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W MASS

To obtain the world average, common systematics between experiments are properly taken into account. The procedure for averaging the LEP data is given in the note LEPEWWG/MASS/2002-01 (March 11, 2002), accessible at http://lepewwg.web.cern.ch/LEPEWWG/lepww/mw/pdg_2002/. The LEP average W mass based on published results is 80.400 ± 0.056 GeV. The combined $p\overline{p}$ collider data yields an average W mass of 80.454 ± 0.059 GeV (ABAZOV 04D).

OUR FIT uses these average LEP and $p\overline{p}$ collider W mass values together with the Z mass, the W to Z mass ratio, and mass difference measurements.

<i>VALUE</i> (GeV)	EVTS	DOCUMENT ID	TECN	COMMENT
80.425± 0.038 OUR F	IT			
$80.41 \pm 0.41 \pm 0.13$	1101	$^{ m 1}$ ABBIENDI	03C OPAL	$E_{\rm cm}^{\it ee} = 183-207 \; {\rm GeV}$
$80.483 \pm \ 0.084$	49247	² ABAZOV	02D D0	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
$80.432 \pm 0.066 \pm 0.045$	2789	³ ABBIENDI	01F OPAL	CIII
80.359± 0.074±0.049	3077	⁴ ABREU	01K DLPH	$+189 \text{ GeV} \ E_{\text{cm}}^{ee} = 161 + 172 + 183 \ +189 \text{ GeV}$
80.433± 0.079	53841	⁵ AFFOLDER	01E CDF	$E_{cm}^{ar{p}} = 1.8 \; TeV$
$80.418 \pm 0.061 \pm 0.047$	2977	⁶ BARATE	00T ALEP	E ^{ee} _{cm} = 161+172+183 +189 GeV
80.61 ± 0.15	801	⁷ ACCIARRI	99 L3	Fee = 161+172+ 183 GeV
• • • We do not use th	ie followii	ng data for average	s, fits, limits	
$80.3 \pm 2.1 \pm 1.2 \pm 1.0$	645	⁸ CHEKANOV	02c ZEUS	$e^- p \rightarrow \nu_e X, \sqrt{s} = 318$
$79.9 \pm \ 2.2 \pm 2.3$	700	⁹ ADLOFF	01A H1	$e^{-}\stackrel{GeV}{p} \rightarrow \nu_{e} X, \ \sqrt{s} \approx 320 \ GeV$
80.482± 0.091	45394	¹⁰ ABBOTT	00 D0	Repl. by ABAZOV 02D
$80.9 \pm 3.7 \pm 3.7$	700	¹¹ ADLOFF	00B H1	$e^+ p \rightarrow \overline{\nu}_e X, \sqrt{s} \approx 300 \text{ GeV}$
$81.4^{+2.7}_{-2.6}\pm2.0^{+3.3}_{-3.0}$	1086	¹² BREITWEG	00D ZEUS	
$80.38 \pm 0.12 \pm 0.05$	701	¹³ ABBIENDI	99c OPAL	
$80.270 \pm 0.137 \pm 0.048$	809	¹⁴ ABREU	99T DLPH	I Repl. by ABREU 01K
$80.423 \pm \ 0.112 \pm 0.054$	812	¹⁵ BARATE	99 ALEP	Repl. by BARATE 00T
$80.80 \ \ ^{+} \ \ 0.48 \ \ \pm 0.03$	20	¹⁶ ACCIARRI	97 L3	Repl. by ACCIARRI 99
$80.5 \begin{array}{c} + & 1.4 \\ - & 2.4 \end{array} \pm 0.3$	94	¹⁷ ACCIARRI	97M L3	Repl. by ACCIARRI 99
$80.71 \ ^{+}_{-} \ 0.34 \ \pm 0.09$	101	¹⁸ ACCIARRI	97s L3	Repl. by ACCIARRI 99

