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***W MASS***

To obtain the world average, common systematics between experiments are properly taken into account. The LEP average  $W$  mass based on published results is  $80.383 \pm 0.035$  GeV. The combined  $p\bar{p}$  collider data yields an average  $W$  mass of  $80.454 \pm 0.059$  GeV (ABAZOV 04D).

OUR FIT uses these average LEP and  $p\bar{p}$  collider  $W$  mass values together with the  $Z$  mass, the  $W$  to  $Z$  mass ratio, and mass difference measurements.

VALUE (GeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>80.403 \pm 0.029</math> OUR FIT</b>				
80.415 $\pm$ 0.042 $\pm$ 0.031	11830	<sup>1</sup> ABBIENDI	06	OPAL $E_{cm}^{ee} = 170\text{--}209$ GeV
80.270 $\pm$ 0.046 $\pm$ 0.031	9909	<sup>2</sup> ACHARD	06	L3 $E_{cm}^{ee} = 161\text{--}209$ GeV
80.440 $\pm$ 0.043 $\pm$ 0.027	8692	<sup>3</sup> SCHael	06	ALEP $E_{cm}^{ee} = 161\text{--}209$ GeV
80.483 $\pm$ 0.084	49247	<sup>4</sup> ABAZOV	02D D0	$E_{cm}^{pp} = 1.8$ TeV
80.359 $\pm$ 0.074 $\pm$ 0.049	3077	<sup>5</sup> ABREU	01K DLPH	$E_{cm}^{ee} = 161+172+183+189$ GeV
80.433 $\pm$ 0.079	53841	<sup>6</sup> AFFOLDER	01E CDF	$E_{cm}^{pp} = 1.8$ TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
82.87 $\pm$ 1.82 $\pm$ 0.30 $-0.16$	1500	<sup>7</sup> AKTAS	06 H1	$e^\pm p \rightarrow \bar{\nu}_e(\nu_e)X$ , $\sqrt{s} \approx 300$ GeV
80.3 $\pm$ 2.1 $\pm$ 1.2 $\pm$ 1.0	645	<sup>8</sup> CHEKANOV	02C ZEUS	$e^- p \rightarrow \nu_e X$ , $\sqrt{s} = 318$ GeV
81.4 $\pm$ 2.7 $-2.6$ $\pm$ 2.0 $\pm$ 3.3 $-3.0$	1086	<sup>9</sup> BREITWEG	00D ZEUS	$e^+ p \rightarrow \bar{\nu}_e X$ , $\sqrt{s} \approx 300$ GeV
80.84 $\pm$ 0.22 $\pm$ 0.83	2065	<sup>10</sup> ALITTI	92B UA2	See $W/Z$ ratio below
80.79 $\pm$ 0.31 $\pm$ 0.84		<sup>11</sup> ALITTI	90B UA2	$E_{cm}^{pp} = 546,630$ GeV
80.0 $\pm$ 3.3 $\pm$ 2.4	22	<sup>12</sup> ABE	89I CDF	$E_{cm}^{pp} = 1.8$ TeV
82.7 $\pm$ 1.0 $\pm$ 2.7	149	<sup>13</sup> ALBAJAR	89 UA1	$E_{cm}^{pp} = 546,630$ GeV
81.8 $\pm$ 6.0 $-5.3$ $\pm$ 2.6	46	<sup>14</sup> ALBAJAR	89 UA1	$E_{cm}^{pp} = 546,630$ GeV
89 $\pm$ 3 $\pm$ 6	32	<sup>15</sup> ALBAJAR	89 UA1	$E_{cm}^{pp} = 546,630$ GeV
81. $\pm$ 5.	6	ARNISON	83 UA1	$E_{cm}^{ee} = 546$ GeV
80. $\pm$ 10. $-6.$	4	BANNER	83B UA2	Repl. by ALITTI 90B

<sup>1</sup> ABBIENDI 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events. The result quoted here is obtained combining this mass value with the results using  $W^+ W^- \rightarrow \ell\nu_\ell\ell'\nu_{\ell'}$  events in the energy range 183–207 GeV (ABBIENDI 03C) and the dependence of the  $WW$  production cross-section on  $m_W$  at threshold. The systematic error includes  $\pm 0.009$  GeV due to the uncertainty on the LEP beam energy.

- <sup>2</sup> ACHARD 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events in the C.M. energy range 189–209 GeV. The result quoted here is obtained combining this mass value with the results obtained from a direct  $W$  mass reconstruction at 172 and 183 GeV and with those from the dependence of the  $WW$  production cross-section on  $m_W$  at 161 and 172 GeV (ACCIARRI 99).
- <sup>3</sup> SCHAEL 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events in the C.M. energy range 183–209 GeV. The result quoted here is obtained combining this mass value with those obtained from the dependence of the  $W$  pair production cross-section on  $m_W$  at 161 and 172 GeV (BARATE 97 and BARATE 97S respectively). The systematic error includes  $\pm 0.009$  GeV due to possible effects of final state interactions in the  $q\bar{q}q\bar{q}$  channel and  $\pm 0.009$  GeV due to the uncertainty on the LEP beam energy.
- <sup>4</sup> ABAZOV 02D improve the measurement of the  $W$ -boson mass including  $W \rightarrow e\nu_e$  events in which the electron is close to a boundary of a central electromagnetic calorimeter module. Properly combining the results obtained by fitting  $m_T(W)$ ,  $p_T(e)$ , and  $p_T(\nu)$ , this sample provides a mass value of  $80.574 \pm 0.405$  GeV. The value reported here is a combination of this measurement with all previous DØ  $W$ -boson mass measurements.
- <sup>5</sup> ABREU 01K obtain this value properly combining results obtained from a direct  $W$  mass reconstruction at 172, 183, and 189 GeV with those from measurements of  $W$ -pair production cross sections at 161, 172, and 183 GeV. The systematic error includes  $\pm 0.017$  GeV due to the beam energy uncertainty and  $\pm 0.033$  GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.
- <sup>6</sup> AFFOLDER 01E fit the transverse mass spectrum of 30115  $W \rightarrow e\nu_e$  events ( $M_W = 80.473 \pm 0.065 \pm 0.092$  GeV) and of 14740  $W \rightarrow \mu\nu_\mu$  events ( $M_W = 80.465 \pm 0.100 \pm 0.103$  GeV) obtained in the run IB (1994–95). Combining the electron and muon results, accounting for correlated uncertainties, yields  $M_W = 80.470 \pm 0.089$  GeV. They combine this value with their measurement of ABE 95P reported in run IA (1992–93) to obtain the quoted value.
- <sup>7</sup> AKTAS 06 fit the  $Q^2$  dependence ( $300 < Q^2 < 30,000$  GeV $^2$ ) of the charged-current differential cross section with a propagator mass. The first error is experimental and the second corresponds to uncertainties due to input parameters and model assumptions.
- <sup>8</sup> CHEKANOV 02C fit the  $Q^2$  dependence ( $200 < Q^2 < 60000$  GeV $^2$ ) of the charged-current differential cross sections with a propagator mass fit. The last error is due to the uncertainty on the probability density functions.
- <sup>9</sup> BREITWEG 00D fit the  $Q^2$  dependence ( $200 < Q^2 < 22500$  GeV $^2$ ) of the charged-current differential cross sections with a propagator mass fit. The last error is due to the uncertainty on the probability density functions.
- <sup>10</sup> ALITTI 92B result has two contributions to the systematic error ( $\pm 0.83$ ): one ( $\pm 0.81$ ) cancels in  $m_W/m_Z$  and one ( $\pm 0.17$ ) is noncancelling. These were added in quadrature. We choose the ALITTI 92B value without using the LEP  $m_Z$  value, because we perform our own combined fit.
- <sup>11</sup> There are two contributions to the systematic error ( $\pm 0.84$ ): one ( $\pm 0.81$ ) which cancels in  $m_W/m_Z$  and one ( $\pm 0.21$ ) which is non-cancelling. These were added in quadrature.
- <sup>12</sup> ABE 89I systematic error dominated by the uncertainty in the absolute energy scale.
- <sup>13</sup> ALBAJAR 89 result is from a total sample of 299  $W \rightarrow e\nu$  events.
- <sup>14</sup> ALBAJAR 89 result is from a total sample of 67  $W \rightarrow \mu\nu$  events.
- <sup>15</sup> ALBAJAR 89 result is from  $W \rightarrow \tau\nu$  events.

## W/Z MASS RATIO

The fit uses the  $W$  and  $Z$  mass, mass difference, and mass ratio measurements.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.88173±0.00032 OUR FIT</b>				
0.8821 ± 0.0011 ± 0.0008	28323	16 ABBOTT	98N D0	$E_{\text{cm}}^{pp} = 1.8 \text{ TeV}$
0.88114 ± 0.00154 ± 0.00252	5982	17 ABBOTT	98P D0	$E_{\text{cm}}^{pp} = 1.8 \text{ TeV}$
0.8813 ± 0.0036 ± 0.0019	156	18 ALITTI	92B UA2	$E_{\text{cm}}^{pp} = 630 \text{ GeV}$
16 ABBOTT 98N obtain this from a study of 28323 $W \rightarrow e\nu_e$ and 3294 $Z \rightarrow e^+e^-$ decays. Of this latter sample, 2179 events are used to calibrate the electron energy scale.				
17 ABBOTT 98P obtain this from a study of 5982 $W \rightarrow e\nu_e$ events. The systematic error includes an uncertainty of ±0.00175 due to the electron energy scale.				
18 Scale error cancels in this ratio.				

