Double- β Decay

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Half-life Measurements and Limits for Double- β Decay

In all cases of double-beta decay, $(Z,A) \rightarrow (Z+2,A) + 2e^- + (0 \text{ or } 2)\overline{\nu}_e$. In the following Listings, only best or comparable limits or lifetimes for each isotope are reported. For 2ν decay, which is well established, only measured half-lives are reported.

$t_{1/2}(10^{21} \text{ yr})$	CL%	ISOTOPE	TR	ANSITION	METHOD	DOCUMENT ID		
ullet $ullet$ $ullet$ We do not use the following data for averages, fits, limits, etc. $ullet$ $ullet$								
$0.57^{+0.13}_{-0.09}\pm0.08$	68	100_{Mo}	2ν	$0^+ \rightarrow 0^+_1$	NEMO-3	¹ ARNOLD	07	
> 89	90	$100_{\mbox{Mo}}$	0ν	$0^+ \rightarrow 0_1^+$	NEMO-3	² ARNOLD	07	
> 1.1	90	$100\mathrm{Mo}$	2ν	$0^+ \rightarrow 2^+$	NEMO-3	³ ARNOLD	07	
> 160	90	100 Mo	0ν	$0^+ \rightarrow 2^+$	NEMO-3	⁴ ARNOLD	07	
22300 + 4400 - 3100	68	76 _{Ge}	0 u		Enriched HPGe	⁵ KLAPDOR-K	. 06A	
> 1800	90	¹³⁰ Te	0ν		Cryog. det.	⁶ ARNABOLDI	05	
> 460	90	¹⁰⁰ Mo	0 u		NEMO-3	⁷ ARNOLD	05A	
> 100	90	⁸² Se	0 u		NEMO-3	⁸ ARNOLD	05A	
$(7.11 \pm 0.02 \pm 0.54)$	4)E-3	3 ¹⁰⁰ Mo	2ν		NEMO-3	⁹ ARNOLD	05A	
$(9.6\pm0.3\pm1.0)$ E	-2	⁸² Se	2ν		NEMO-3	¹⁰ ARNOLD	05A	
> 550	90	¹³⁰ Te	0ν		Cryog. det.	¹¹ ARNABOLDI	04	
> 310	90	¹⁰⁰ Mo	0ν		NEMO-3	¹² ARNOLD	04	
> 140	90	⁸² Se	0ν		NEMO-3	¹³ ARNOLD	04	
$(7.68 \pm 0.02 \pm 0.54)$	4)E-3	3 ¹⁰⁰ Mo	2ν		NEMO-3	¹⁴ ARNOLD	04	
$(10.3 \pm 0.3 \pm 0.7)$	E-2	⁸² Se	2ν		NEMO-3	¹⁵ ARNOLD	04	
$0.14^{+0.04}_{-0.02} \pm 0.03$	68	¹⁵⁰ Nd	$0\nu+2\nu$	$0^+ \rightarrow 0^+_1$	γ in Ge det.	¹⁶ BARABASH	04	
$11900 \stackrel{+29900}{-} 5000$	99.7	76 _{Ge}	0ν	-	Enriched HPGe	¹⁷ KLAPDOR-K	. 04A	
> 14	90	⁴⁸ Ca	0ν		CaF ₂ scint.	¹⁸ OGAWA	04	
> 210	90	¹³⁰ Te	0ν		Cryog. det.	¹⁹ ARNABOLDI	03	
> 31	90	¹³⁰ Te	0 u	$0^+ \rightarrow 2^+$	Cryog. det.	²⁰ ARNABOLDI	03	
$0.61 \pm 0.14 \substack{+0.29 \\ -0.35}$	90	¹³⁰ Te	2ν		Cryog. det.	²¹ ARNABOLDI	03	
> 110	90	¹²⁸ Te	0 u		Cryog. det.	²² ARNABOLDI	03	
(0.029 + 0.004) - 0.003)		^{116}Cd	2ν		¹¹⁶ CdWO ₄ scint	. ²³ DANEVICH	03	
> 170	90	116 Cd	0 u		¹¹⁶ CdWO₄ scint	²⁴ DANEVICH	03	
> 29	90	116 Cd	0 u	$0^+ \rightarrow 2^+$	¹¹⁶ CdWO ₄ scint	. ²⁵ DANEVICH	03	
> 14	90	^{116}Cd	0 u	$0^+ \rightarrow 0_1^+$	¹¹⁶ CdWO ₄ scint	. ²⁶ DANEVICH	03	
> 6	90	^{116}Cd	0ν	$0^+ \rightarrow 0^+_2$	¹¹⁶ CdWO ₄ scint	. ²⁷ DANEVICH	03	
$1.74 \pm 0.01 \substack{+0.18 \\ 0.16}$		76 _{Ge}	2ν	2	Enriched HPGe	²⁸ DOERR	03	
>15700	90	76 _{Ge}	0ν		Enriched HPGe	²⁹ AALSETH	0 2B	
> 58	90	¹³⁴ Xe	0ν		Liquid Xe Scint.	³⁰ BERNABEI	0 2D	
> 1200	90	136 _{Xe}	0 u		Liquid Xe Scint.	³¹ BERNABEI	02D	
$15000 \ +168000 \ -7500$)	76 _{Ge}	0ν		Enriched HPGe	³² KLAPDOR-K	. 02 D	