80.41	± 0.18		8986	¹⁹ ABE	95P CDF	Repl. by AF- FOLDER 01E
80.84	± 0.22	± 0.83	2065	²⁰ ALITTI	92B UA2	See W/Z ratio below
80.79	\pm 0.31	± 0.84		²¹ ALITTI	90B UA2	$E_{\rm cm}^{p\overline{p}}$ = 546,630 GeV
80.0	\pm 3.3	± 2.4	22	²² ABE	891 CDF	$E_{cm}^{ar{p}} = 1.8 \; TeV$
82.7	\pm 1.0	±2.7	149	²³ ALBAJAR	89 UA1	$E_{ m cm}^{p\overline{p}} =$ 546,630 GeV
81.8	$^{+}$ 6.0 $^{-}$ 5.3	±2.6	46	²⁴ ALBAJAR	89 UA1	$E_{\rm cm}^{p \overline{p}} = 546,630 \; {\rm GeV}$
89	± 3	± 6	32	²⁵ ALBAJAR	89 UA1	$E_{\rm cm}^{p\overline{p}}$ = 546,630 GeV
81.	± 5.		6	ARNISON	83 UA1	<i>E</i> ^{ee} _{cm} = 546 GeV
80.	+10. $-6.$		4	BANNER	83B UA2	Repl. by ALITTI 90B

- ¹ ABBIENDI 03C determine the mass of the W boson using fully leptonic decays $W^+W^- \to \ell\nu_\ell\ell'\nu_{\ell'}$. They use the measured energies of the charged leptons and an approximate kinematic reconstruction of the event (both neutrinos are assumed in the same plane as the charged leptons) to get a W pseudo-mass. All these variables are combined in a simultaneous maximum likelihood fit. The systematic error is dominated by the uncertainty on the lepton energy.
- ²ABAZOV 02D improve the measurement of the W-boson mass including $W \to e \nu_e$ events in which the electron is close to a boundary of a central electromagnetic calorimeter module. Properly combining the results obtained by fitting $m_T(W)$, $p_T(e)$, and $p_T(\nu)$, this sample provides a mass value of 80.574 \pm 0.405 GeV. The value reported here is a combination of this measurement with all previous DØ W-boson mass measurements.
- 3 ABBIENDI 01F obtain this value properly combining results obtained from a direct W mass reconstruction at 172, 183, and 189 GeV with that from measurement of the W-pair production cross section at 161 GeV. The systematic error includes ± 0.017 GeV due to LEP energy uncertainty and ± 0.028 GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.
- ⁴ ABREU 01K obtain this value properly combining results obtained from a direct W mass reconstruction at 172, 183, and 189 GeV with those from measurements of W-pair production cross sections at 161, 172, and 183 GeV. The systematic error includes ± 0.017 GeV due to the beam energy uncertainty and ± 0.033 GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.
- 5 AFFOLDER 01E fit the transverse mass spectrum of 30115 $W\to e\nu_e$ events $(M_W=80.473\pm0.065\pm0.092~{\rm GeV})$ and of 14740 $W\to \mu\nu_\mu$ events $(M_W=80.465\pm0.100\pm0.103~{\rm GeV})$ obtained in the run IB (1994-95). Combining the electron and muon results, accounting for correlated uncertainties, yields $M_W=80.470\pm0.089~{\rm GeV}$. They combine this value with their measurement of ABE 95P reported in run IA (1992-93) to obtain the quoted value.
- 6 BARATE 00T obtain this value properly combining results obtained from a direct W mass reconstruction at 172, 183, and 189 GeV with those from measurements of W-pair production cross sections at 161 and 172 GeV. The systematic error includes ± 0.017 GeV due to LEP energy uncertainty and ± 0.019 GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.
- ⁷ ACCIARRI 99 obtain this value properly combining results obtained from a direct W mass reconstruction at 172 and 183 GeV with those from the measurements of the total W-pair production cross sections at 161 and 172 GeV. The value of the mass obtained from the direct reconstruction at 172 and 183 GeV is $M(W) = 80.58 \pm 0.14 \pm 0.08$ GeV.
- ⁸ CHEKANOV 02C fit the Q^2 dependence (200< Q^2 <60000 GeV 2) of the charged-current differential cross sections with a propagator mass fit. The last error is due to the uncertainty on the probability density functions.
- 9 ADLOFF 01A fit the Q^2 dependence (150 $< Q^2 <$ 30000 GeV 2) of the charged-current double-differential cross sections with a propagator mass fit. The second error includes 2.1 GeV due to the theoretical uncertainties.

