtracted.			

$\Gamma(a_2(1320)\pi)/\Gamma_{\text{total}}$   $\Gamma_{15}/\Gamma$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •			
not seen	AMELIN 00	VES	37 $\pi^- p \rightarrow \eta\pi^+\pi^- n$

$\Gamma(K\bar{K})/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{16}/\Gamma_2$

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.015±0.010		<sup>60</sup> DELCOURT 81B	DM1		$e^+e^- \rightarrow \bar{K}K$
<0.04	95	BINGHAM 72B	HBC	0	9.3 $\gamma p$
<sup>60</sup> Assuming $\rho(1700)$ and $\omega$ radial excitations to be degenerate in mass.					

$\Gamma(K\bar{K})/\Gamma(K\bar{K}^*(892)+c.c.)$   $\Gamma_{16}/\Gamma_{13}$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.052±0.026	BUON 82	DM1	$e^+e^- \rightarrow \text{hadrons}$

$\Gamma(\pi^0\omega)/\Gamma_{\text{total}}$   $\Gamma_{18}/\Gamma$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	1.6k	ACHASOV 12	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
not seen	2382	AKHMETSHIN 03B	CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
seen		ACHASOV 97	RVUE	$e^+e^- \rightarrow \omega\pi^0$

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ATKINSON	86B	ZPHY C30 531	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
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