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$(7.2 \pm 0.9 \pm 1.8)$ E-	-3	100 _{Mo}	2ν		Lia. Ar ioniz.	³³ ASHITKOV	01
> 4.9	90	100 _{Mo}	0ν		Lia. Ar ioniz.	³⁴ ASHITKOV	01
> 1.3	90	160 Gd	0ν		GdっSiO⊑:Ce	³⁵ DANEVICH	01
> 1.3	90	160 Gd	0ν	$0^+ \rightarrow 2^+$	GdoSiO⊑:Ce	³⁶ DANEVICH	01
$0.59^{+0.17}_{-0.11} \pm 0.06$		100 _{Mo}	$0\nu+2\nu$	$0^+ \rightarrow 0_1^+$	Ge coinc.	37 DEBRAECKEL	.01
> 55	90	¹⁰⁰ Mo	0ν , $\langle m_{ij}$		ELEGANT V	³⁸ EJIRI	01
> 42	90	100_{Mo}	$0\nu,\langle\lambda\rangle^{\nu}$	/	ELEGANT V	³⁸ EJIRI	01
> 49	90	100_{Mo}	$0\nu,\langle\eta\rangle$		ELEGANT V	³⁸ EJIRI	01
>19000	90	76 _{Ge}	0ν		Enriched HPGe	³⁹ KLAPDOR-K	01
$1.55 \pm 0.001 + 0.19$	90	76 _{Ge}	2ν		Enriched HPGe	⁴⁰ KLAPDOR-K	01
(9.4 ± 3.2) E-3	90	⁹⁶ Zr	$0\nu+2\nu$		Geochem	⁴¹ WIESER	01
0.042 + 0.033		⁴⁸ Ca	2ν		Ge spectrometer	⁴² BRUDANIN	00
$0.021 + 0.008 \pm 0.0000 \pm 0.00000000000000000000$	02	⁹⁶ Zr	2ν		NEMO-2	⁴³ ARNOLD	99
> 1.0	90	⁹⁶ Zr	0ν		NEMO-2	⁴³ ARNOLD	99
$(8.3 \pm 1.0 \pm 0.7)$ E	-2	⁸² Se	2ν		NEMO-2	⁴⁴ ARNOLD	98
> 9.5	90	82 _{Se}	0 u		NEMO-2	⁴⁵ ARNOLD	98
> 2.8	90	⁸² Se	0 u	$0^+ \rightarrow 2^+$	NEMO-2	⁴⁶ ARNOLD	98
$(7.6^{+2.2}_{-1.4})$ E-3		100_{Mo}	2ν		Si(Li)	⁴⁷ ALSTON	97
$(6.82 + 0.38 \pm 0.68)$)E-3	100_{Mo}	2ν		ТРС	⁴⁸ DESILVA	97
(6.75 + 0.37 = 0.68))E-3	$^{150}\mathrm{Nd}$	2ν		ТРС	⁴⁹ DESILVA	97
> 1.2	90	¹⁵⁰ Nd	0ν		ТРС	⁵⁰ DESILVA	97
$(3.75 \pm 0.35 \pm 0.2$	1)E-2	2^{116} Cd	2ν	$0^+ ightarrow 0^+$	NEMO 2	⁵¹ ARNOLD	96
$0.043^{+0.024}_{-0.011} \pm 0.00$)14	⁴⁸ Ca	2ν		ТРС	⁵² BALYSH	96
0.79 ± 0.10		¹³⁰ Te	$0\nu+2\nu$		Geochem	⁵³ такаока	96
$0.61 \substack{+0.18 \\ -0.11}$		$100\mathrm{Mo}$	$0\nu+2\nu$	$0^+ \rightarrow 0^+_1$	γ in HPGe	⁵⁴ BARABASH	95
$(9.5 \pm 0.4 \pm 0.9)$ E-	-3	100_{Mo}	2ν	1	NEMO 2	DASSIE	95
> 0.6	90	100 Mo	0 u	$0^+ \rightarrow 0^+_1$	NEMO 2	DASSIE	95
$0.026 \substack{+0.009 \\ -0.005}$		$^{116}\mathrm{Cd}$	2ν	$0^+ \rightarrow 0^+$	ELEGANT IV	EJIRI	95
$0.017 \substack{+0.010 \\ -0.005} \pm 0.0$	035	$^{150}\mathrm{Nd}$	2ν	$0^+ \rightarrow 0^+$	ТРС	ARTEMEV	93
0.039 ± 0.009		⁹⁶ Zr	$0\nu+2\nu$		Geochem	KAWASHIMA	93
2.7 ± 0.1		¹³⁰ Te	$0\nu+2\nu$		Geochem	BERNATOW	92
7200 ± 400		¹²⁸ Te	$0\nu+2\nu$		Geochem	⁵⁵ BERNATOW	92
> 27	68	⁸² Se	0ν	$0^+ ightarrow 0^+$	ТРС	ELLIOTT	92
$0.108 \substack{+0.026 \\ -0.006}$		⁸² Se	2ν	$0^+ \rightarrow 0^+$	ТРС	ELLIOTT	92
2.0 ± 0.6		238 _U	$0\nu+2\nu$		Radiochem	⁵⁶ TURKEVICH	91
> 9.5	76	⁴⁸ Ca	0 u		CaF ₂ scint.	YOU	91
$0.12 \pm 0.01 \pm 0.04$	68	⁸² Se	0 u+2 u		Geochem.	⁵⁷ LIN	88
$0.75 \pm 0.03 \pm 0.23$	68	¹³⁰ Te	0 u+2 u		Geochem.	⁵⁸ LIN	88
1800 ± 700	68	¹²⁸ Te	$0\nu+2\nu$		Geochem.	⁵⁹ LIN	88 B
2.60 ± 0.28		¹³⁰ Te	$0\nu+2\nu$		Geochem	⁶⁰ KIRSTEN	83

¹ First exclusive measurement of 2ν -decay to the first excited 0_1^+ -state of daughter nucleus. ARNOLD 07 use the NEMO-3 tracking calorimeter to detect all particles emitted in decay. Result agrees with the inclusive ($0\nu + 2\nu$) measurement of DEBRAECKELEER 01. ² Limit on 0ν -decay to the first excited 0_1^+ -state of daughter nucleus using NEMO-3 tracking calorimeter. Supersedes DASSIE 95.

- ³Limit on 2ν -decay to the first excited 2⁺-state of daughter nucleus using NEMO-3 tracking calorimeter.
- ⁴Limit on 0ν -decay to the first excited 2⁺-state of daughter nucleus using NEMO-3 tracking calorimeter.
- ⁵ KLAPDOR-KLEINGROTHAUS 06A present re-analysis of data originally published in KLAPDOR-KLEINGROTHAUS 04A. Modified pulse shape analysis leads the authors to claim improved 6σ statistical evidence for observation of 0ν -decay, compared to 4.2σ in KLAPDOR-KLEINGROTHAUS 04A. Analysis of the systematic uncertainty is not presented. Supersedes KLAPDOR-KLEINGROTHAUS 04A.
- ⁶ Supersedes ARNABOLDI 04. Bolometric TeO₂ detector array CUORICINO is used for high resolution search for $0\nu\beta\beta$ decay. The half-life limit is derived from 3.09 kg yr 130 Te exposure.
- ⁷NEMO-3 tracking calorimeter containing 6.9 kg of enriched ¹⁰⁰Mo is used in ARNOLD 05A. A limit for $0\nu\beta\beta$ half-life of ¹⁰⁰Mo is reported. Supersedes ARNOLD 04.
- ⁸ NEMO-3 tracking calorimeter is used in ARNOLD 05A to place limit on $0\nu\beta\beta$ half-life of ⁸²Se. Detector contains 0.93 kg of enriched ⁸²Se. Supersedes ARNOLD 04.
- ⁹ ARNOLD 05A use the NEMO-3 tracking calorimeter to determine the $2\nu\beta\beta$ half-life of 100 Mo with high statistics and low background (389 days of data taking). Supersedes ARNOLD 04.
- ¹⁰ ARNOLD 05A use the NEMO-3 tracking detector to determine the $2\nu\beta\beta$ half-life of ⁸²Se with high statistics and low background (389 days of data taking). Supersedes , ARNOLD 04.
- ¹¹ Supersedes ARNABOLDI 03. Bolometric TeO₂ detector array Cuoricino used for high resolution search for $0\nu\beta\beta$ decay.
- ¹² ARNOLD 04 use the NEMO-3 tracking detector to determine the limit for $0\nu\beta\beta$ halflife of ¹⁰⁰Mo. This represents an improvement, by a factor of \sim 6, when compared with EJIRI 01.
- ¹³ ARNOLD 04 use the NEMO-3 tracking detector to determine the limit for $0\nu\beta\beta$ halflife of ⁸²Se. This represents an improvement, by a factor of ~ 10, when compared with ELLIOTT 92. It supersedes the limit of ARNOLD 98 for this decay using NEMO-2.
- ¹⁴ ARNOLD 04 use the NEMO-3 tracking detector to determine the $2\nu\beta\beta$ halflife of ¹⁰⁰ Mo with high statistics and low background. The halflife is determined assuming the Single State Dominance. It is in agreement with, and more accurate than, previous determinations. Supersedes DASSIE 95 determination of this quantity with NEMO-2.
- ¹⁵ ARNOLD 04 use the NEMO-3 tracking detector to determine the $2\nu\beta\beta$ halflife of ⁸²Se. The halflife is in agreement with ARNOLD 98 with NEMO-2 which it supersedes. ¹⁶ BARABASH 04 perform an inclusive measurement of the $\beta\beta$ decay of ¹⁵⁰Nd into the
- ¹⁶ BARABASH 04 perform an inclusive measurement of the $\beta\beta$ decay of ¹⁵⁰Nd into the first excited (0^+_1) state of ¹⁵⁰Sm. Gamma radiation emitted in decay of the excited state is detected.
- ¹⁷ Supersedes KLAPDOR-KLEINGROTHAUS 02D. Authors present new analysis of event excess seen in Heidelberg-Moscow experiment at $\beta\beta$ -decay energy. Enhanced statistics leads to a 4.2 σ evidence for observation of $0\nu\beta\beta$ -decay and a finite Majorana neutrino mass. Stated error is purely statistical. No systematic errors are mentioned in the paper. More details can be found in KLAPDOR-KLEINGROTHAUS 04C.
- $^{18}\,\text{CaF}_2$ scintillation calorimeter ELEGANT VI used to set limit on $0\nu\,\beta\,\beta$ -decay rate of $^{48}\,\text{Ca.}$ The stated half-life limit benefits from a downward fluctuation on the number of background events. The experimental sensitivity is 5.9×10^{21} yr. at 90 % CL. Replaces YOU 91 as the most stringest experiment using $^{48}\,\text{Ca.}$
- ¹⁹ Supersedes ALESSANDRELLO 00. Array of TeO₂ crystals in high resolution cryogenic calorimeter. Some enriched in ¹³⁰Te. Ground state to ground state decay.
- ²⁰ Decay into first excited state of daughter nucleus.
- ²¹ Two neutrino decay into ground state. Relatively large error mainly due to uncertainties in background determination. Reported value is shorter than the geochemical measurements of KIRSTEN 83 and BERNATOWICZ 92 but in agreement with LIN 88 and TAKAOKA 96.