- ¹⁰ ABBOTT 00 use $W \to e \nu_e$ events to measure the W mass with a fit to the transverse mass distribution. The result quoted here corresponds to electrons detected both in the forward and in the central calorimeters for the data recorded in 1992–1995. For the large rapidity electrons recorded in 1994–1995, the analysis combines results obtained from m_T , $p_T(e)$, and $p_T(\nu)$.
- ¹¹ ADLOFF 00B fit the Q^2 dependence (300 $< Q^2 <$ 15000 GeV²) of the charged-current double-differential cross sections with a propagator mass fit. The second error is due to the theoretical uncertainties
- the theoretical uncertainties, 12 BREITWEG 00D fit the Q^2 dependence (200 $< Q^2 <$ 22500 GeV 2) of the charged-current differential cross sections with a propagator mass fit. The last error is due to the uncertainty on the probability density functions.
- 13 ABBIENDI 99C obtain this value properly combining results from a direct W mass reconstruction at 172 and 183 GeV with that from the measurement of the total W-pair production cross section at 161 GeV. The systematic error includes an uncertainty of ± 0.02 GeV due to the possible color-reconnection and Bose-Einstein effects in the purely hadronic final states and an uncertainty of ± 0.02 GeV due to the beam energy.
- 14 ABREU 99T obtain this value properly combining results obtained from a direct W mass reconstruction at 172 and 183 GeV with those from measurement of W-pair production cross sections at 161, 172, and 183 GeV. The systematic error includes ± 0.021 GeV due to the beam energy uncertainty and ± 0.030 GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.
- 15 BARATE 99 obtain this value properly combining results from a direct W mass reconstruction at 172 and 183 GeV with those from the measurements of the total W-pair production cross sections at 161 and 172 GeV. The systematic error includes ± 0.023 GeV due to LEP energy uncertainty and ± 0.021 GeV due to theory uncertainty on account of possible color reconnection and Bose-Einstein correlations.
- 16 ACCIARRI 97 derive this value from their measured W-W production cross section $\sigma_{WW}=2.89^{+0.81}_{-0.70}\pm0.14$ pb using the Standard Model dependence of σ_{WW} on M_W at the given c.m. energy. Statistical and systematic errors are added in quadrature and the last error of ±0.03 GeV arises from the beam energy uncertainty. The same result is given by a fit of the production cross sections to the data.
- 17 ACCIARRI 97M derive this value from their measured WW production cross section $\sigma_{WW}=12.27^{+1.41}_{-1.32}\pm0.23$ pb using the Standard Model dependence of σ_{WW} on M_W at the given c.m. energy. Combining with ACCIARRI 97 authors find $M(W)=80.78^{+0.45}_{-0.41}\pm0.03$ GeV where the last error is due to beam energy uncertainty.
- 18 ACCIARRI 97S obtain this value from a fit to the reconstructed W mass distribution. The W width was taken as its Standard Model value at the fitted W mass. When both W mass and width are varied they obtain $M(W)=80.72 {+} 0.31 \pm 0.09$ GeV. The systematic error includes ± 0.03 GeV due to the beam energy uncertainty and ± 0.05 GeV due to the possible color reconnection and Bose-Einstein effects in the purely hadronic final state. Combining with ACCIARRI 97 and ACCIARRI 97M authors find: $M(W)=80.75 {+} 0.26 \pm 0.03$ (LEP) GeV.
- 19 ABE 95P use 3268 $W\to \mu\nu_{\mu}$ events to find $M=80.310\pm0.205\pm0.130$ GeV and 5718 $W\to e\nu_{e}$ events to find $M=80.490\pm0.145\pm0.175$ GeV. The result given here combines these while accounting for correlated uncertainties.
- ²⁰ ALITTI 92B result has two contributions to the systematic error (± 0.83); one (± 0.81) cancels in m_W/m_Z and one (± 0.17) is noncancelling. These were added in quadrature. We choose the ALITTI 92B value without using the LEP m_Z value, because we perform our own combined fit.
- There are two contributions to the systematic error (± 0.84) : one (± 0.81) which cancels in m_W/m_Z and one (± 0.21) which is non-cancelling. These were added in quadrature.
- ABE 891 systematic error dominated by the uncertainty in the absolute energy scale.
- 23 ALBAJAR 89 result is from a total sample of 299 W
 ightarrow e
 u events.
- 24 ALBAJAR 89 result is from a total sample of 67 $W
 ightarrow ~\mu
 u$ events.
- ²⁵ ALBAJAR 89 result is from $W \rightarrow \tau \nu$ events.