### $m_Z - m_W$

The fit uses the  $W$  and  $Z$  mass, mass difference, and mass ratio measurements.

<u>VALUE (GeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>10.785±0.029 OUR FIT</b>			
<b>10.4 ±1.4 ±0.8</b>	ALBAJAR	89 UA1	$E_{\text{cm}}^{pp} = 546,630 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
11.3 ±1.3 ±0.9	ANSARI	87 UA2	$E_{\text{cm}}^{pp} = 546,630 \text{ GeV}$

### $m_{W^+} - m_{W^-}$

Test of *CPT* invariance.

<u>VALUE (GeV)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>-0.19±0.58</b>	1722	ABE	90G CDF	$E_{\text{cm}}^{pp} = 1.8 \text{ TeV}$

## W WIDTH

The CDF and DØ widths labelled “extracted value” are obtained by measuring  $R = [\sigma(W)/\sigma(Z)] [\Gamma(W \rightarrow \ell\nu_\ell)] / [B(Z \rightarrow \ell\ell)\Gamma(W)]$  where the bracketed quantities can be calculated with plausible reliability.  $\Gamma(W)$  is then extracted by using a value of  $B(Z \rightarrow \ell\ell)$  measured at LEP. The UA1 and UA2 widths used  $R = [\sigma(W)/\sigma(Z)] [\Gamma(W \rightarrow \ell\nu_\ell)/\Gamma(Z \rightarrow \ell\ell)] \Gamma(Z)/\Gamma(W)$  and the measured value of  $\Gamma(Z)$ . The Standard Model prediction is  $2.0910 \pm 0.0015 \text{ GeV}$  (see Review on “Electroweak model and constraints on new physics” in this Edition).

To obtain OUR FIT, the correlation between systematics within LEP experiments and within Tevatron experiments is properly taken into account as given in the LEP note accessible at [http://lepewwg.web.cern.ch/LEPEWWG/leppw/mw/pdg\\_2006/](http://lepewwg.web.cern.ch/LEPEWWG/leppw/mw/pdg_2006/) and in the combined Tevatron paper of ABAZOV 04D. The respective average

values ( $2.164 \pm 0.085$  GeV from LEP and  $2.115 \pm 0.105$  GeV from Tevatron) yield an average  $W$  width of  $2.145 \pm 0.066$  GeV coming from direct measurements. ABAZOV 04D also determine the average extracted  $W$  width using CDF and DØ data to obtain a value of  $2.141 \pm 0.057$  GeV.

They further combine the Tevatron direct and extracted  $W$  widths to obtain an average Tevatron width of  $2.135 \pm 0.050$  GeV. Finally combining this with the LEP  $W$  width and the extracted  $W$  width values from UA1 and UA2 one obtains the quoted value.

VALUE (GeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.141±0.041 OUR FIT</b>					
1.996 $\pm 0.096 \pm 0.102$	10729	19	ABBIENDI	06	OPAL $E_{cm}^{ee} = 170\text{--}209$ GeV
2.18 $\pm 0.11 \pm 0.09$	9795	20	ACHARD	06	L3 $E_{cm}^{ee} = 172\text{--}209$ GeV
2.14 $\pm 0.09 \pm 0.06$	8717	21	SCHAEL	06	ALEP $E_{cm}^{ee} = 183\text{--}209$ GeV
2.23 $^{+0.15}_{-0.14} \pm 0.10$	294	22	ABAZOV	02E D0	Direct meas.
2.266 $\pm 0.176 \pm 0.076$	3005	23	ABREU	01K DLPH	$E_{cm}^{ee} = 183, 189$ GeV
2.152 $\pm 0.066$	79176	24	ABBOTT	00B D0	Extracted value
2.05 $\pm 0.10 \pm 0.08$	662	25	AFFOLDER	00M CDF	Direct meas.
2.064 $\pm 0.060 \pm 0.059$		26	ABE	95W CDF	Extracted value
2.10 $^{+0.14}_{-0.13} \pm 0.09$	3559	27	ALITTI	92 UA2	Extracted value
2.18 $^{+0.26}_{-0.24} \pm 0.04$		28	ALBAJAR	91 UA1	Extracted value
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2.30 $\pm 0.19 \pm 0.06$		29	ALITTI	90C UA2	Extracted value
2.8 $^{+1.4}_{-1.5} \pm 1.3$	149	30	ALBAJAR	89 UA1	$E_{cm}^{pp} = 546, 630$ GeV
< 7	90	119	APPEL	86 UA2	$E_{cm}^{pp} = 546, 630$ GeV
< 6.5	90	86	ARNISON	86 UA1	$E_{cm}^{pp} = 546, 630$ GeV

19 ABBIENDI 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events. The systematic error includes  $\pm 0.003$  GeV due to the uncertainty on the LEP beam energy.

20 ACHARD 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events in the C.M. energy range 189–209 GeV. The result quoted here is obtained combining this value of the width with the result obtained from a direct  $W$  mass reconstruction at 172 and 183 GeV (ACCIARRI 99).

21 SCHAEL 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events. The systematic error includes  $\pm 0.05$  GeV due to possible effects of final state interactions in the  $q\bar{q}q\bar{q}$  channel and  $\pm 0.01$  GeV due to the uncertainty on the LEP beam energy.

22 ABAZOV 02E obtain this result fitting the high-end tail (90–200 GeV) of the transverse-mass spectrum in semileptonic  $W \rightarrow e\nu_e$  decays.

23 ABREU 01K obtain this value properly combining results obtained at 183 and 189 GeV using  $WW \rightarrow \ell\bar{\nu}_\ell q\bar{q}$  and  $WW \rightarrow q\bar{q}q\bar{q}$  decays. The systematic error includes an uncertainty of  $\pm 0.052$  GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.

24 ABBOTT 00B measure  $R = 10.43 \pm 0.27$  for the  $W \rightarrow e\nu_e$  decay channel. They use the SM theoretical predictions for  $\sigma(W)/\sigma(Z)$  and  $\Gamma(W \rightarrow e\nu_e)$  and the world average for  $B(Z \rightarrow ee)$ . The value quoted here is obtained combining this result ( $2.169 \pm 0.070$  GeV) with that of ABBOTT 99H.

- <sup>25</sup> AFFOLDER 00M fit the high transverse mass (100–200 GeV)  $W \rightarrow e\nu_e$  and  $W \rightarrow \mu\nu_\mu$  events to obtain  $\Gamma(W) = 2.04 \pm 0.11(\text{stat}) \pm 0.09(\text{syst})$  GeV. This is combined with the earlier CDF measurement (ABE 95C) to obtain the quoted result.
- <sup>26</sup> ABE 95W measured  $R = 10.90 \pm 0.32 \pm 0.29$ . They use  $m_W = 80.23 \pm 0.18$  GeV,  $\sigma(W)/\sigma(Z) = 3.35 \pm 0.03$ ,  $\Gamma(W \rightarrow e\nu) = 225.9 \pm 0.9$  MeV,  $\Gamma(Z \rightarrow e^+ e^-) = 83.98 \pm 0.18$  MeV, and  $\Gamma(Z) = 2.4969 \pm 0.0038$  GeV.
- <sup>27</sup> ALITTI 92 measured  $R = 10.4^{+0.7}_{-0.6} \pm 0.3$ . The values of  $\sigma(Z)$  and  $\sigma(W)$  come from  $O(\alpha_s^2)$  calculations using  $m_W = 80.14 \pm 0.27$  GeV, and  $m_Z = 91.175 \pm 0.021$  GeV along with the corresponding value of  $\sin^2\theta_W = 0.2274$ . They use  $\sigma(W)/\sigma(Z) = 3.26 \pm 0.07 \pm 0.05$  and  $\Gamma(Z) = 2.487 \pm 0.010$  GeV.
- <sup>28</sup> ALBAJAR 91 measured  $R = 9.5^{+1.1}_{-1.0}$  (stat. + syst.).  $\sigma(W)/\sigma(Z)$  is calculated in QCD at the parton level using  $m_W = 80.18 \pm 0.28$  GeV and  $m_Z = 91.172 \pm 0.031$  GeV along with  $\sin^2\theta_W = 0.2322 \pm 0.0014$ . They use  $\sigma(W)/\sigma(Z) = 3.23 \pm 0.05$  and  $\Gamma(Z) = 2.498 \pm 0.020$  GeV. This measurement is obtained combining both the electron and muon channels.
- <sup>29</sup> ALITTI 90C used the same technique as described for ABE 90. They measured  $R = 9.38^{+0.82}_{-0.72} \pm 0.25$ , obtained  $\Gamma(W)/\Gamma(Z) = 0.902 \pm 0.074 \pm 0.024$ . Using  $\Gamma(Z) = 2.546 \pm 0.032$  GeV, they obtained the  $\Gamma(W)$  value quoted above and the limits  $\Gamma(W) < 2.56$  (2.64) GeV at the 90% (95%) CL.  $E_{\text{cm}}^{p\bar{p}} = 546,630$  GeV.
- <sup>30</sup> ALBAJAR 89 result is from a total sample of 299  $W \rightarrow e\nu$  events.
- <sup>31</sup> If systematic error is neglected, result is  $2.7^{+1.4}_{-1.5}$  GeV. This is enhanced subsample of 172 total events.