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- 22 Supersedes ALESSANDRELLO 00. Array of TeO $_2$ crystals in high resolution cryogenic calorimeter. Some enriched in ¹²⁸Te. Ground state to ground state decay.
- 23 Calorimetric measurement of 2ν ground state decay of 116 Cd using enriched CdWO₄ scintillators. Agrees with EJIRI 95 and ARNOLD 96. Supersedes DANEVICH 00.
- ²⁴Limit on 0ν decay of ¹¹⁶Cd using enriched CdWO₄ scintillators. Supersedes DANEVICH 00.
- 25 Limit on 0 ν decay of 116 Cd into first excited 2⁺ state of daughter nucleus using enriched CdWO₄ scintillators. Supersedes DANEVICH 00.
- 26 Limit on 0ν decay of 116 Cd into first excited 0^+ state of daughter nucleus using enriched CdWO_4 scintillators. Supersedes DANEVICH 00.
- 27 Limit on 0ν decay of 116 Cd into second excited 0⁺ state of daughter nucleus using enriched CdWO₄ scintillators. Supersedes DANEVICH 00.
- 28 Results of the Heidelberg-Moscow experiment (KLAPDOR-KLEINGROTHAUS 01 and GUENTHER 97) are reanalyzed using a new simulation of the complete background spectrum. The $\beta\beta 2\nu$ -decay rate is deduced from a 41.57 kg-y exposure. The result is in agreement and supersedes the above referenced halflives with similar statistical and systematic errors.
- 29 AALSETH 02B limit is based on 117 mol \cdot yr of data using enriched Ge detectors. Background reduction by means of pulse shape analysis is applied to part of the data set. Reported limit is slightly less restrictive than that in KLAPDOR-KLEINGROTHAUS 01 However, it excludes part of the allowed half-life range reported in KLAPDOR-KLEINGROTHAUS 01B for the same nuclide. The analysis has been criticized in KLAPDOR-KLEINGROTHAUS 04B. The criticism was addressed and disputed in AALSETH 04.
- 30 BERNABEI 02D report a limit for the 0 $\nu,$ 0 $^+~\rightarrow~0^+$ decay of 134 Xe, present in the source at 17%, by considering the maximum number of events for this mode compatible with the fitted smooth background.
- 31 BERNABEI 02D report a limit for the 0u, 0 $^+$ ightarrow 0 $^+$ decay of 136 Xe, by considering the maximum number of events for this mode compatible with the fitted smooth background. The quoted sensitivity is 450×10^{21} yr. The Feldman and Cousins method is used to obtain the quoted limit.
- ³² KLAPDOR-KLEINGROTHAUS 02D is expanded an version of KLAPDOR-KLEINGROTHAUS 01B. The authors re-evaluate the data collected by the Heidelberg-Moscow experiment (KLAPDOR-KLEINGROTHAUS 01) and present a more detailed description of their analysis of an excess of counts at the energy expected for neutrinoless double-beta decay. They interpret this excess, which has a significance of 2.2 to 3.1 σ depending on the data analysis, as evidence for the observation of Lepton Number violation and violation of Baryon minus Lepton Number. The analysis has been criticized by AALSETH 02 and others. The criticisms have been addressed in KLAPDOR-KLEINGROTHAUS 02. See also KLAPDOR-KLEINGROTHAUS 02B. GRO-MOV 06 analysis of the background supports the assignment of the weak γ transitions near 2040 keV to the decay of 214 Bi as claimed in KLAPDOR-KLEINGROTHAUS 04A and KLAPDOR-KLEINGROTHAUS 04C, and in earlier works. 33 ASHITKOV 01 result for 2ν of 100 Mo is in agreement with other determinations of that
- halflife.
- 34 ASHITKOV 01 result for 0 ν of 100 Mo is less stringent than EJIRI 01.
- 35 DANEVICH 01 place limit on 0 ν decay of 160 Gd using Gd₂SiO₅:Ce crystal scintillators. The limit is more stringent than KOBAYASHI 95.
- 36 DANEVICH 01 place limits on 0 ν decay of $^{160}{\rm Gd}$ into excited 2 $^+$ state of daughter nucleus using Gd₂SiO₅:Ce crystal scintillators.
- 37 DEBRAECKELEER 01 performed an inclusive measurement of the etaeta decay into the second excited state of the daughter nucleus. A novel coincidence technique counting the de-excitation photons is employed. The result agrees with BARABASH 95.
- 38 EJIRI 01 uses tracking calorimeter and isotopically enriched passive source. Efficiencies were calculated assuming $\langle m_{\nu} \rangle$, $\langle \lambda \rangle$, or $\langle \eta \rangle$ driven decay. This is a continuation of EJIRI 96 which it supersedes.