W/Z MASS RATIO

The fit uses the W and Z mass, mass difference, and mass ratio measurements.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.88197±0.00042 OUR FIT				
$0.8821\ \pm0.0011\ \pm0.0008$	28323	²⁶ ABBOTT	98N D0	$E_{cm}^{oldsymbol{p}\overline{oldsymbol{p}}}$ $= 1.8\;TeV$
$0.88114 \!\pm\! 0.00154 \!\pm\! 0.00252$	5982	²⁷ ABBOTT	98P D0	$E_{cm}^{oldsymbol{p}\overline{oldsymbol{p}}}$ $= 1.8\;TeV$
$0.8813 \pm 0.0036 \pm 0.0019$	156	²⁸ ALITTI	92B UA2	$E_{ m cm}^{p\overline{p}}=$ 630 GeV

 $^{^{26}}$ ABBOTT 98N obtain this from a study of 28323 $W\to e\nu_e$ and 3294 $Z\to e^+e^-$ decays. Of this latter sample, 2179 events are used to calibrate the electron energy scale.

$m_Z - m_W$

The fit uses the W and Z mass, mass difference, and mass ratio measurements.

VALUE	(GeV)		DOCUMENT ID		TECN	COMMENT
10.76	3±0.03	8 OUR FIT				
10.4	±1.4	±0.8	ALBAJAR	89	UA1	$E_{\rm cm}^{p\overline{p}}$ = 546,630 GeV
• • •	We do	$not\ use\ the\ following$	data for averages	s, fits	, limits,	etc. • • •
11.3	± 1.3	± 0.9	ANSARI	87	UA2	$E_{cm}^{p\overline{p}} = 546,630 \; GeV$

$$m_{W^+} - m_{W^-}$$

Test of CPT invariance.

VALUE (GeV)	EVTS	DOCUMENT ID	TECN	COMMENT
-0.19 ± 0.58	1722	ABE	90G CDF	$E_{cm}^{p\overline{p}} = 1.8 \text{ TeV}$

W WIDTH

The CDF and DØ widths labelled "extracted value" are obtained by measuring $R = [\sigma(W)/\sigma(Z)] \ [\Gamma(W \to \ell \nu_\ell)]/(\mathsf{B}(Z \to \ell \ell)\Gamma(W))$ where the bracketed quantities can be calculated with plausible reliability. $\Gamma(W)$ is then extracted by using a value of $\mathsf{B}(Z \to \ell \ell)$ measured at LEP. The UA1 and UA2 widths used $R = [\sigma(W)/\sigma(Z)] \ [\Gamma(W \to \ell \nu_\ell)/\Gamma(Z \to \ell \ell)] \ \Gamma(Z)/\Gamma(W)$ and the measured value of $\Gamma(Z)$. The Standard Model prediction is 2.0921 \pm 0.0025 GeV (see Review on "Electroweak model and constraints on new physics" in this Edition).

To obtain OUR FIT, the correlation between systematics within LEP experiments and within Tevatron experiments is properly taken into account as given in the LEP note LEPEWWG/MASS/2002-01 (http://lepewwg.web.cern.ch/LEPEWWG/lepww/mw/pdg_2002/) and the combined Tevatron paper of ABAZOV 04D. The respective average values (2.17 \pm 0.12 GeV from LEP and 2.115 \pm 0.105 GeV from Tevatron)

²⁷ ABBOTT 98P obtain this from a study of 5982 $W \rightarrow e \nu_e$ events. The systematic error includes an uncertainty of ± 0.00175 due to the electron energy scale.

²⁸ Scale error cancels in this ratio.

yield an average W width of 2.139 \pm 0.079 GeV coming from direct measurements. ABAZOV 04D also determine the average extracted W width using CDF and DØ data to obtain a value of 2.141 \pm 0.057 GeV.

They further combine the Tevatron direct and extracted W widths to obtain an average Tevatron width of 2.135 ± 0.050 GeV. Finally combining this with the LEP W width and the extracted W width values from UA1 and UA2 one obtains the quoted value.

