graviton

J = 2

## graviton MASS

All of the following limits are obtained assuming Yukawa potential in weak field limit. VANDAM 70 argue that a massive field cannot approach general relativity in the zero-mass limit; however, see GOLD-HABER 10 and references therein.  $h_0$  is the Hubble constant in units of 100 km s<sup>-1</sup> Mpc<sup>-1</sup>.

The following conversions are useful: 1 eV =  $1.783\times10^{-33}$  g =  $1.957\times10^{-6}~m_e$ ;  $\lambda_C$  =  $1.973\times10^{-7}$  m.

VALUE (eV)	DOCUMENT ID		COMMENT		
$<7 \times 10^{-32}$	<sup>1</sup> CHOUDHURY	04	Weak gravitational lensing		
ullet $ullet$ $ullet$ We do not use the following data for averages, fits, limits, etc. $ullet$ $ullet$					
$< 7.6 \times 10^{-20}$	<sup>2</sup> FINN	02	Binary Pulsars		
	<sup>3</sup> DAMOUR	91	Binary pulsar PSR 1913+16		
$ \begin{array}{l} < 2 \times 10^{-29} \ h_0^{-1} \\ < 7 \ \times 10^{-28} \\ < 8 \ \times 10^4 \end{array} $	GOLDHABER	74	Rich clusters		
$< 7 \times 10^{-28}$	HARE	73	Galaxy		
$< 8 \times 10^{4}$	HARE	73	$2\gamma$ decay		
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<sup>1</sup>CHOUDHURY 04 sets limits based on nonobservation of a distortion in the measured values of the variance of the power spectrum.

 $^2$  FINN 02 analyze the orbital decay rates of PSR B1913+16 and PSR B1534+12 with a possible graviton mass as a parameter. The combined frequentist mass limit is at 90%CL.

<sup>3</sup>DAMOUR 91 is an analysis of the orbital period change in binary pulsar PSR 1913+16, and confirms the general relativity prediction to 0.8%. "The theoretical importance of the [rate of orbital period decay] measurement has long been recognized as a direct confirmation that the gravitational interaction propagates with velocity *c* (which is the immediate cause of the appearance of a damping force in the binary pulsar system) and thereby as a test of the existence of gravitational radiation and of its quadrupolar nature." TAYLOR 93 adds that orbital parameter studies now agree with general relativity to 0.5%, and set limits on the level of scalar contribution in the context of a family of tensor [spin 2]-biscalar theories.

## graviton REFERENCES

GOLDHABER	10	RMP 82 939	A.F. Goldhaber, M.M. Nieto	(STON, LANL)
CHOUDHURY	04	ASP 21 559	S.R. Choudhury <i>et al.</i>	(DELPH, MELB)
FINN	02	PR D65 044022	L.S. Finn, P.J. Sutton	
TAYLOR	93	NAT 355 132	J.N. Taylor <i>et al.</i>	(PRIN, ARCBO, BURE+) J
DAMOUR	91	APJ 366 501	T. Damour, J.H. Taylor	(BURE, MEUD, PRIN)
GOLDHABER	74	PR D9 1119	A.S. Goldhaber, M.M. Nieto	(LANL, STON)
HARE	73	CJP 51 431	M.G. Hare	(SASK)
VANDAM	70	NP B22 397	H. van Dam, M. Veltman	(UTRE)
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