- ³⁹ KLAPDOR-KLEINGROTHAUS 01 is a continuation of the work published in BAUDIS 99. Isotopically enriched Ge detectors are used in calorimetric measurement. The most stringent bound is derived from the data set in which pulse-shape analysis has been used to reduce background. Exposure time is 35.5 kg y. Supersedes BAUDIS 99 as most stringent result.
- 40 KLAPDOR-KLEINGROTHAUS 01 is a measurement of the etaeta 2
 u-decay rate with higher statistics than GUENTHER 97. The reported value has a larger systematic error than their previous result.
- ⁴¹WIESER 01 reports an inclusive geochemical measurement of 96 Zr $_{\beta\beta}$ half life. Their result agrees within 2σ with ARNOLD 99 but only marginally, within 3σ , with KAWASHIMA 93.
- 42 BRUDANIN 00 determine the 2 ν halflife of 48 Ca. Their value is less accurate than BALYSH 96.
- 43 ARNOLD 99 measure directly the 2
 u decay of Zr for the first time, using the NEMO-2 tracking detector and an isotopically enriched source. The lifetime is more accurate than the geochemical result of KAWASHIMA 93.
- 44 ARNOLD 98 measure the 2ν decay of 82 Se by comparing the spectra in an enriched and natural selenium source using the NEMO-2 tracking detector. The measured half-life is in agreement, perhaps slightly shorter, than ELLIOTT 92.
- 45 ARNOLD 98 determine the limit for 0u decay to the ground state of 82 Se using the NEMO-2 tracking detector. The half-life limit is in agreement, but less stringent, than ELLIOTT 92.
- ⁴⁶ ARNOLD 98 determine the limit for 0ν decay to the excited 2⁺ state of ⁸²Se using the NEMO-2 tracking detector.
- ⁴⁷ALSTON-GARNJOST 97 report evidence for 2ν decay of ¹⁰⁰Mo. This decay has been also observed by EJIRI 91, DASSIE 95, and DESILVA 97.
- 48 DESILVA 97 result for 2ν decay of 100 Mo is in agreement with ALSTON-GARNJOST 97 and DASSIE 95. This measurement has the smallest errors. DESILVA 97 result for 2ν decay of 150 Nd is in marginal agreement with ARTEMEV 93.
- It has smaller errors.
- 50 DESILVA 97 do not explain whether their efficiency for 0u decay of 150 Nd was calculated under the assumption of a $\langle m_{\mu} \rangle$, $\langle \lambda \rangle$, or $\langle \eta \rangle$ driven decay.
- 51 ARNOLD 96 measure the 2 ν decay of 116 Cd. This result is in agreement with EJIRI 95, but has smaller errors. Supersedes ARNOLD 95.
- 52 BALYSH 96 measure the 2 ν decay of 48 Ca, using a passive source of enriched 48 Ca in a TPC.
- 53 TAKAOKA 96 measure the geochemical half-life of 130 Te. Their value is in disagreement with the quoted values of BERNATOWICZ 92 and KIRSTEN 83; but agrees with several other unquoted determinations, e.g., MANUEL 91.
- 54 BARABASH 95 cannot distinguish 0 ν and 2 ν , but it is inferred indirectly that the 0 ν mode accounts for less than 0.026% of their event sample. They also note that their result disagrees with the previous experiment by the NEMO group (BLUM 92).
- ⁵⁵ BERNATOWICZ 92 finds 128 Te/130 Te activity ratio from slope of 128 Xe/132 Xe vs 130 Xe/ 132 Xe ratios during extraction, and normalizes to lead-dated ages for the 130 Te lifetime. The authors state that their results imply that "(a) the double beta decay of 128 Te has been firmly established and its half-life has been determined ... without any ambiguity due to trapped Xe interferences... (b) Theoretical calculations ... underestimate the [long half-lives of ¹²⁸Te ¹³⁰Te] by 1 or 2 orders of magnitude, pointing to a real suppression in the 2ν decay rate of these isotopes. (c) Despite [this], most $\beta\beta$ -models predict a ratio of 2ν decay widths ... in fair agreement with observation." Further details of the experiment are given in BERNATOWICZ 93. Our listed half-life has been revised downward from the published value by the authors, on the basis of reevaluated cosmic-ray ¹²⁸Xe production corrections.
- 56 TURKEVICH 91 observes activity in old U sample. The authors compare their results with theoretical calculations. They state "Using the phase-space factors of Boehm and Vogel (BOEHM 87) leads to matrix element values for the ²³⁸U transition in the same

range as deduced for 130 Te and 76 Ge. On the other hand, the latest theoretical estimates (STAUDT 90) give an upper limit that is 10 times lower. This large discrepancy implies either a defect in the calculations or the presence of a faster path than the standard two-neutrino mode in this case." See BOEHM 87 and STAUDT 90.

- 57 Result agrees with direct determination of ELLIOTT 92.
- ⁵⁸ Inclusive half life inferred from mass spectroscopic determination of abundance of $\beta\beta$ decay product ¹³⁰Te in mineral kitkaite (NiTeSe). Systematic uncertainty reflects variations in U-Xe gas-retention-age derived from different uranite samples. Agrees with geochemical determination of TAKAOKA 96 and direct measurement of ARNABOLDI 03. Inconsistent with results of KIRSTEN 83 and BERNATOWICZ 92.
- Inconsistent with results of KIRSTEN 83 and BERNATOWICZ 92. ⁵⁹ Ratio of inclusive double beta half lives of ¹²⁸Te and ¹³⁰Te determined from minerals melonite (NiTe₂) and altaite (PbTe) by means of mass spectroscopic measurement of abundance of $\beta\beta$ -decay products. As gas-retention-age could not be determined the authors use half life of ¹³⁰Te (LIN 88) to infer the half life of ¹²⁸Te. No estimate of the systematic uncertainty of this method is given. The directly determined half life ratio agrees with BERNATOWICZ 92. However, the inferred ¹²⁸Te half life disagrees with KIRSTEN 83 and BERNATOWICZ 92.
- 60 KIRSTEN 83 reports " 2σ " error. References are given to earlier determinations of the 130 Te lifetime.

$\langle m_{\nu} \rangle$, The Effective Weighted Sum of Majorana Neutrino Masses Contributing to Neutrinoless Double- β Decay

 $\langle m_{\nu} \rangle = |\Sigma \ U_{1\,j}^2 m_{\nu_j}|$, where the sum goes from 1 to *n* and where *n* = number of neutrino generations, and ν_j is a Majorana neutrino. Note that $U_{e\,j}^2$, not $|U_{e\,j}|^2$, occurs in the sum. The possibility of cancellations has been stressed. In the following Listings, only best or comparable limits or lifetimes for each isotope are reported.