VALUE (GeV)	CL%	EVTS	DOCUMENT ID		TECN	COMMENT
2.138 ± 0.044	OUR FIT					
$2.23 \begin{array}{l} +0.15 \\ -0.14 \end{array}$	±0.10	294	²⁹ ABAZOV	02E	D0	Direct meas.
2.04 ± 0.16	± 0.09	2756	³⁰ ABBIENDI	01F	OPAL	E ^{ee} _{cm} = 172+183 +189 GeV
2.266 ± 0.176	± 0.076	3005	³¹ ABREU	01 K	DLPH	Eee = 183+189 GeV
2.152 ± 0.066		79176	³² ABBOTT	00 B	D0	Extracted value
2.05 ± 0.10	± 0.08	662	³³ AFFOLDER	00M	CDF	Direct meas.
2.24 ± 0.20	± 0.13	1711	³⁴ BARATE	00T	ALEP	$E_{\rm cm}^{\it ee} = 189 \; {\rm GeV}$
1.97 ± 0.34	± 0.17	687	³⁵ ACCIARRI	99	L3	E ^{ee} _{cm} = 172+183 GeV
2.064 ± 0.060	± 0.059		³⁶ ABE	95W	CDF	Extracted value
$2.10 \begin{array}{l} +0.14 \\ -0.13 \end{array}$	± 0.09	3559	³⁷ ALITTI	92	UA2	Extracted value
$2.18 \begin{array}{l} +0.26 \\ -0.24 \end{array}$			³⁸ ALBAJAR		UA1	Extracted value
• • • We do no	ot use the fo	llowing da	ata for averages, fit	s, lim	its, etc.	• • •
1.84 ± 0.32	± 0.20	674	³⁹ ABBIENDI	99 C	OPAL	Repl. by ABBI- ENDI 01F
2.044 ± 0.097	,	11858	⁴⁰ ABBOTT	99н	D0	Repl. by AB- BOTT 00B
2.48 ± 0.40	± 0.10	737	⁴¹ ABREU	99T	DLPH	Repl. by ABREU 01K
$2.126 ^{+0.052}_{-0.048}$			⁴² BARATE	991	ALEP	E ^{ee} _{cm} = 161+172+183 GeV
$1.74 \begin{array}{l} +0.88 \\ -0.78 \end{array}$	± 0.25	101	⁴³ ACCIARRI	97 S	L3	Repl. by ACCIA- RRI 99
2.11 ± 0.28	± 0.16	58	⁴⁴ ABE	95 C	CDF	Repl. by AF- FOLDER 00M
2.30 ± 0.19	± 0.06		⁴⁵ ALITTI	90 C	UA2	Extracted value
$2.8 \begin{array}{c} +1.4 \\ -1.5 \end{array}$	± 1.3	149	⁴⁶ ALBAJAR	89	UA1	$E_{\rm cm}^{p\overline{p}} = 546,630 \text{ GeV}$
<7	90	119	APPEL	86	UA2	$E_{\rm cm}^{p\overline{p}} = 546,630 \; {\rm GeV}$
<6.5	90	86	⁴⁷ ARNISON	86	UA1	$E_{cm}^{p\overline{p}} = 546,630 \; GeV$

²⁹ ABAZOV 02E obtain this result fitting the high-end tail (90–200 GeV) of the transversemass spectrum in semileptonic $W \rightarrow e \nu_e$ decays.

 $^{^{30}}$ ABBIENDI 01F obtain this value from a fit to the reconstructed W mass distribution using data at 172, 183, and 189 GeV. The systematic error includes ± 0.010 GeV due to LEP energy uncertainty and ± 0.078 GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.

³¹ ABREU 01K obtain this value properly combining results obtained at 183 and 189 GeV using $WW \to \ell \overline{\nu}_\ell q \overline{q}$ and $WW \to q \overline{q} q \overline{q}$ decays. The systematic error includes an