## $W^+$ DECAY MODES

$W^-$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1 \ell^+ \nu$	[a] $(10.80 \pm 0.09) \%$	
$\Gamma_2 e^+ \nu$	$(10.75 \pm 0.13) \%$	
$\Gamma_3 \mu^+ \nu$	$(10.57 \pm 0.15) \%$	
$\Gamma_4 \tau^+ \nu$	$(11.25 \pm 0.20) \%$	
$\Gamma_5$ hadrons	$(67.60 \pm 0.27) \%$	
$\Gamma_6 \pi^+ \gamma$	$< 8 \times 10^{-5}$	95%
$\Gamma_7 D_s^+ \gamma$	$< 1.3 \times 10^{-3}$	95%
$\Gamma_8 cX$	$(33.4 \pm 2.6) \%$	
$\Gamma_9 c\bar{s}$	$(31^{+13}_{-11}) \%$	
$\Gamma_{10}$ invisible	[b] $(1.4 \pm 2.8) \%$	

[a]  $\ell$  indicates each type of lepton ( $e$ ,  $\mu$ , and  $\tau$ ), not sum over them.

[b] This represents the width for the decay of the  $W$  boson into a charged particle with momentum below detectability,  $p < 200$  MeV.

## W PARTIAL WIDTHS

### $\Gamma(\text{invisible})$

### $\Gamma_{10}$

This represents the width for the decay of the  $W$  boson into a charged particle with momentum below detectability,  $p < 200$  MeV.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$30^{+52}_{-48} \pm 33$	32 BARATE 99I ALEP	$E_{\text{cm}}^{\text{ee}} = 161+172+183$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •	33 BARATE 99L ALEP	$E_{\text{cm}}^{\text{ee}} = 161+172+183$ GeV	

- 32 BARATE 99I measure this quantity using the dependence of the total cross section  $\sigma_{WW}$  upon a change in the total width. The fit is performed to the  $WW$  measured cross sections at 161, 172, and 183 GeV. This partial width is  $< 139$  MeV at 95%CL.
  - 33 BARATE 99L use  $W$ -pair production to search for effectively invisible  $W$  decays, tagging with the decay of the other  $W$  boson to Standard Model particles. The partial width for effectively invisible decay is  $< 27$  MeV at 95%CL.
- 

## W BRANCHING RATIOS

Overall fits are performed to determine the branching ratios of the  $W$ . LEP averages on  $W \rightarrow e\nu_e$ ,  $W \rightarrow \mu\nu_\mu$ , and  $W \rightarrow \tau\nu_\tau$ , and their correlations are first obtained by combining results from the four experiments taking properly into account the common systematics. The procedure is described in the note LEPEWWG/XSEC/2001-02, 30 March 2001, at <http://lepewwg.web.cern.ch/LEPEWWG/lepww/4f/PDG01>. The LEP average values so obtained, using published data, are given in the note LEPEWWG/XSEC/2005-01 accessible at <http://lepewwg.web.cern.ch/LEPEWWG/lepww/4f/PDG05/>. These results, together with results from the  $p\bar{p}$  colliders are then used in fits to obtain the world average  $W$  branching ratios. A first fit determines three individual leptonic branching ratios,  $B(W \rightarrow e\nu_e)$ ,  $B(W \rightarrow \mu\nu_\mu)$ , and  $B(W \rightarrow \tau\nu_\tau)$ . This fit has a  $\chi^2 = 4.7$  for 10 degrees of freedom. A second fit assumes lepton universality and determines the leptonic branching ratio  $B(W \rightarrow \ell\nu_\ell)$  and the hadronic branching ratio is derived as  $B(W \rightarrow \text{hadrons}) = 1-3B(W \rightarrow \ell\nu)$ . This fit has a  $\chi^2=11.3$  for 12 degrees of freedom.

The LEP  $W \rightarrow \ell\nu$  data are obtained by the Collaborations using individual leptonic channels and are, therefore, not included in the overall fits to avoid double counting.

### $\Gamma(\ell^+\nu)/\Gamma_{\text{total}}$

### $\Gamma_1/\Gamma$

$\ell$  indicates average over  $e$ ,  $\mu$ , and  $\tau$  modes, not sum over modes.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.1080 \pm 0.0009</math> OUR FIT</b>				
$0.1085 \pm 0.0014 \pm 0.0008$	13600	ABDALLAH	04G DLPH	$E_{\text{cm}}^{\text{ee}} = 161-209$ GeV
$0.1083 \pm 0.0014 \pm 0.0010$	11246	ACHARD	04J L3	$E_{\text{cm}}^{\text{ee}} = 161-209$ GeV
$0.1096 \pm 0.0012 \pm 0.0005$	16116	SCHAEL	04A ALEP	$E_{\text{cm}}^{\text{ee}} = 183-209$ GeV
$0.1056 \pm 0.0020 \pm 0.0009$	5778	ABBIENDI,G	00 OPAL	$E_{\text{cm}}^{\text{ee}} = 161+172+183$ +189 GeV
$0.1102 \pm 0.0052$	11858	<sup>34</sup> ABBOTT	99H D0	$E_{\text{cm}}^{p\bar{p}} = 1.8$ TeV
$0.104 \pm 0.008$	3642	<sup>35</sup> ABE	92I CDF	$E_{\text{cm}}^{p\bar{p}} = 1.8$ TeV

<sup>34</sup> ABBOTT 99H measure  $R \equiv [\sigma_W B(W \rightarrow \ell\nu_\ell)]/[\sigma_Z B(Z \rightarrow \ell\ell)] = 10.90 \pm 0.52$  combining electron and muon channels. They use  $M_W = 80.39 \pm 0.06$  GeV and the SM theoretical predictions for  $\sigma(W)/\sigma(Z)$  and  $B(Z \rightarrow \ell\ell)$ .

<sup>35</sup> 1216  $\pm 38^{+27}_{-31}$   $W \rightarrow \mu\nu$  events from ABE 92I and 2426  $W \rightarrow e\nu$  events of ABE 91C. ABE 92I give the inverse quantity as  $9.6 \pm 0.7$  and we have inverted.

### $\Gamma(e^+\nu)/\Gamma_{\text{total}}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
<b>0.1075 <math>\pm 0.0013</math> OUR FIT</b>					
0.1061 $\pm 0.0028$		<sup>36</sup> ABAZOV	04D TEVA	$E_{\text{cm}}^{p\bar{p}} = 1.8$ TeV	
0.1055 $\pm 0.0031 \pm 0.0014$	1804	ABDALLAH	04G DLPH	$E_{\text{cm}}^{ee} = 161\text{--}209$ GeV	
0.1078 $\pm 0.0029 \pm 0.0013$	1576	ACHARD	04J L3	$E_{\text{cm}}^{ee} = 161\text{--}209$ GeV	
0.1078 $\pm 0.0027 \pm 0.0010$	2142	SCHAEL	04A ALEP	$E_{\text{cm}}^{ee} = 183\text{--}209$ GeV	
0.1046 $\pm 0.0042 \pm 0.0014$	801	ABBIENDI,G	00 OPAL	$E_{\text{cm}}^{ee} = 161\text{--}172\text{--}183$ +189 GeV	
0.10 $\pm 0.014$ $+0.02$ $-0.03$	248	<sup>37</sup> ANSARI	87C UA2	$E_{\text{cm}}^{p\bar{p}} = 546,630$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
seen	119	APPEL	86 UA2	$E_{\text{cm}}^{p\bar{p}} = 546,630$ GeV	
seen	172	ARNISON	86 UA1	$E_{\text{cm}}^{p\bar{p}} = 546,630$ GeV	

<sup>36</sup> ABAZOV 04D take into account all correlations to properly combine the CDF (ABE 95W) and DØ (ABBOTT 00B) measurements of the ratio R in the electron channel. The ratio R is defined as  $[\sigma_W \cdot B(W \rightarrow e\nu_e)] / [\sigma_Z \cdot B(Z \rightarrow ee)]$ . The combination gives  $R^{\text{Tevatron}} = 10.59 \pm 0.23$ .  $\sigma_W / \sigma_Z$  is calculated at next-to-next-to-leading order ( $3.360 \pm 0.051$ ). The branching fraction  $B(Z \rightarrow ee)$  is taken from this Review as ( $3.363 \pm 0.004$ )%.

<sup>37</sup> The first error was obtained by adding the statistical and systematic experimental uncertainties in quadrature. The second error reflects the dependence on theoretical prediction of total  $W$  cross section:  $\sigma(546 \text{ GeV}) = 4.7^{+1.4}_{-0.7}$  nb and  $\sigma(630 \text{ GeV}) = 5.8^{+1.8}_{-1.0}$  nb. See ALTARELLI 85B.

### $\Gamma(\mu^+\nu)/\Gamma_{\text{total}}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
<b>0.1057 <math>\pm 0.0015</math> OUR FIT</b>					
0.1065 $\pm 0.0026 \pm 0.0008$	1998	ABDALLAH	04G DLPH	$E_{\text{cm}}^{ee} = 161\text{--}209$ GeV	
0.1003 $\pm 0.0029 \pm 0.0012$	1423	ACHARD	04J L3	$E_{\text{cm}}^{ee} = 161\text{--}209$ GeV	
0.1087 $\pm 0.0025 \pm 0.0008$	2216	SCHAEL	04A ALEP	$E_{\text{cm}}^{ee} = 183\text{--}209$ GeV	
0.1050 $\pm 0.0041 \pm 0.0012$	803	ABBIENDI,G	00 OPAL	$E_{\text{cm}}^{ee} =$ 161+172+183 +189 GeV	
0.10 $\pm 0.01$	1216	<sup>38</sup> ABE	92I CDF	$E_{\text{cm}}^{p\bar{p}} = 1.8$ TeV	

<sup>38</sup> ABE 92I quote the inverse quantity as  $9.9 \pm 1.2$  which we have inverted.