VALUE (eV)	<u>CL%</u>	<u>ISOTOPE</u>	TRANSITION	METHOD	DOCUMENT ID		
ullet $ullet$ $ullet$ We do not use the following data for averages, fits, limits, etc. $ullet$ $ullet$							
< 9.3–60	90	100 _{Mo}	$0^+ \rightarrow 0^+_1$	NEMO-3	⁶¹ ARNOLD	07	
< 6500	90	¹⁰⁰ Mo	$0^+ \rightarrow 2^+$	NEMO-3	⁶² ARNOLD	07	
0.32 ± 0.03	68	⁷⁶ Ge	0 u	Enriched HPGe	⁶³ KLAPDOR-K	. 06A	
< 0.2–1.1	90	¹³⁰ Te		Cryog. det.	⁶⁴ ARNABOLDI	05	
< 0.7–2.8	90	¹⁰⁰ Mo	0 u	NEMO-3	⁶⁵ ARNOLD	05A	
< 1.7–4.9	90	⁸² Se	0 u	NEMO-3	⁶⁶ ARNOLD	05A	
< 0.37–1.9	90	¹³⁰ Te		Cryog. det.	⁶⁷ ARNABOLDI	04	
< 0.8–1.2	90	¹⁰⁰ Mo	0 u	NEMO-3	⁶⁸ ARNOLD	04	
< 1.5 - 3.1	90	⁸² Se	0 u	NEMO-3	⁶⁸ ARNOLD	04	
0.1-0.9	99.	7 ⁷⁶ Ge		Enriched HP Ge	⁶⁹ KLAPDOR-K	. 04A	
< 7.2–44.7	90	⁴⁸ Ca		CaF ₂ scint.	⁷⁰ OGAWA	04	
< 1.1–2.6	90	¹³⁰ Te		Cryog. det.	⁷¹ ARNABOLDI	03	
< 1.5 - 1.7	90	^{116}Cd	0 u	¹¹⁶ CdWO ₄ scint.	⁷² DANEVICH	03	
< 0.33–1.35	90			Enriched HPGe	⁷³ AALSETH	0 2B	
<2.9	90	¹³⁶ Xe	0 u	Liquid Xe Scint.	⁷⁴ BERNABEI	0 2D	
$0.39\substack{+0.17 \\ -0.28}$		76 _{Ge}	0ν	Enriched HPGe	⁷⁵ KLAPDOR-K	. 02 D	
< 2.1–4.8	90	¹⁰⁰ Mo	0 u	ELEGANT V	⁷⁶ EJIRI	01	
< 0.35	90	76 _{Ge}		Enriched HPGe	⁷⁷ KLAPDOR-K	. 01	
<23	90	⁹⁶ Zr		NEMO-2	⁷⁸ ARNOLD	99	
< 1.1 - 1.5		¹²⁸ Te		Geochem	⁷⁹ BERNATOW	. 92	
<5	68	⁸² Se		ТРС	⁸⁰ ELLIOTT	92	
<8.3	76	⁴⁸ Ca	0ν	CaF ₂ scint.	YOU	91	

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- ⁶¹ ARNOLD 07 use NEMO-3 half life limit for 0ν -decay of ¹⁰⁰Mo to the first excited 0_1^+ state of daughter nucleus to obtain neutrino mass limit. The spread reflects the choice of two different nuclear matrix elements. This limit is not competitive when compared to the decay to the ground state.
- ⁶² ARNOLD 07 use NEMO-3 half life limit for 0ν -decay of ¹⁰⁰Mo to the first excited 2⁺state of daughter nucleus to obtain neutrino mass limit. This limit is not competitive when compared to the decay to the ground state.
- ⁶³ Re-analysis of data originally published in KLAPDOR-KLEINGROTHAUS 04A. Modified pulse shape analysis leads the authors to claim 6σ statistical evidence for observation of 0ν -decay. Authors use matrix element of STAUDT 90. Uncertainty of nuclear matrix element is not reflected in stated error. Supersedes KLAPDOR-KLEINGROTHAUS 04A.
- ⁶⁴ Supersedes ARNABOLDI 04. Reported range of limits due to use of different nuclear matrix element calculations.
- ⁶⁵ Mass limits reported in ARNOLD 05A are derived from ¹⁰⁰Mo data, obtained by the NEMO-3 collaboration. The range reflects the spread of matrix element calculations considered in this work. Supersedes ARNOLD 04.
- ⁶⁶ Neutrino mass limits based on ⁸²Se data utilizing the NEMO-3 detector. The range reported in ARNOLD 05A reflects the spread of matrix element calculations considered in this work. Supersedes ARNOLD 04.
- ⁶⁷ Supersedes ARNABOLDI 03. Reported range of limits due to use of different nuclear matrix element calculations.
- ⁶⁸ ARNOLD 04 limit is based on the nuclear matrix elements of SIMKOVIC 99, STOICA 01 and CIVITARESE 03.
- ⁶⁹ Supersedes KLAPDOR-KLEINGROTHAUS 02D. Event excess at $\beta\beta$ -decay energy is used to derive Majorana neutrino mass using the nuclear matrix elements of STAUDT 90. The mass range shown is based on the authors evaluation of the uncertainties of the STAUDT 90 matrix element calculation. If this uncertainty is neglected, and only statistical errors are considered, the range in $\langle m \rangle$ becomes (0.2–0.6) eV at the 3 σ level.
- ⁷⁰ Calorimetric CaF₂ scintillator. Range of limits reflects authors' estimate of the uncertainty of the nuclear matrix elements. Replaces YOU 91 as the most stringest limit based on ⁴⁸Ca.
- ⁷¹ Supersedes ALESSANDRELLO 00. Cryogenic calorimeter search. Reported a range reflecting uncertainty in nuclear matrix element calculations.
- 72 Limit for $\langle m_{\nu}\rangle$ is based on the nuclear matrix elements of STAUDT 90 and ARNOLD 96. Supersedes DANEVICH 00.
- 73 AALSETH 02B reported range of limits on $\langle m_{\nu} \rangle$ reflects the spread of theoretical nuclear matrix elements. Excludes part of allowed mass range reported in KLAPDOR-__KLEINGROTHAUS 01B.
- ⁷⁴ BERNABEI 02D limit is based on the matrix elements of SIMKOVIC 02. The range of neutrino masses based on a variety of matrix elements is 1.1–2.9 eV.
- ⁷⁵ KLAPDOR-KLEINGROTHAUS 02D is a detailed description of the analysis of the data collected by the Heidelberg-Moscow experiment, previously presented in KLAPDOR-KLEINGROTHAUS 01B. Matrix elements in STAUDT 90 have been used. See the footnote in the preceding table for further details. See also KLAPDOR-KLEINGROTHAUS 02B.
- ⁷⁶ The range of the reported $\langle m_{\nu} \rangle$ values reflects the spread of the nuclear matrix elements. On axis value assuming $\langle \lambda \rangle = \langle \eta \rangle = 0$.
- ⁷⁷ KLAPDOR-KLEINGROTHAUS 01 uses the calculation by STAUDT 90. Using several other models in the literature could worsen the limit up to 1.2 eV. This is the most stringent experimental bound on m_{ν} . It supersedes BAUDIS 99B.
- $\frac{78}{22}$ ARNOLD 99 limit based on the nuclear matrix elements of STAUDT 90.
- ⁷⁹ BERNATOWICZ 92 finds these majorana neutrino mass limits assuming that the measured geochemical decay width is a limit on the 0ν decay width. The range is the range found using matrix elements from HAXTON 84, TOMODA 87, and SUHONEN 91. Further details of the experiment are given in BERNATOWICZ 93.
- ⁸⁰ ELLIOTT 92 uses the matrix elements of HAXTON 84.