- uncertainty of ± 0.052 GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.
- 32 ABBOTT 00B measure $R=10.43\pm0.27$ for the $W\to e\nu_e$ decay channel. They use the SM theoretical predictions for $\sigma(W)/\sigma(Z)$ and $\Gamma(W\to e\nu_e)$ and the world average for B($Z\to ee$). The value quoted here is obtained combining this result (2.169 \pm 0.070 GeV) with that of ABBOTT 99H.
- 33 AFFOLDER 00M fit the high transverse mass (100–200 GeV) $W\to e\nu_e$ and $W\to \mu\nu_\mu$ events to obtain $\Gamma(W)=2.04\pm0.11({\rm stat})\pm0.09({\rm syst})$ GeV. This is combined with the earlier CDF measurement (ABE 95C) to obtain the quoted result.
- ³⁴ BARATE 00T obtain this value using $WW\to q\overline{q}q\overline{q}$, $WW\to e\nu_e q\overline{q}$, and $WW\to \mu\nu_\mu q\overline{q}$ decays. The systematic error includes ± 0.015 GeV due to LEP energy uncertainty and ± 0.080 GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.
- 35 ACCIARRI 99 obtain this value from a fit to the reconstructed W mass distribution using data at 172 and 183 GeV.
- ³⁶ ABE 95W measured $R=10.90\pm0.32\pm0.29$. They use $m_{W}=80.23\pm0.18$ GeV, $\sigma(W)/\sigma(Z)=3.35\pm0.03$, $\Gamma(W\to e\nu)=225.9\pm0.9$ MeV, $\Gamma(Z\to e^+e^-)=83.98\pm0.18$ MeV, and $\Gamma(Z)=2.4969\pm0.0038$ GeV.
- 37 ALITTI 92 measured $R=10.4^{+0.7}_{-0.6}\pm0.3$. The values of $\sigma(Z)$ and $\sigma(W)$ come from $O(\alpha_s^2)$ calculations using $m_W=80.14\pm0.27$ GeV, and $m_Z=91.175\pm0.021$ GeV along with the corresponding value of $\sin^2\!\theta_W=0.2274$. They use $\sigma(W)/\sigma(Z)=3.26\pm0.07\pm0.05$ and $\Gamma(Z)=2.487\pm0.010$ GeV.
- 38 ALBAJAR 91 measured $R=9.5^{+1.1}_{-1.0}$ (stat. + syst.). $\sigma(W)/\sigma(Z)$ is calculated in QCD at the parton level using $m_W=80.18\pm0.28$ GeV and $m_Z=91.172\pm0.031$ GeV along with $\sin^2\!\theta_W=0.2322\pm0.0014$. They use $\sigma(W)/\sigma(Z)=3.23\pm0.05$ and $\Gamma(Z)=2.498\pm0.020$ GeV. This measurement is obtained combining both the electron and muon channels.
- ABBIENDI 99C obtain this value from a fit to the reconstructed W mass distribution using data at 172 and 183 GeV. The systematic error includes an uncertainty of ± 0.12 GeV due to the possible color-reconnection and Bose-Einstein effects in the purely hadronic final states and an uncertainty of ± 0.01 GeV due to the beam energy.
- ⁴⁰ ABBOTT 99H measure R= 10.90 \pm 0.52 combining electron and muon channels. They use $M_W = 80.39 \pm 0.06$ GeV and the SM theoretical predictions for $\sigma(W)/\sigma(Z)$, B($Z \rightarrow \ell \ell$), and $\Gamma(W \rightarrow \ell \nu_{\ell})$.
- ⁴¹ ABREU 99T obtain this value using $WW \to \ell \overline{\nu}_{\ell} q \overline{q}$ and $WW \to q \overline{q} q \overline{q}$ events. The systematic error includes an uncertainty of ± 0.080 GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.
- ⁴²BARATE 99I obtain this result with a fit to the *WW* measured cross sections at 161, 172, and 183 GeV. The theoretical prediction takes into account the sensitivity to the *W* total width.
- 43 ACCIARRI 97S obtain this value from a fit to the reconstructed W mass distribution.
- ⁴⁴ ABE 95C use the tail of the transverse mass distribution of $W
 ightarrow e
 u_e$ decays.
- ⁴⁵ ALITTI 90C used the same technique as described for ABE 90. They measured $R=9.38^{+0.82}_{-0.72}\pm0.25$, obtained $\Gamma(W)/\Gamma(Z)=0.902\pm0.074\pm0.024$. Using $\Gamma(Z)=2.546\pm0.032$ GeV, they obtained the $\Gamma(W)$ value quoted above and the limits $\Gamma(W)<<2.56$ (2.64) GeV at the 90% (95%) CL. $E_{\rm CM}^{p\overline{p}}=546,630$ GeV.
- ⁴⁶ ALBAJAR 89 result is from a total sample of 299 $W \rightarrow e \nu$ events.
- 47 If systematic error is neglected, result is $2.7^{+1.4}_{-1.5}$ GeV. This is enhanced subsample of 172 total events.

W⁺ DECAY MODES

 W^- modes are charge conjugates of the modes below.