### $\Gamma(\tau^+\nu)/\Gamma_{\text{total}}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
<b>0.1125 <math>\pm 0.0020</math> OUR FIT</b>					
0.1146 $\pm 0.0039 \pm 0.0019$	2034	ABDALLAH	04G DLPH	$E_{\text{cm}}^{ee} = 161\text{--}209$ GeV	
0.1189 $\pm 0.0040 \pm 0.0020$	1375	ACHARD	04J L3	$E_{\text{cm}}^{ee} = 161\text{--}209$ GeV	
0.1125 $\pm 0.0032 \pm 0.0020$	2070	SCHAEL	04A ALEP	$E_{\text{cm}}^{ee} = 183\text{--}209$ GeV	
0.1075 $\pm 0.0052 \pm 0.0021$	794	ABBIENDI,G	00 OPAL	$E_{\text{cm}}^{ee} =$ 161+172+183 +189 GeV	

### $\Gamma(\text{hadrons})/\Gamma_{\text{total}}$

OUR FIT value is obtained by a fit to the lepton branching ratio data assuming lepton universality.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.6760±0.0027 OUR FIT</b>				
0.6745±0.0041±0.0024	13600	ABDALLAH	04G DLPH	$E_{\text{cm}}^{\text{ee}} = 161\text{--}209 \text{ GeV}$
0.6750±0.0042±0.0030	11246	ACHARD	04J L3	$E_{\text{cm}}^{\text{ee}} = 161\text{--}209 \text{ GeV}$
0.6713±0.0037±0.0015	16116	SCHAEL	04A ALEP	$E_{\text{cm}}^{\text{ee}} = 183\text{--}209 \text{ GeV}$
0.6832±0.0061±0.0028	5778	ABBIENDI,G	00 OPAL	$E_{\text{cm}}^{\text{ee}} = 161+172+183+189 \text{ GeV}$

### $\Gamma(\mu^+\nu)/\Gamma(e^+\nu)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.983±0.018 OUR FIT</b>				
0.89 ± 0.10	13k	<sup>39</sup> ABACHI	95D D0	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$
1.02 ± 0.08	1216	<sup>40</sup> ABE	92I CDF	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$
1.00 ± 0.14 ± 0.08	67	ALBAJAR	89 UA1	$E_{\text{cm}}^{p\bar{p}} = 546,630 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.24 <sup>+0.6</sup> <sub>-0.4</sub>	14	ARNISON	84D UA1	Repl. by ALBA-JAR 89

<sup>39</sup> ABACHI 95D obtain this result from the measured  $\sigma_W B(W \rightarrow \mu\nu) = 2.09 \pm 0.23 \pm 0.11 \text{ nb}$  and  $\sigma_W B(W \rightarrow e\nu) = 2.36 \pm 0.07 \pm 0.13 \text{ nb}$  in which the first error is the combined statistical and systematic uncertainty, the second reflects the uncertainty in the luminosity.

<sup>40</sup> ABE 92I obtain  $\sigma_W B(W \rightarrow \mu\nu) = 2.21 \pm 0.07 \pm 0.21$  and combine with ABE 91C  $\sigma_W B((W \rightarrow e\nu))$  to give a ratio of the couplings from which we derive this measurement.

### $\Gamma(\tau^+\nu)/\Gamma(e^+\nu)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.046±0.023 OUR FIT</b>				
0.961±0.061	980	<sup>41</sup> ABBOTT	00D D0	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$
0.94 ± 0.14	179	<sup>42</sup> ABE	92E CDF	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$
1.04 ± 0.08 ± 0.08	754	<sup>43</sup> ALITTI	92F UA2	$E_{\text{cm}}^{p\bar{p}} = 630 \text{ GeV}$
1.02 ± 0.20 ± 0.12	32	ALBAJAR	89 UA1	$E_{\text{cm}}^{p\bar{p}} = 546,630 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.995±0.112±0.083	198	ALITTI	91C UA2	Repl. by ALITTI 92F
1.02 ± 0.20 ± 0.10	32	ALBAJAR	87 UA1	Repl. by ALBA-JAR 89

<sup>41</sup> ABBOTT 00D measure  $\sigma_W \times B(W \rightarrow \tau\nu_\tau) = 2.22 \pm 0.09 \pm 0.10 \pm 0.10 \text{ nb}$ . Using the ABBOTT 00B result  $\sigma_W \times B(W \rightarrow e\nu_e) = 2.31 \pm 0.01 \pm 0.05 \pm 0.10 \text{ nb}$ , they quote the ratio of the couplings from which we derive this measurement.

<sup>42</sup> ABE 92E use two procedures for selecting  $W \rightarrow \tau\nu_\tau$  events. The missing  $E_T$  trigger leads to  $132 \pm 14 \pm 8$  events and the  $\tau$  trigger to  $47 \pm 9 \pm 4$  events. Proper statistical and systematic correlations are taken into account to arrive at  $\sigma B(W \rightarrow \tau\nu) = 2.05 \pm 0.27 \text{ nb}$ . Combined with ABE 91C result on  $\sigma B(W \rightarrow e\nu)$ , ABE 92E quote a ratio of the couplings from which we derive this measurement.

<sup>43</sup> This measurement is derived by us from the ratio of the couplings of ALITTI 92F.

### $\Gamma_5/\Gamma$

### $\Gamma(\pi^+\gamma)/\Gamma(e^+\nu)$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma_2$
$< 7 \times 10^{-4}$	95	ABE	98H	CDF	$E_{cm}^{pp} = 1.8 \text{ TeV}$
$< 4.9 \times 10^{-3}$	95	<sup>44</sup> ALITTI	92D	UA2	$E_{cm}^{pp} = 630 \text{ GeV}$
$< 58 \times 10^{-3}$	95	<sup>45</sup> ALBAJAR	90	UA1	$E_{cm}^{pp} = 546, 630 \text{ GeV}$
<sup>44</sup> ALITTI 92D limit is $3.8 \times 10^{-3}$ at 90%CL.					
<sup>45</sup> ALBAJAR 90 obtain $< 0.048$ at 90%CL.					

### $\Gamma(D_s^+\gamma)/\Gamma(e^+\nu)$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma_2$
$< 1.2 \times 10^{-2}$	95	ABE	98P	CDF	$E_{cm}^{pp} = 1.8 \text{ TeV}$

### $\Gamma(cX)/\Gamma(\text{hadrons})$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma_5$
<b>0.49 ± 0.04 OUR AVERAGE</b>					
0.481 ± 0.042 ± 0.032	3005	<sup>46</sup> ABBIENDI	00V	OPAL	$E_{cm}^{ee} = 183 + 189 \text{ GeV}$
0.51 ± 0.05 ± 0.03	746	<sup>47</sup> BARATE	99M	ALEP	$E_{cm}^{ee} = 172 + 183 \text{ GeV}$
<sup>46</sup> ABBIENDI 00V tag $W \rightarrow cX$ decays using measured jet properties, lifetime information, and leptons produced in charm decays. From this result, and using the additional measurements of $\Gamma(W)$ and $B(W \rightarrow \text{hadrons})$ , $ V_{cs} $ is determined to be $0.969 \pm 0.045 \pm 0.036$ .					
<sup>47</sup> BARATE 99M tag $c$ jets using a neural network algorithm. From this measurement $ V_{cs} $ is determined to be $1.00 \pm 0.11 \pm 0.07$ .					

### $R_{cs} = \Gamma(c\bar{s})/\Gamma(\text{hadrons})$

VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma_5$
<b>0.46<sup>+0.18</sup><sub>-0.14</sub> ± 0.07</b>	<sup>48</sup> ABREU	98N	DLPH	$E_{cm}^{ee} = 161+172 \text{ GeV}$

<sup>48</sup> ABREU 98N tag  $c$  and  $s$  jets by identifying a charged kaon as the highest momentum particle in a hadronic jet. They also use a lifetime tag to independently identify a  $c$  jet, based on the impact parameter distribution of charged particles in a jet. From this measurement  $|V_{cs}|$  is determined to be  $0.94^{+0.32}_{-0.26} \pm 0.13$ .

## AVERAGE PARTICLE MULTIPLICITIES IN HADRONIC $W$ DECAY

Summed over particle and antiparticle, when appropriate.

### $\langle N_{\pi^\pm} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_{10}/\Gamma_5$
<b>15.70 ± 0.35</b>	<sup>49</sup> ABREU,P	00F	DLPH	$E_{cm}^{ee} = 189 \text{ GeV}$
<sup>49</sup> ABREU,P 00F measure $\langle N_{\pi^\pm} \rangle = 31.65 \pm 0.48 \pm 0.76$ and $15.51 \pm 0.38 \pm 0.40$ in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.				

### $\langle N_{K^\pm} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_{11}/\Gamma_5$
<b>2.20 ± 0.19</b>	<sup>50</sup> ABREU,P	00F	DLPH	$E_{cm}^{ee} = 189 \text{ GeV}$
<sup>50</sup> ABREU,P 00F measure $\langle N_{K^\pm} \rangle = 4.38 \pm 0.42 \pm 0.12$ and $2.23 \pm 0.32 \pm 0.17$ in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.				