Limits on Lepton-Number Violating (V+A) Current Admixture

For reasons given in the discussion at the beginning of this section, we list only results from 1989 and later. $\langle \lambda \rangle = \lambda \sum U_{ej} V_{ej}$ and $\langle \eta \rangle = \eta \sum U_{ej} V_{ej}$, where the sum is over the number of neutrino generations. This sum vanishes for massless or unmixed neutrinos. In the following Listings, only best or comparable limits or lifetimes for each isotope are reported.

$\left<\lambda\right>$ (10 ⁻⁶)	CL%	$\left<\eta\right>$ (10 ⁻⁸)	CL%	ISOTOPE	METHOD	DOCUMENT ID		
• • We do not use the following data for averages, fits, limits, etc. • • •								
<120	90			100 _{Mo}	$0^+ \rightarrow 2^+$	⁸¹ ARNOLD	07	
0.692 + 0.05	8 68	$0.305 \substack{+0.02 \\ -0.02}$	6 5 68	^{76}Ge	Enriched HPGe	⁸² KLAPDOR-K	. 06A	
< 2.5	90		-	¹⁰⁰ Mo	0ν , NEMO-3	⁸³ ARNOLD	05A	
< 3.8	90			⁸² Se	0ν , NEMO-3	⁸⁴ ARNOLD	05A	
< 1.5 - 2.0	90			¹⁰⁰ Mo	0ν , NEMO-3	⁸⁵ ARNOLD	04	
< 3.2–3.8	90			⁸² Se	0ν , NEMO-3	⁸⁶ ARNOLD	04	
< 1.6 - 2.4	90	< 0.9–5.3	90	¹³⁰ Te	Cryog. det.	⁸⁷ ARNABOLDI	03	
< 2.2	90	<2.5	90	^{116}Cd	¹¹⁶ CdWO ₄ scint.	⁸⁸ DANEVICH	03	
< 3.2–4.7	90	< 2.4–2.7	90	¹⁰⁰ Mo	ELEGANT V	⁸⁹ EJIRI	01	
< 1.1	90	<0.64	90	76 Ge	Enriched HPGe	⁹⁰ GUENTHER	97	
< 4.4	90	<2.3	90	¹³⁶ Xe	ТРС	⁹¹ VUILLEUMIER	93	
		<5.3		¹²⁸ Te	Geochem	⁹² BERNATOW	92	

⁸¹ ARNOLD 07 use NEMO-3 half life limit for 0 ν -decay of ¹⁰⁰Mo to the first excited 2⁺-state of daughter nucleus to limit the right-right handed admixture of weak currents $\langle \lambda \rangle$. This limit is not competitive when compared to the decay to the ground state.

- ⁸² Re-analysis of data originally published in KLAPDOR-KLEINGROTHAUS 04A. Modified pulse shape analysis leads the authors to claim 6σ statistical evidence for observation of 0ν -decay. Authors use matrix element of MUTO 89 to determine $\langle \lambda \rangle$ and $\langle \eta \rangle$. Uncertainty of nuclear matrix element is not reflected in stated errors.
- ⁸³ ARNOLD 05A derive limit for $\langle \lambda \rangle$ based on ¹⁰⁰Mo data collected with NEMO-3 detector. No limit for $\langle \eta \rangle$ is given. Supersedes ARNOLD 04.
- ⁸⁴ ARNOLD 05A derive limit for $\langle \lambda \rangle$ based on ⁸²Se data collected with NEMO-3 detector. No limit for $\langle \eta \rangle$ is given. Supersedes ARNOLD 04.
- ⁸⁵ ARNOLD 04 use the matrix elements of SUHONEN 94 to obtain a limit for $\langle \lambda \rangle$, no limit for $\langle \eta \rangle$ is given. This limit is more stringent than the limit in EJIRI 01 for the same nucleus.
- ⁸⁶ ARNOLD 04 use the matrix elements of TOMODA 91 and SUHONEN 91 to obtain a limit for $\langle \lambda \rangle$, no limit for $\langle \eta \rangle$ is given.
- ⁸⁷ Supersedes ALESSANDRELLO 00. Cryogenic calorimeter search. Reported a range reflecting uncertainty in nuclear matrix element calculations.
- 88 Limits for $\langle\lambda\rangle$ and $\langle\eta\rangle$ are based on nuclear matrix elements of STAUDT 90. Supersedes a DANEVICH 00.
- ⁸⁹ The range of the reported $\langle \lambda \rangle$ and $\langle \eta \rangle$ values reflects the spread of the nuclear matrix elements. On axis value assuming $\langle m_{\mu} \rangle = 0$ and $\langle \lambda \rangle = \langle \eta \rangle = 0$, respectively.
- 90 GUENTHER 97 limits use the matrix elements of STAUDT 90. Supersedes BALYSH 95 and BALYSH 92.
- 91 VUILLEUMIER 93 uses the matrix elements of MUTO 89. Based on a half-life limit 2.6×10^{23} y at 90%CL.
- ⁹² BERNATOWICZ 92 takes the measured geochemical decay width as a limit on the 0ν width, and uses the SUHONEN 91 coefficients to obtain the least restrictive limit on η . Further details of the experiment are given in BERNATOWICZ 93.