	Mode	Fraction (Γ_i/Γ)	Confidence level
$\overline{\Gamma_1}$	$\ell^+ \nu$	[a] (10.80± 0.09) %	
_	$e^+ u$	$(10.75 \pm \ 0.13) \%$	
Γ ₃	$\mu^+ \nu$	$(10.57 \pm \ 0.15) \ \%$	
Γ_4	$ au^+ u$	$(11.25 \pm \ 0.20) \%$	
Γ_5	hadrons	$(67.60 \pm \ 0.27) \%$	
Γ_6	$\pi^+\gamma$	< 8 ×	10^{-5} 95%
Γ_7	$D_s^+ \gamma$	< 1.3 ×	10^{-3} 95%
Γ ₈	cX	(33.4 \pm 2.6) %	
Γ ₉	c s	$(31 {}^{+13}_{-11}) \%$	
Γ_{10}	invisible	[b] (1.4 \pm 2.8) %	

- [a] ℓ indicates each type of lepton $(e, \mu, \text{ and } \tau)$, not sum over them.
- [b] This represents the width for the decay of the W boson into a charged particle with momentum below detectability, p< 200 MeV.

W PARTIAL WIDTHS

Γ(invisible)

Γ₁₀

This represents the width for the decay of the W boson into a charged particle with momentum below detectability, p< 200 MeV.

VALUE (MeV)	DOCUMENT ID		TECN	COMMENT
$30^{+52}_{-48}\pm33$	⁴⁸ BARATE	991	ALEP	$E_{\rm cm}^{\rm ee} = 161 + 172 + 183$

• • We do not use the following data for averages, fits, limits, etc.

⁴⁹ BARATE 99L ALEP
$$E_{\text{cm}}^{\text{ee}} = 161 + 172 + 183$$

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 48 BARATE 991 measure this quantity using the dependence of the total cross section σ_{WW} upon a change in the total width. The fit is performed to the WW measured cross sections at 161, 172, and 183 GeV. This partial width is < 139 MeV at 95%CL.

W BRANCHING RATIOS

Overall fits are performed to determine the branching ratios of the W. LEP averages on $W \to e \nu_e$, $W \to \mu \nu_\mu$, and $W \to \tau \nu_\tau$, and their correlations are first obtained by combining results from the four experiments taking properly into account the common systematics. The procedure is described in the note LEPEWWG/XSEC/2001-02, 30 March 2001, at http://lepewwg.web.cern.ch/LEPEWWG/lepww/4f/PDG01. The LEP average values so obtained, using published data, are given in the note LEPEWWG/XSEC/2005-01 accessible at http://lepewwg.web.cern.ch/

⁴⁹ BARATE 99L use W-pair production to search for effectively invisible W decays, tagging with the decay of the other W boson to Standard Model particles. The partial width for effectively invisible decay is < 27 MeV at 95%CL.

LEPEWWG/lepww/4f/PDG05/. These results, together with results from the $p\overline{p}$ colliders are then used in fits to obtain the world average W branching ratios. A first fit determines three individual leptonic branching ratios, $\mathsf{B}(W \to e \nu_e), \, \mathsf{B}(W \to \mu \nu_\mu), \, \mathsf{and} \, \, \mathsf{B}(W \to \tau \nu_\tau). \, \, \mathsf{This} \, \, \mathsf{fit} \, \, \mathsf{has} \, \, \mathsf{a} \, \, \chi^2 =$ 4.7 for 10 degrees of freedom. A second fit assumes lepton universality and determines the leptonic branching ratio $\mathsf{B}(W \to \ell \nu_{\ell})$ and the hadronic branching ratio is derived as $B(W \to hadrons) = 1-3 B(W \to \ell \nu)$. This fit has a χ^2 =11.3 for 12 degrees of freedom.

The LEP $W
ightarrow \ell
u$ data are obtained by the Collaborations using individual leptonic channels and are, therefore, not included in the overall fits to avoid double counting.

 $\Gamma(\ell^+\nu)/\Gamma_{\text{total}}$ Γ_1/Γ ℓ indicates average over e, μ , and τ modes, not sum over modes.