**$\langle N_p \rangle$** 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.92±0.14</b>	51 ABREU,P	00F DLPH	$E_{\text{cm}}^{\text{ee}} = 189 \text{ GeV}$

51 ABREU,P 00F measure  $\langle N_p \rangle = 1.82 \pm 0.29 \pm 0.16$  and  $0.94 \pm 0.23 \pm 0.06$  in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.

 **$\langle N_{\text{charged}} \rangle$** 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>19.39±0.08 OUR AVERAGE</b>			

$19.38 \pm 0.05 \pm 0.08$	52 ABBIENDI	06A OPAL	$E_{\text{cm}}^{\text{ee}} = 189-209 \text{ GeV}$
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$19.44 \pm 0.17$	53 ABREU,P	00F DLPH	$E_{\text{cm}}^{\text{ee}} = 183+189 \text{ GeV}$
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$19.3 \pm 0.3 \pm 0.3$	54 ABBIENDI	99N OPAL	$E_{\text{cm}}^{\text{ee}} = 183 \text{ GeV}$
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$19.23 \pm 0.74$	55 ABREU	98C DLPH	$E_{\text{cm}}^{\text{ee}} = 172 \text{ GeV}$
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52 ABBIENDI 06A measure  $\langle N_{\text{charged}} \rangle = 38.74 \pm 0.12 \pm 0.26$  when both  $W$  bosons decay hadronically and  $\langle N_{\text{charged}} \rangle = 19.39 \pm 0.11 \pm 0.09$  when one  $W$  boson decays semileptonically. The value quoted here is obtained under the assumption that there is no color reconnection between  $W$  bosons; the value is a weighted average taking into account correlations in the systematic uncertainties.

53 ABREU,P 00F measure  $\langle N_{\text{charged}} \rangle = 39.12 \pm 0.33 \pm 0.36$  and  $38.11 \pm 0.57 \pm 0.44$  in the fully hadronic final states at 189 and 183 GeV respectively, and  $\langle N_{\text{charged}} \rangle = 19.49 \pm 0.31 \pm 0.27$  and  $19.78 \pm 0.49 \pm 0.43$  in the semileptonic final states. The value quoted is a weighted average without assuming any correlations.

54 ABBIENDI 99N use the final states  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  to derive this value.

55 ABREU 98C combine results from both the fully hadronic as well semileptonic  $WW$  final states after demonstrating that the  $W$  decay charged multiplicity is independent of the topology within errors.

**TRIPLE GAUGE COUPLINGS (TGC'S)**

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 **$g_1^Z$** 

OUR FIT below is obtained by combining the measurements taking into account properly the common systematic errors (see LEPEWWG/TGC/2005-01 at <http://lepewwg.web.cern.ch/LEPEWWG/lepww/tgc>).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.984<sup>+0.022</sup><sub>-0.019</sub> OUR FIT</b>				

$1.001 \pm 0.027 \pm 0.013$	9310	56 SCHABEL	05A ALEP	$E_{\text{cm}}^{\text{ee}} = 183-209 \text{ GeV}$
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$0.987^{+0.034}_{-0.033}$	9800	57 ABBIENDI	04D OPAL	$E_{\text{cm}}^{\text{ee}} = 183-209 \text{ GeV}$
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$0.966^{+0.034}_{-0.032} \pm 0.015$	8325	58 ACHARD	04D L3	$E_{\text{cm}}^{\text{ee}} = 161-209 \text{ GeV}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

	2.3	59 ABAZOV	05S D0	$E_{\text{cm}}^{pp} = 1.96 \text{ TeV}$
$0.98 \pm 0.07 \pm 0.01$	2114	60 ABREU	01I DLPH	$E_{\text{cm}}^{\text{ee}} = 183+189 \text{ GeV}$
	331	61 ABBOTT	99I D0	$E_{\text{cm}}^{pp} = 1.8 \text{ TeV}$

- 56 SCHael 05A study single-photon, single- $W$ , and  $WW$ -pair production from 183 to 209 GeV. The result quoted here is derived from the  $WW$ -pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 57 ABBIENDI 04D combine results from  $W^+ W^-$  in all decay channels. Only  $CP$ -conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is  $0.923 < g_1^Z < 1.054$ .
- 58 ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained from the  $WW$ -pair production sample including data from 161 to 183 GeV, ACCIA-RRI 99Q. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 59 ABAZOV 05S study  $\bar{p}p \rightarrow WZ$  production with a subsequent trilepton decay to  $\ell\nu\ell'\bar{\ell}'$  ( $\ell$  and  $\ell' = e$  or  $\mu$ ). Three events (estimated background  $0.71 \pm 0.08$  events) with  $WZ$  decay characteristics are observed from which they derive limits on the anomalous  $WWZ$  couplings. The 95% CL limit for a form factor scale  $\Lambda = 1.5$  TeV is  $0.51 < g_1^Z < 1.66$ , fixing  $\lambda_Z$  and  $\kappa_Z$  to their Standard Model values.
- 60 ABREU 01I combine results from  $e^+ e^-$  interactions at 189 GeV leading to  $W^+ W^-$  and  $We\nu_e$  final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is  $0.84 < g_1^Z < 1.13$ .
- 61 ABBOTT 99I perform a simultaneous fit to the  $W\gamma$ ,  $WW \rightarrow$  dilepton,  $WW/WZ \rightarrow e\nu jj$ ,  $WW/WZ \rightarrow \mu\nu jj$ , and  $WZ \rightarrow$  trilepton data samples. For  $\Lambda = 2.0$  TeV, the 95%CL limits are  $0.63 < g_1^Z < 1.57$ , fixing  $\lambda_Z$  and  $\kappa_Z$  to their Standard Model values, and assuming Standard Model values for the  $WW\gamma$  couplings.

 **$\kappa_\gamma$** 

OUR FIT below is obtained by combining the measurements taking into account properly the common systematic errors (see LEPEWWG/TGC/2005-01 at <http://lepewwg.web.cern.ch/LEPEWWG/leppw/tgc>).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.973^{+0.044}_{-0.045}</math> OUR FIT</b>				
$0.971 \pm 0.055 \pm 0.030$	10689	62 SCHael	05A ALEP	$E_{cm}^{ee} = 183\text{--}209$ GeV
$0.88^{+0.09}_{-0.08}$	9800	63 ABBIENDI	04D OPAL	$E_{cm}^{ee} = 183\text{--}209$ GeV
$1.013^{+0.067}_{-0.064} \pm 0.026$	10575	64 ACHARD	04D L3	$E_{cm}^{ee} = 161\text{--}209$ GeV
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
	17	65 ABAZOV	06H D0	$E_{cm}^{pp} = 1.96$ TeV
	141	66 ABAZOV	05J D0	$E_{cm}^{pp} = 1.96$ TeV
$1.25^{+0.21}_{-0.20} \pm 0.06$	2298	67 ABREU	01I DLPH	$E_{cm}^{ee} = 183\text{+}189$ GeV
		68 BREITWEG	00 ZEUS	$e^+ p \rightarrow e^+ W^\pm X$ , $\sqrt{s} \approx 300$ GeV
$0.92 \pm 0.34$	331	69 ABBOTT	99I D0	$E_{cm}^{pp} = 1.8$ TeV

- 62 SCHael 05A study single-photon, single- $W$ , and  $WW$ -pair production from 183 to 209 GeV. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 63 ABBIENDI 04D combine results from  $W^+ W^-$  in all decay channels. Only  $CP$ -conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is  $0.73 < \kappa_\gamma < 1.07$ .
- 64 ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained including data from 161 to 183 GeV, ACCIARRI 99Q. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 65 ABAZOV 06H study  $\bar{p}p \rightarrow WW$  production with a subsequent decay  $WW \rightarrow e^+ \nu_e e^- \bar{\nu}_e$ ,  $WW \rightarrow e^\pm \nu_e \mu^\mp \nu_\mu$  or  $WW \rightarrow \mu^+ \nu_\mu \mu^- \bar{\nu}_\mu$ . The 95% C.L. limit for a form factor scale  $\Lambda = 1$  TeV is  $-0.05 < \kappa_\gamma < 2.29$ , fixing  $\lambda_\gamma = 0$ . With the assumption that the  $WW\gamma$  and  $WWZ$  couplings are equal the 95% C.L. one-dimensional limit ( $\Lambda = 2$  TeV) is  $0.68 < \kappa < 1.45$ .
- 66 ABAZOV 05J perform a likelihood fit to the photon  $E_T$  spectrum of  $W\gamma + X$  events, where the  $W$  decays to an electron or muon which is required to be well separated from the photon. For  $\Lambda = 2.0$  TeV the 95% CL limits are  $0.12 < \kappa_\gamma < 1.96$ . In the fit  $\lambda_\gamma$  is kept fixed to its Standard Model value.
- 67 ABREU 01I combine results from  $e^+ e^-$  interactions at 189 GeV leading to  $W^+ W^-$ ,  $W e \nu_e$ , and  $\nu \bar{\nu} \gamma$  final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is  $0.87 < \kappa_\gamma < 1.68$ .
- 68 BREITWEG 00 search for  $W$  production in events with large hadronic  $p_T$ . For  $p_T > 20$  GeV, the upper limit on the cross section gives the 95%CL limit  $-3.7 < \kappa_\gamma < 2.5$  (for  $\lambda_\gamma = 0$ ).
- 69 ABBOTT 99I perform a simultaneous fit to the  $W\gamma$ ,  $WW \rightarrow$  dilepton,  $WW/WZ \rightarrow e\nu jj$ ,  $WW/WZ \rightarrow \mu\nu jj$ , and  $WZ \rightarrow$  trilepton data samples. For  $\Lambda = 2.0$  TeV, the 95%CL limits are  $0.75 < \kappa_\gamma < 1.39$ .