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GROMOV	06	PPNL 3 157	K. Gromov <i>et al.</i>	
KLAPDOR-K	06A	MPL A21 1547	H.V. Klapdor-Kleingrothaus,	I.V. Krivosheina
ARNABOLDI	05	PRL 95 142501	C. Arnaboldi <i>et al.</i>	(CUORICINO Collab.)
ARNOLD	05A	PRL 95 182302	R. Arnold <i>et al.</i>	(NEMO-3 Collab.)
	04	PR D/U U/8302	C.E. Aalseth <i>et al.</i>	
	04	PL B384 200	C. Arnaboldi <i>et al.</i>	(NEMO2 Detector Colleb.)
ARNOLD	04	Translated from 7FTFP	80 429	(NEWOS Detector Collab.)
BARABASH	04	JETPL 79 10	A.S. Barabash <i>et al.</i>	
KLAPDOR-K	04A	PL B586 198	H.V. Klapdor-Kleingrothaus	et al.
KLAPDOR-K	04B	PR D70 078301	H.V. Klapdor-Kleingrothaus,	A. Dietz, I.V. Krivosheina
KLAPDOR-K	04C	NIM A522 371	H.V. Klapdor-Kleingrothaus	et al.
OGAWA	04	NP A730 215	I. Ogawa <i>et al.</i>	
ARNABOLDI	03	PL B557 167	C. Arnaboldi <i>et al.</i>	
CIVITARESE	03	NP A/29 867	O. Civitarese, J. Suhonen	
	03	PR C08 035501	F.A. Danevich <i>et al.</i>	
	03	MDI A17 1475	C. Doerr, H.V. Klapdor-Klei	ngrotnaus
	02 02B	PR D65 002007	C.E. Adiseth <i>et al.</i>	(IGEX Collab.)
BERNABEI	02D	PI B546 23	R Bernabei <i>et al</i>	(DAMA Collab.)
KLAPDOR-K	02	hep-ph/0205228	H V Klapdor-Kleingrothaus	
KLAPDOR-K	02B	JINRRC 110 57	H.V. Klapdor-Kleingrothaus.	A. Dietz. I.V. Krivosheina
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SIMKOVIC	02	hep-ph/0204278	F. Simkovic, P. Domin, A. I	Faessler
ASHITKOV	01	JETPL 74 529	V.D. Ashitkov <i>et al.</i>	
		Translated from ZETFP	74 601.	
DANEVICH	01	NP A694 375	F.A. Danevich <i>et al.</i>	
DEBRAECKEL.	01	PRL 86 3510	L. De Braeckeleer <i>et al.</i>	
EJIKI	01	PR C63 065501	H. Ejiri <i>et al.</i>	
KLAPDOR-K	01	EPJ A12 14/	H.V. Klapdor-Kleingrothaus	et al.
	010	NPL A10 2409	S Stoice HV Klander Klai	el al.
WIESER	01	PR C64 024308	M F Wieser I R De Laete	r
ALESSAND	00	PI B486 13	A Alessandrello <i>et al</i>	•
BRUDANIN	00	PL B495 63	V.B. Brudanin <i>et al.</i>	
DANEVICH	00	PR C62 045501	F.A. Danevich et al.	
ARNOLD	99	NP A658 299	R. Arnold <i>et al.</i>	(NEMO Collab.)
BAUDIS	99	PR D59 022001	L. Baudis <i>et al.</i>	(Heidelberg-Moscow Collab.)
BAUDIS	99B	PRL 83 41	L. Baudis <i>et al.</i>	(Heidelberg-Moscow Collab.)
SIMKOVIC	99	PR C60 055502	F. Simkovic <i>et al.</i>	
ARNOLD	98	NP A636 209		
ALSTON	07		R. Arnold <i>et al.</i>	(INEIMO-2 Collab.)
	97 07	PR C55 474	R. Arnold <i>et al.</i> M. Alston-Garnjost <i>et al.</i>	(NEMO-2 Collab.) (LBL, MTHO+)
	97 97 07	PR C55 474 PR C56 2451 PR D55 54	 A. Alston-Garnjost <i>et al.</i> A. de Silva <i>et al.</i> 	(NEMO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg Massey Collab.)
DESILVA GUENTHER	97 97 97 96	PR C55 474 PR C56 2451 PR D55 54 ZPHV C72 220	 Arnold <i>et al.</i> M. Alston-Garnjost <i>et al.</i> A. de Silva <i>et al.</i> M. Gunther <i>et al.</i> B. Arnold <i>et al.</i> 	(NEMIO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.)
DESILVA GUENTHER ARNOLD BALVSH	97 97 97 96 96	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRI 77 5186	 R. Arnold <i>et al.</i> M. Alston-Garnjost <i>et al.</i> A. de Silva <i>et al.</i> M. Gunther <i>et al.</i> R. Arnold <i>et al.</i> A. Balveb <i>et al.</i> 	(NEMIO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCL CIT)
DESILVA GUENTHER ARNOLD BALYSH F IIRI	97 97 97 96 96 96	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85	 R. Arnold <i>et al.</i> M. Alston-Garnjost <i>et al.</i> A. de Silva <i>et al.</i> M. Gunther <i>et al.</i> R. Arnold <i>et al.</i> A. Balysh <i>et al.</i> H. Fiiri <i>et al.</i> 	(NEMIO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK)
DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA	97 97 96 96 96 96 96	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557	 K. Arnold <i>et al.</i> M. Alston-Garnjost <i>et al.</i> A. de Silva <i>et al.</i> M. Gunther <i>et al.</i> R. Arnold <i>et al.</i> A. Balysh <i>et al.</i> H. Ejiri <i>et al.</i> N. Takaoka, Y. Motomura. 	(NEMIO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY)
DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA ARNOLD	97 97 96 96 96 96 96 95	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557 JETPL 61 170	 R. Arnold <i>et al.</i> M. Alston-Garnjost <i>et al.</i> A. de Silva <i>et al.</i> M. Gunther <i>et al.</i> R. Arnold <i>et al.</i> A. Balysh <i>et al.</i> H. Ejiri <i>et al.</i> N. Takaoka, Y. Motomura, R.G. Arnold <i>et al.</i> 	(NEMIO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY) (NEMO Collab.)
DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA ARNOLD	97 97 96 96 96 96 96 95	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557 JETPL 61 170 Translated from ZETFP	 R. Arnold <i>et al.</i> M. Alston-Garnjost <i>et al.</i> A. de Silva <i>et al.</i> M. Gunther <i>et al.</i> R. Arnold <i>et al.</i> A. Balysh <i>et al.</i> H. Ejiri <i>et al.</i> N. Takaoka, Y. Motomura, R.G. Arnold <i>et al.</i> 61 168. 	(NEMO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY) (NEMO Collab.)
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DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA ARNOLD BALYSH BARABASH	97 97 96 96 96 96 95 95 95	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557 JETPL 61 170 Translated from ZETFP PL B356 450 PL B345 408	R. Arnold <i>et al.</i> M. Alston-Garnjost <i>et al.</i> A. de Silva <i>et al.</i> M. Gunther <i>et al.</i> R. Arnold <i>et al.</i> A. Balysh <i>et al.</i> H. Ejiri <i>et al.</i> N. Takaoka, Y. Motomura, R.G. Arnold <i>et al.</i> 61 168. A. Balysh <i>et al.</i> A.S. Barabash <i>et al.</i>	(NEMO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY) (NEMO Collab.) (Heidelberg-Moscow Collab.) (ITEP, SCUC, PNL+)
DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA ARNOLD BALYSH BARABASH DASSIE	97 97 97 96 96 96 96 96 95 95 95 95	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557 JETPL 61 170 Translated from ZETFP PL B356 450 PL B345 408 PR D51 2090	 R. Arnold et al. M. Alston-Garnjost et al. A. de Silva et al. M. Gunther et al. R. Arnold et al. R. Arnold et al. H. Ejiri et al. H. Ejiri et al. N. Takaoka, Y. Motomura, R.G. Arnold et al. 61 168. A. Balysh et al. A. Balysh et al. A.S. Barabash et al. D. Dassie et al. 	(NEMO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY) (NEMO Collab.) (Heidelberg-Moscow Collab.) (ITEP, SCUC, PNL+) (NEMO Collab.)
DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA ARNOLD BALYSH BARABASH DASSIE EJIRI	97 97 97 96 96 96 96 96 95 95 95 95 95	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557 JETPL 61 170 Translated from ZETFP PL B356 450 PL B345 408 PR D51 2090 JPSJ 64 339	 R. Arnold et al. M. Alston-Garnjost et al. A. de Silva et al. M. Gunther et al. R. Arnold et al. A. Balysh et al. H. Ejiri et al. N. Takaoka, Y. Motomura, R.G. Arnold et al. 61 168. A. Balysh et al. A.S. Barabash et al. D. Dassie et al. H. Ejiri et al. M. Ejiri et al. 	(NEMO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY) (NEMO Collab.) (Heidelberg-Moscow Collab.) (ITEP, SCUC, PNL+) (NEMO Collab.) (OSAK, KIEV)
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DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA ARNOLD BALYSH BARABASH DASSIE EJIRI KOBAYASHI SUHONEN APTEMEV	97 97 96 96 96 96 95 95 95 95 95 95 95 95 95	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557 JETPL 61 170 Translated from ZETFP PL B356 450 PL B345 408 PR D51 2090 JPSJ 64 339 NP A586 457 PR C49 3055 JETPL 52 262	 R. Arnold et al. M. Alston-Garnjost et al. A. de Silva et al. M. Gunther et al. R. Arnold et al. A. Balysh et al. H. Ejiri et al. N. Takaoka, Y. Motomura, R.G. Arnold et al. 61 168. A. Balysh et al. A.S. Barabash et al. D. Dassie et al. H. Ejiri et al. M. Kobayashi, M. Kobayash J. Suhonen, O. Civitarese 	(NEMO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY) (NEMO Collab.) (Heidelberg-Moscow Collab.) (ITEP, SCUC, PNL+) (NEMO Collab.) (OSAK, KIEV) i (KEK, SAGA)
DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA ARNOLD BALYSH BARABASH DASSIE EJIRI KOBAYASHI SUHONEN ARTEMEV	97 97 96 96 96 95 95 95 95 95 95 95 95 94 93	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557 JETPL 61 170 Translated from ZETFP PL B356 450 PL B345 408 PR D51 2090 JPSJ 64 339 NP A586 457 PR C49 3055 JETPL 58 262 Translated from ZETEP	 R. Arnold et al. M. Alston-Garnjost et al. A. de Silva et al. M. Gunther et al. R. Arnold et al. A. Balysh et al. H. Ejiri et al. N. Takaoka, Y. Motomura, R.G. Arnold et al. 61 168. A. Balysh et al. A.S. Barabash et al. D. Dassie et al. H. Ejiri et al. M. Kobayashi, M. Kobayash J. Suhonen, O. Civitarese V.A. Artemiev et al. 58 256 	(NEMO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY) (NEMO Collab.) (Heidelberg-Moscow Collab.) (ITEP, SCUC, PNL+) (NEMO Collab.) (OSAK, KIEV) i (KEK, SAGA)
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DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA ARNOLD BALYSH BARABASH DASSIE EJIRI KOBAYASHI SUHONEN ARTEMEV BERNATOW KAWASHIMA	97 97 96 96 96 95 95 95 95 95 95 95 95 95 95 95 93 93 93	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557 JETPL 61 170 Translated from ZETFP PL B356 450 PL B345 408 PR D51 2090 JPSJ 64 339 NP A586 457 PR C49 3055 JETPL 58 262 Translated from ZETFP PR C47 806 PR C47 R2452	 K. Arnold et al. M. Alston-Garnjost et al. A. de Silva et al. M. Gunther et al. R. Arnold et al. A. Balysh et al. H. Ejiri et al. N. Takaoka, Y. Motomura, R.G. Arnold et al. 61 168. A. Balysh et al. A.S. Barabash et al. D. Dassie et al. H. Ejiri et al. M. Kobayashi, M. Kobayash J. Suhonen, O. Civitarese V.A. Artemiev et al. 58 256. T. Bernatowicz et al. A. Kawashima, K. Takahash 	(NEMO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY) (NEMO Collab.) (Heidelberg-Moscow Collab.) (NEMO Collab.) (ITEP, SCUC, PNL+) (NEMO Collab.) (OSAK, KIEV) i (KEK, SAGA) (ITEP, INRM) (WUSL, TATA) i, A. Masuda (TOKYC+)
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DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA ARNOLD BALYSH BARABASH DASSIE EJIRI KOBAYASHI SUHONEN ARTEMEV BERNATOW KAWASHIMA VUILLEUMIER BALYSH	97 97 96 96 96 95 95 95 95 95 95 95 95 95 94 93 93 93 93 92	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557 JETPL 61 170 Translated from ZETFP PL B356 450 PL B345 408 PR D51 2090 JPSJ 64 339 NP A586 457 PR C49 3055 JETPL 58 262 Translated from ZETFP PR C47 806 PR C47 R2452 PR D48 1009 PL B283 32	 K. Arnold et al. M. Alston-Garnjost et al. A. de Silva et al. M. Gunther et al. R. Arnold et al. A. Balysh et al. H. Ejiri et al. N. Takaoka, Y. Motomura, R.G. Arnold et al. 61 168. A. Balysh et al. A.S. Barabash et al. D. Dassie et al. H. Ejiri et al. M. Kobayashi, M. Kobayash J. Suhonen, O. Civitarese V.A. Artemiev et al. 58 256. T. Bernatowicz et al. A. Kawashima, K. Takahash J.C. Vuilleumier et al. A. Balysh et al. 	(NEMO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY) (NEMO Collab.) (Heidelberg-Moscow Collab.) (ITEP, SCUC, PNL+) (NEMO Collab.) (ITEP, SCUC, PNL+) (NEMO Collab.) (OSAK, KIEV) i (KEK, SAGA) (ITEP, INRM) (WUSL, TATA) i, A. Masuda (TOKYC+) (NEUC, CIT, VILL) (MPIH, KIAE, SASSO)
DESILVA GUENTHER ARNOLD BALYSH EJIRI TAKAOKA ARNOLD BALYSH BARABASH DASSIE EJIRI KOBAYASHI SUHONEN ARTEMEV BERNATOW BERNATOW	97 97 96 96 96 95 95 95 95 95 95 95 95 95 95 93 93 93 93 93 92 92	PR C55 474 PR C56 2451 PR D55 54 ZPHY C72 239 PRL 77 5186 NP A611 85 PR C53 1557 JETPL 61 170 Translated from ZETFP PL B356 450 PL B345 408 PR D51 2090 JPSJ 64 339 NP A586 457 PR C49 3055 JETPL 58 262 Translated from ZETFP PR C47 806 PR C47 R2452 PR D48 1009 PL B283 32 PRL 69 2341	 R. Arnold et al. M. Alston-Garnjost et al. A. de Silva et al. M. Gunther et al. R. Arnold et al. R. Balysh et al. H. Ejiri et al. N. Takaoka, Y. Motomura, R.G. Arnold et al. 61 168. A. Balysh et al. A.S. Barabash et al. D. Dassie et al. H. Ejiri et al. M. Kobayashi, M. Kobayash J. Suhonen, O. Civitarese V.A. Artemiev et al. 58 256. T. Bernatowicz et al. A. Kawashima, K. Takahash J.C. Vuilleumier et al. A. Balysh et al. T. Bernatowicz et al. T. Bernatowicz et al. 	(NEMO-2 Collab.) (LBL, MTHO+) (UCI) (Heidelberg-Moscow Collab.) (BCEN, CAEN, JINR+) (KIAE, UCI, CIT) (OSAK) K. Nagao (KYUSH, OKAY) (NEMO Collab.) (Heidelberg-Moscow Collab.) (NEMO Collab.) (ITEP, SCUC, PNL+) (NEMO Collab.) (OSAK, KIEV) i (KEK, SAGA) (ITEP, INRM) (WUSL, TATA) (MEUC, CIT, VILL) (MPIH, KIAE, SASSO) (WUSL, TATA)
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