$\lambda_\gamma$

OUR FIT below is obtained by combining the measurements taking into account properly the common systematic errors (see LEPEWWG/TGC/2005-01 at <http://lepewwg.web.cern.ch/LEPEWWG/leppww/tgc>).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.028^{+0.020}_{-0.021}</math> OUR FIT</b>				
$-0.012 \pm 0.027 \pm 0.011$	10689	70 SCHael	05A ALEP	$E_{cm}^{ee} = 183\text{--}209$ GeV
$-0.060^{+0.034}_{-0.033}$	9800	71 ABBIENDI	04D OPAL	$E_{cm}^{ee} = 183\text{--}209$ GeV
$-0.021^{+0.035}_{-0.034} \pm 0.017$	10575	72 ACHARD	04D L3	$E_{cm}^{ee} = 161\text{--}209$ GeV
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
	17	73 ABAZOV	06H D0	$E_{cm}^{p\bar{p}} = 1.96$ TeV
	141	74 ABAZOV	05J D0	$E_{cm}^{p\bar{p}} = 1.96$ TeV
$0.05 \pm 0.09 \pm 0.01$	2298	75 ABREU	01I DLPH	$E_{cm}^{ee} = 183\text{+}189$ GeV
		76 BREITWEG	00 ZEUS	$e^+ p \rightarrow e^+ W^\pm X$ , $\sqrt{s} \approx 300$ GeV
$0.00^{+0.10}_{-0.09}$	331	77 ABBOTT	99I D0	$E_{cm}^{p\bar{p}} = 1.8$ TeV

- 70 SCHAEL 05A study single-photon, single- $W$ , and  $WW$ -pair production from 183 to 209 GeV. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 71 ABBIENDI 04D combine results from  $W^+ W^-$  in all decay channels. Only  $CP$ -conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is  $-0.13 < \lambda_\gamma < 0.01$ .
- 72 ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained including data from 161 to 183 GeV, ACCIARRI 99Q. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 73 ABAZOV 06H study  $\bar{p}p \rightarrow WW$  production with a subsequent decay  $WW \rightarrow e^+ \nu_e e^- \bar{\nu}_e$ ,  $WW \rightarrow e^\pm \nu_e \mu^\mp \nu_\mu$  or  $WW \rightarrow \mu^+ \nu_\mu \mu^- \bar{\nu}_\mu$ . The 95% C.L. limit for a form factor scale  $\Lambda = 1$  TeV is  $-0.97 < \lambda_\gamma < 1.04$ , fixing  $\kappa_\gamma = 1$ . With the assumption that the  $WW\gamma$  and  $WWZ$  couplings are equal the 95% C.L. one-dimensional limit ( $\Lambda = 2$  TeV) is  $-0.29 < \lambda < 0.30$ .
- 74 ABAZOV 05J perform a likelihood fit to the photon  $E_T$  spectrum of  $W\gamma + X$  events, where the  $W$  decays to an electron or muon which is required to be well separated from the photon. For  $\Lambda = 2.0$  TeV the 95% CL limits are  $-0.20 < \lambda_\gamma < 0.20$ . In the fit  $\kappa_\gamma$  is kept fixed to its Standard Model value.
- 75 ABREU 01I combine results from  $e^+ e^-$  interactions at 189 GeV leading to  $W^+ W^-$ ,  $W e \nu_e$ , and  $\nu \bar{\nu} \gamma$  final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is  $-0.11 < \lambda_\gamma < 0.23$ .
- 76 BREITWEG 00 search for  $W$  production in events with large hadronic  $p_T$ . For  $p_T > 20$  GeV, the upper limit on the cross section gives the 95%CL limit  $-3.2 < \lambda_\gamma < 3.2$  for  $\kappa_\gamma$  fixed to its Standard Model value.
- 77 ABBOTT 99I perform a simultaneous fit to the  $W\gamma$ ,  $WW \rightarrow$  dilepton,  $WW/WZ \rightarrow e\nu jj$ ,  $WW/WZ \rightarrow \mu\nu jj$ , and  $WZ \rightarrow$  trilepton data samples. For  $\Lambda = 2.0$  TeV, the 95%CL limits are  $-0.18 < \lambda_\gamma < 0.19$ .

 **$\kappa_Z$** 

This coupling is  $CP$ -conserving ( $C$ - and  $P$ - separately conserving).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.924^{+0.059}_{-0.056} \pm 0.024</math></b>	7171	78 ACHARD	04D L3	$E_{cm}^{ee} = 189\text{--}209$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

17	79 ABAZOV	06H D0	$E_{cm}^{PP} = 1.96$ TeV
2.3	80 ABAZOV	05S D0	$E_{cm}^{PP} = 1.96$ TeV

- 78 ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the  $WW$ -pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 79 ABAZOV 06H study  $\bar{p}p \rightarrow WW$  production with a subsequent decay  $WW \rightarrow e^+ \nu_e e^- \bar{\nu}_e$ ,  $WW \rightarrow e^\pm \nu_e \mu^\mp \nu_\mu$  or  $WW \rightarrow \mu^+ \nu_\mu \mu^- \bar{\nu}_\mu$ . The 95% C.L. limit for a form factor scale  $\Lambda = 2$  TeV is  $0.55 < \kappa_Z < 1.55$ , fixing  $\lambda_Z = 0$ . With the assumption that the  $WW\gamma$  and  $WWZ$  couplings are equal the 95% C.L. one-dimensional limit ( $\Lambda = 2$  TeV) is  $0.68 < \kappa < 1.45$ .
- 80 ABAZOV 05S study  $\bar{p}p \rightarrow WZ$  production with a subsequent trilepton decay to  $\ell\nu\ell'\bar{\nu}$  ( $\ell$  and  $\ell' = e$  or  $\mu$ ). Three events (estimated background  $0.71 \pm 0.08$  events) with  $WZ$  decay characteristics are observed from which they derive limits on the anomalous  $WWZ$  couplings. The 95% CL limit for a form factor scale  $\Lambda = 1$  TeV is  $-1.0 < \kappa_Z < 3.4$ , fixing  $\lambda_Z$  and  $g_1^Z$  to their Standard Model values.

$\lambda_Z$ 

This coupling is *CP*-conserving (*C*- and *P*- separately conserving).

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$-0.088^{+0.060}_{-0.057} \pm 0.023$	7171	81 ACHARD	04D L3	$E_{\text{cm}}^{\text{ee}} = 189\text{--}209 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

17	82 ABAZOV	06H D0	$E_{\text{cm}}^{p\bar{p}} = 1.96 \text{ TeV}$
2.3	83 ABAZOV	05S D0	$E_{\text{cm}}^{p\bar{p}} = 1.96 \text{ TeV}$

81 ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the  $WW$ -pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.

82 ABAZOV 06H study  $\bar{p}p \rightarrow WW$  production with a subsequent decay  $WW \rightarrow e^+ \nu_e e^- \bar{\nu}_e$ ,  $WW \rightarrow e^\pm \nu_e \mu^\mp \nu_\mu$  or  $WW \rightarrow \mu^+ \nu_\mu \mu^- \bar{\nu}_\mu$ . The 95% C.L. limit for a form factor scale  $\Lambda = 2 \text{ TeV}$  is  $-0.39 < \lambda_Z < 0.39$ , fixing  $\kappa_Z = 1$ . With the assumption that the  $WW\gamma$  and  $WWZ$  couplings are equal the 95% C.L. one-dimensional limit ( $\Lambda = 2 \text{ TeV}$ ) is  $-0.29 < \lambda < 0.30$ .

83 ABAZOV 05S study  $\bar{p}p \rightarrow WZ$  production with a subsequent trilepton decay to  $\ell\nu\ell'\bar{\ell}'$  ( $\ell$  and  $\ell' = e$  or  $\mu$ ). Three events (estimated background  $0.71 \pm 0.08$  events) with  $WZ$  decay characteristics are observed from which they derive limits on the anomalous  $WWZ$  couplings. The 95% CL limit for a form factor scale  $\Lambda = 1.5 \text{ TeV}$  is  $-0.48 < \lambda_Z < 0.48$ , fixing  $g_1^Z$  and  $\kappa_Z$  to their Standard Model values.

 $g_5^Z$ 

This coupling is *CP*-conserving but *C*- and *P*-violating.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>0.93 \pm 0.09 \text{ OUR AVERAGE}</math></b>				Error includes scale factor of 1.1.
$0.96^{+0.13}_{-0.12}$	9800	84 ABBIENDI	04D OPAL	$E_{\text{cm}}^{\text{ee}} = 183\text{--}209 \text{ GeV}$
$1.00 \pm 0.13 \pm 0.05$	7171	85 ACHARD	04D L3	$E_{\text{cm}}^{\text{ee}} = 189\text{--}209 \text{ GeV}$

$0.56^{+0.23}_{-0.22} \pm 0.12$       1154      86 ACCIARRI      99Q L3       $E_{\text{cm}}^{\text{ee}} = 161 + 172 + 183 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.84 \pm 0.23$       87 EBOLI      00 THEO LEP1, SLC+ Tevatron

84 ABBIENDI 04D combine results from  $W^+ W^-$  in all decay channels. Only *CP*-conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is  $0.72 < g_5^Z < 1.21$ .

85 ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the  $WW$ -pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.

86 ACCIARRI 99Q study  $W$ -pair, single- $W$ , and single photon events.

87 EBOLI 00 extract this indirect value of the coupling studying the non-universal one-loop contributions to the experimental value of the  $Z \rightarrow b\bar{b}$  width ( $\Lambda=1 \text{ TeV}$  is assumed).

$g_4^Z$ 

This coupling is *CP*-violating (*C*-violating and *P*-conserving).

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>-0.02<sup>+0.32</sup><sub>-0.33</sub></b>	1065	88	ABBIENDI	01H OPAL $E_{\text{cm}}^{\text{ee}} = 189 \text{ GeV}$

88 ABBIENDI 01H study  $W$ -pair events, with one leptonically and one hadronically decaying  $W$ . The coupling is extracted using information from the  $W$  production angle together with decay angles from the leptonically decaying  $W$ .

 $\tilde{\kappa}_Z$ 

This coupling is *CP*-violating (*C*-conserving and *P*-violating).

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>-0.20<sup>+0.10</sup><sub>-0.07</sub></b>	1065	89	ABBIENDI	01H OPAL $E_{\text{cm}}^{\text{ee}} = 189 \text{ GeV}$

89 ABBIENDI 01H study  $W$ -pair events, with one leptonically and one hadronically decaying  $W$ . The coupling is extracted using information from the  $W$  production angle together with decay angles from the leptonically decaying  $W$ .

 $\tilde{\lambda}_Z$ 

This coupling is *CP*-violating (*C*-conserving and *P*-violating).

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>-0.18<sup>+0.24</sup><sub>-0.16</sub></b>	1065	90	ABBIENDI	01H OPAL $E_{\text{cm}}^{\text{ee}} = 189 \text{ GeV}$

90 ABBIENDI 01H study  $W$ -pair events, with one leptonically and one hadronically decaying  $W$ . The coupling is extracted using information from the  $W$  production angle together with decay angles from the leptonically decaying  $W$ .

## **$W$ ANOMALOUS MAGNETIC MOMENT**

The full magnetic moment is given by  $\mu_W = e(1+\kappa + \lambda)/2m_W$ . In the Standard Model, at tree level,  $\kappa = 1$  and  $\lambda = 0$ . Some papers have defined  $\Delta\kappa = 1-\kappa$  and assume that  $\lambda = 0$ . Note that the electric quadrupole moment is given by  $-e(\kappa-\lambda)/m_W^2$ . A description of the parameterization of these moments and additional references can be found in HAGIWARA 87 and BAUR 88. The parameter  $\Lambda$  appearing in the theoretical limits below is a regularization cutoff which roughly corresponds to the energy scale where the structure of the  $W$  boson becomes manifest.

<u>VALUE (<math>e/2m_W</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>2.22<sup>+0.20</sup><sub>-0.19</sub></b>	2298	91	ABREU	01I DLPH $E_{\text{cm}}^{\text{ee}} = 183+189 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

92 ABE	95G	CDF
93 ALITTI	92C	UA2
94 SAMUEL	92	THEO
95 SAMUEL	91	THEO
96 GRIFOLS	88	THEO
97 GROTCHE	87	THEO
98 VANDERBIJ	87	THEO
99 GRAU	85	THEO
100 SUZUKI	85	THEO
101 HERZOG	84	THEO

- 91 ABREU 01I combine results from  $e^+e^-$  interactions at 189 GeV leading to  $W^+W^-$ ,  $We\nu_e$ , and  $\nu\bar{\nu}\gamma$  final states with results from ABREU 99L at 183 GeV to determine  $\Delta g_1^Z$ ,  $\Delta\kappa_\gamma$ , and  $\lambda_\gamma$ .  $\Delta\kappa_\gamma$  and  $\lambda_\gamma$  are simultaneously floated in the fit to determine  $\mu_W$ .
- 92 ABE 95G report  $-1.3 < \kappa < 3.2$  for  $\lambda=0$  and  $-0.7 < \lambda < 0.7$  for  $\kappa=1$  in  $p\bar{p} \rightarrow e\nu_e\gamma X$  and  $\mu\nu_\mu\gamma X$  at  $\sqrt{s} = 1.8$  TeV.
- 93 ALITTI 92C measure  $\kappa = 1^{+2.6}_{-2.2}$  and  $\lambda = 0^{+1.7}_{-1.8}$  in  $p\bar{p} \rightarrow e\nu\gamma + X$  at  $\sqrt{s} = 630$  GeV. At 95%CL they report  $-3.5 < \kappa < 5.9$  and  $-3.6 < \lambda < 3.5$ .
- 94 SAMUEL 92 use preliminary CDF and UA2 data and find  $-2.4 < \kappa < 3.7$  at 96%CL and  $-3.1 < \kappa < 4.2$  at 95%CL respectively. They use data for  $W\gamma$  production and radiative  $W$  decay.
- 95 SAMUEL 91 use preliminary CDF data for  $p\bar{p} \rightarrow W\gamma X$  to obtain  $-11.3 \leq \Delta\kappa \leq 10.9$ . Note that their  $\kappa = 1 - \Delta\kappa$ .
- 96 GRIFOLS 88 uses deviation from  $\rho$  parameter to set limit  $\Delta\kappa \lesssim 65 (m_W^2/\Lambda^2)$ .
- 97 GROTCHE 87 finds the limit  $-37 < \Delta\kappa < 73.5$  (90% CL) from the experimental limits on  $e^+e^- \rightarrow \nu\bar{\nu}\gamma$  assuming three neutrino generations and  $-19.5 < \Delta\kappa < 56$  for four generations. Note their  $\Delta\kappa$  has the opposite sign as our definition.
- 98 VANDERBIJ 87 uses existing limits to the photon structure to obtain  $|\Delta\kappa| < 33 (m_W/\Lambda)$ . In addition VANDERBIJ 87 discusses problems with using the  $\rho$  parameter of the Standard Model to determine  $\Delta\kappa$ .
- 99 GRAU 85 uses the muon anomaly to derive a coupled limit on the anomalous magnetic dipole and electric quadrupole ( $\lambda$ ) moments  $1.05 > \Delta\kappa \ln(\Lambda/m_W) + \lambda/2 > -2.77$ . In the Standard Model  $\lambda = 0$ .
- 100 SUZUKI 85 uses partial-wave unitarity at high energies to obtain  $|\Delta\kappa| \lesssim 190 (m_W/\Lambda)^2$ . From the anomalous magnetic moment of the muon, SUZUKI 85 obtains  $|\Delta\kappa| \lesssim 2.2/\ln(\Lambda/m_W)$ . Finally SUZUKI 85 uses deviations from the  $\rho$  parameter and obtains a very qualitative, order-of-magnitude limit  $|\Delta\kappa| \lesssim 150 (m_W/\Lambda)^4$  if  $|\Delta\kappa| \ll 1$ .
- 101 HERZOG 84 consider the contribution of  $W$ -boson to muon magnetic moment including anomalous coupling of  $WW\gamma$ . Obtain a limit  $-1 < \Delta\kappa < 3$  for  $\Lambda \gtrsim 1$  TeV.

## ANOMALOUS $W/Z$ QUARTIC COUPLINGS

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#### $a_0/\Lambda^2$ , $a_c/\Lambda^2$ , $a_n/\Lambda^2$

Using the  $WW\gamma$  final state, the LEP combined 95% CL limits on the anomalous contributions to the  $WW\gamma\gamma$  and  $WWZ\gamma$  vertices (as of summer 2003) are given below:

(See P. Wells, “Experimental Tests of the Standard Model,” Int. Europhysics Conference on High-Energy Physics, Aachen, Germany, 17–23 July 2003)

$$\begin{aligned} -0.02 &< a_0^W/\Lambda^2 < 0.02 \text{ GeV}^{-2}, \\ -0.05 &< a_c^W/\Lambda^2 < 0.03 \text{ GeV}^{-2}, \\ -0.15 &< a_n^W/\Lambda^2 < 0.15 \text{ GeV}^{-2}. \end{aligned}$$

VALUE	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •		

102	ABBIENDI	04B	OPAL
103	ABBIENDI	04L	OPAL
104	HEISTER	04A	ALEP
105	ABDALLAH	03I	DLPH
106	ACHARD	02F	L3

- 102 ABBIENDI 04B select 187  $e^+ e^- \rightarrow W^+ W^- \gamma$  events in the C.M. energy range 180–209 GeV, where  $E_\gamma > 2.5$  GeV, the photon has a polar angle  $|\cos\theta_\gamma| < 0.975$  and is well isolated from the nearest jet and charged lepton, and the effective masses of both fermion-antifermion systems agree with the  $W$  mass within  $3\Gamma_W$ . The measured differential cross section as a function of the photon energy and photon polar angle is used to extract the 95% CL limits:  $-0.020 \text{ GeV}^{-2} < a_0/\Lambda^2 < 0.020 \text{ GeV}^{-2}$ ,  $-0.053 \text{ GeV}^{-2} < a_c/\Lambda^2 < 0.037 \text{ GeV}^{-2}$  and  $-0.16 \text{ GeV}^{-2} < a_n/\Lambda^2 < 0.15 \text{ GeV}^{-2}$ .
- 103 ABBIENDI 04L select 20  $e^+ e^- \rightarrow \nu\bar{\nu}\gamma\gamma$  acoplanar events in the energy range 180–209 GeV and 176  $e^+ e^- \rightarrow q\bar{q}\gamma\gamma$  events in the energy range 130–209 GeV. These samples are used to constrain possible anomalous  $W^+ W^- \gamma\gamma$  and  $Z Z \gamma\gamma$  quartic couplings. Further combining with the  $W^+ W^- \gamma$  sample of ABBIENDI 04B the following one-parameter 95% CL limits are obtained:  $-0.007 < a_0^Z/\Lambda^2 < 0.023 \text{ GeV}^{-2}$ ,  $-0.029 < a_c^Z/\Lambda^2 < 0.029 \text{ GeV}^{-2}$ ,  $-0.020 < a_0^W/\Lambda^2 < 0.020 \text{ GeV}^{-2}$ ,  $-0.052 < a_c^W/\Lambda^2 < 0.037 \text{ GeV}^{-2}$ .
- 104 In the CM energy range 183 to 209 GeV HEISTER 04A select 30  $e^+ e^- \rightarrow \nu\bar{\nu}\gamma\gamma$  events with two acoplanar, high energy and high transverse momentum photons. The photon-photon acoplanarity is required to be  $> 5^\circ$ ,  $E_\gamma/\sqrt{s} > 0.025$  (the more energetic photon having energy  $> 0.2\sqrt{s}$ ),  $p_{T\gamma}/E_{\text{beam}} > 0.05$  and  $|\cos\theta_\gamma| < 0.94$ . A likelihood fit to the photon energy and recoil missing mass yields the following one-parameter 95% CL limits:  $-0.012 < a_0^Z/\Lambda^2 < 0.019 \text{ GeV}^{-2}$ ,  $-0.041 < a_c^Z/\Lambda^2 < 0.044 \text{ GeV}^{-2}$ ,  $-0.060 < a_0^W/\Lambda^2 < 0.055 \text{ GeV}^{-2}$ ,  $-0.099 < a_c^W/\Lambda^2 < 0.093 \text{ GeV}^{-2}$ .
- 105 ABDALLAH 03I select 122  $e^+ e^- \rightarrow W^+ W^- \gamma$  events in the C.M. energy range 189–209 GeV, where  $E_\gamma > 5$  GeV, the photon has a polar angle  $|\cos\theta_\gamma| < 0.95$  and is well isolated from the nearest charged fermion. A fit to the photon energy spectra yields  $a_c/\Lambda^2 = 0.000^{+0.019}_{-0.040} \text{ GeV}^{-2}$ ,  $a_0/\Lambda^2 = -0.004^{+0.018}_{-0.010} \text{ GeV}^{-2}$ ,  $\tilde{a}_0/\Lambda^2 = -0.007^{+0.019}_{-0.008} \text{ GeV}^{-2}$ ,  $a_n/\Lambda^2 = -0.09^{+0.16}_{-0.05} \text{ GeV}^{-2}$ , and  $\tilde{a}_n/\Lambda^2 = +0.05^{+0.07}_{-0.15} \text{ GeV}^{-2}$ , keeping the other parameters fixed to their Standard Model values (0). The 95% CL limits are:  $-0.063 \text{ GeV}^{-2} < a_c/\Lambda^2 < +0.032 \text{ GeV}^{-2}$ ,  $-0.020 \text{ GeV}^{-2} < a_0/\Lambda^2 < +0.020 \text{ GeV}^{-2}$ ,  $-0.020 \text{ GeV}^{-2} < \tilde{a}_0/\Lambda^2 < +0.020 \text{ GeV}^{-2}$ ,  $-0.18 \text{ GeV}^{-2} < a_n/\Lambda^2 < +0.14 \text{ GeV}^{-2}$ ,  $-0.16 \text{ GeV}^{-2} < \tilde{a}_n/\Lambda^2 < +0.17 \text{ GeV}^{-2}$ .
- 106 ACHARD 02F select 86  $e^+ e^- \rightarrow W^+ W^- \gamma$  events at 192–207 GeV, where  $E_\gamma > 5$  GeV and the photon is well isolated. They also select 43 acoplanar  $e^+ e^- \rightarrow \nu\bar{\nu}\gamma\gamma$  events in this energy range, where the photon energies are  $> 5$  GeV and  $> 1$  GeV and the photon polar angles are between  $14^\circ$  and  $166^\circ$ . All these 43 events are in the recoil mass region corresponding to the  $Z$  (75–110 GeV). Using the shape and normalization of the photon spectra in the  $W^+ W^- \gamma$  events, and combining with the 42 event sample from 189 GeV data (ACCIARRI 00T), they obtain:  $a_0/\Lambda^2 = 0.000 \pm 0.010 \text{ GeV}^{-2}$ ,  $a_c/\Lambda^2 = -0.013 \pm 0.023 \text{ GeV}^{-2}$ , and  $a_n/\Lambda^2 = -0.002 \pm 0.076 \text{ GeV}^{-2}$ . Further combining the analyses of  $W^+ W^- \gamma$  events with the low recoil mass region of  $\nu\bar{\nu}\gamma\gamma$  events (including samples collected at 183 + 189 GeV), they obtain the following one-parameter 95% CL limits:  $-0.015 \text{ GeV}^{-2} < a_0/\Lambda^2 < 0.015 \text{ GeV}^{-2}$ ,  $-0.048 \text{ GeV}^{-2} < a_c/\Lambda^2 < 0.026 \text{ GeV}^{-2}$ , and  $-0.14 \text{ GeV}^{-2} < a_n/\Lambda^2 < 0.13 \text{ GeV}^{-2}$ .

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Also		PR D52 4784	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	95W	PR D52 2624	F. Abe <i>et al.</i>	(CDF Collab.)
Also		PRL 73 220	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	92E	PRL 68 3398	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	92I	PRL 69 28	F. Abe <i>et al.</i>	(CDF Collab.)
ALITTI	92	PL B276 365	J. Alitti <i>et al.</i>	(UA2 Collab.)
ALITTI	92B	PL B276 354	J. Alitti <i>et al.</i>	(UA2 Collab.)
ALITTI	92C	PL B277 194	J. Alitti <i>et al.</i>	(UA2 Collab.)

ALITTI	92D	PL B277 203	J. Alitti <i>et al.</i>	(UA2 Collab.)
ALITTI	92F	PL B280 137	J. Alitti <i>et al.</i>	(UA2 Collab.)
SAMUEL	92	PL B280 124	M.A. Samuel <i>et al.</i>	(OKSU, CARL)
ABE	91C	PR D44 29	F. Abe <i>et al.</i>	(CDF Collab.)
ALBAJAR	91	PL B253 503	C. Albajar <i>et al.</i>	(UA1 Collab.)
ALITTI	91C	ZPHY C52 209	J. Alitti <i>et al.</i>	(UA2 Collab.)
SAMUEL	91	PRL 67 9	M.A. Samuel <i>et al.</i>	(OKSU, CARL)
Also		PRL 67 2920 (erratum)	M.A. Samuel <i>et al.</i>	
ABE	90	PRL 64 152	F. Abe <i>et al.</i>	(CDF Collab.)
Also		PR D44 29	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	90G	PRL 65 2243	F. Abe <i>et al.</i>	(CDF Collab.)
Also		PR D43 2070	F. Abe <i>et al.</i>	(CDF Collab.)
ALBAJAR	90	PL B241 283	C. Albajar <i>et al.</i>	(UA1 Collab.)
ALITTI	90B	PL B241 150	J. Alitti <i>et al.</i>	(UA2 Collab.)
ALITTI	90C	ZPHY C47 11	J. Alitti <i>et al.</i>	(UA2 Collab.)
ABE	89I	PRL 62 1005	F. Abe <i>et al.</i>	(CDF Collab.)
ALBAJAR	89	ZPHY C44 15	C. Albajar <i>et al.</i>	(UA1 Collab.)
BAUR	88	NP B308 127	U. Baur, D. Zeppenfeld	(FSU, WISC)
GRIFOLS	88	IJMP A3 225	J.A. Grifols, S. Peris, J. Sola	(BARC, DESY)
Also		PL B197 437	J.A. Grifols, S. Peris, J. Sola	(BARC, DESY)
ALBAJAR	87	PL B185 233	C. Albajar <i>et al.</i>	(UA1 Collab.)
ANSARI	87	PL B186 440	R. Ansari <i>et al.</i>	(UA2 Collab.)
ANSARI	87C	PL B194 158	R. Ansari <i>et al.</i>	(UA2 Collab.)
GROTCHE	87	PR D36 2153	H. Grotch, R.W. Robinett	(PSU)
HAGIWARA	87	NP B282 253	K. Hagiwara <i>et al.</i>	(KEK, UCLA, FSU)
VANDERBIJ	87	PR D35 1088	J.J. van der Bij	(FNAL)
APPEL	86	ZPHY C30 1	J.A. Appel <i>et al.</i>	(UA2 Collab.)
ARNISON	86	PL 166B 484	G.T.J. Arnison <i>et al.</i>	(UA1 Collab.) J
ALTARELLI	85B	ZPHY C27 617	G. Altarelli, R.K. Ellis, G. Martinelli	(CERN+)
GRAU	85	PL 154B 283	A. Grau, J.A. Grifols	(BARC)
SUZUKI	85	PL 153B 289	M. Suzuki	(LBL)
ARNISON	84D	PL 134B 469	G.T.J. Arnison <i>et al.</i>	(UA1 Collab.)
HERZOG	84	PL 148B 355	F. Herzog	(WISC)
Also		PL 155B 468 (erratum)	F. Herzog	(WISC)
ARNISON	83	PL 122B 103	G.T.J. Arnison <i>et al.</i>	(UA1 Collab.)
BANNER	83B	PL 122B 476	M. Banner <i>et al.</i>	(UA2 Collab.)

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