VALUE	EV/EC	DOCUMENT ID	TECN	COMMENT
VALUE	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	COMMENT
0.1080 ± 0.0009 OUR FIT	-			
$0.1085 \pm 0.0014 \pm 0.0008$	13600	ABDALLAH	04G DLPH	$E_{cm}^{ee} = 161209 \; GeV$
$0.1083 \pm 0.0014 \pm 0.0010$	11246	ACHARD	04J L3	$E_{cm}^{ee} = 161209 \; GeV$
$0.1096 \pm 0.0012 \pm 0.0005$	16116	SCHAEL	04A ALEP	$E_{\rm cm}^{\it ee} = 183 – 209 \; {\rm GeV}$
$0.1056 \pm 0.0020 \pm 0.0009$	5778	ABBIENDI,G	00 OPAL	$E_{\rm cm}^{\rm ee} = 161 + 172 + 183$
				<u>+</u> 189 GeV
$0.1102\!\pm\!0.0052$	11858	⁵⁰ ABBOTT	99H D0	$^{+189}$ GeV $E_{cm}^{oldsymbol{p}\overline{oldsymbol{p}}}$ = 1.8 TeV
$0.104\ \pm0.008$	3642	⁵¹ ABE	92ı CDF	$E_{cm}^{ar{p}} = 1.8 \; TeV$
ullet $ullet$ We do not use the	following	data for averages,	fits, limits, et	tc. • • •
$0.1071 \pm 0.0024 \pm 0.0014$	4843	ABREU	00k DLPH	Repl. by ABDAL- LAH 04G
$0.1060 \pm 0.0023 \pm 0.0011$	5328	ACCIARRI	00V L3	Repl. by ACHARD 04J
$0.1101 \pm 0.0022 \pm 0.0011$	5258	BARATE	00J ALEP	Repl. by SCHAEL 04A
$0.107\ \pm0.004\ \pm0.002$	1440	ABBIENDI	99D OPAL	Repl. by ABBI- ENDI,G 00
50 ARROTT OOL massu	ro R — 1	$[\sigma \dots R(W) \times \ell_{W}]$)]/[a = B(7	$\rightarrow \ell\ell) = 10.90 \pm 0.52$

⁵⁰ ABBOTT 99H measure $R \equiv [\sigma_W \ B(W \to \ell \nu_\ell)]/[\sigma_Z \ B(Z \to \ell \ell)] = 10.90 \pm 0.52$ combining electron and muon channels. They use $M_W = 80.39 \pm 0.06$ GeV and the SM theoretical predictions for $\sigma(W)/\sigma(Z)$ and $B(Z \to \ell \ell)$.

⁵¹ $1216 \pm 38^{+27}_{-31} \ W \to \mu \nu$ events from ABE 92I and 2426 $W \to e \nu$ events of ABE 91C.

ABE 921 give the inverse quantity as 9.6 \pm 0.7 and we have inverted.

$\Gamma(e^+\nu)/\Gamma_{ m total}$				Γ_2/Γ
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.1075 ± 0.0013 OUR FIT				
0.1061 ± 0.0028		⁵² ABAZOV	04D TEVA	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
$0.1055 \pm 0.0031 \pm 0.0014$	1804	ABDALLAH	04G DLPH	$E_{cm}^{ee} = 161209 \; GeV$
$0.1078 \pm 0.0029 \pm 0.0013$	1576	ACHARD	04J L3	$E_{cm}^{ee} = 161209 \; GeV$
$0.1078 \pm 0.0027 \pm 0.0010$	2142	SCHAEL	04A ALEP	E ^{ee} _{cm} = 183–209 GeV
$0.1046 \pm 0.0042 \pm 0.0014$	801	ABBIENDI,G	00 OPAL	Eee = 161+172+183 +189 GeV
$0.10 \pm 0.014 \ ^{+ 0.02}_{- 0.03}$	248	⁵³ ANSARI	87C UA2	$E_{\rm cm}^{p \overline{p}} = 546,630 \; {\rm GeV}$

ullet ullet We do not use the following data for averages, fits, limits, etc. ullet ullet

$0.1044 \pm 0.0015 \pm 0.0028$	67318	⁵⁴ ABBOTT	00B D0	Repl. by
				ABAZOV 04D
$0.1018 \pm 0.0054 \pm 0.0026$	527	ABREU	00k DLPH	Repl. by ABDAL-
				LAH 04G
$0.1077 \pm 0.0045 \pm 0.0016$	715	ACCIARRI	00∨ L3	Repl. by ACHARD 04J
$0.1135 \pm 0.0046 \pm 0.0017$	720	BARATE	00J ALEP	Repl. by SCHAEL 04A
$0.117\ \pm0.009\ \pm0.002$	224	ABBIENDI	99D OPAL	Repl. by ABBI-
				ENDI,G 00
$0.1094 \pm 0.0033 \pm 0.0031$		⁵⁵ ABE	95w CDF	Repl. by
				ABAZOV 04D
seen	119	APPEL	86 UA2	$E_{\rm cm}^{p\overline{p}} = 546,630 \; {\rm GeV}$
seen	172	ARNISON	86 UA1	$E_{\rm cm}^{p} = 546,630 {\rm GeV}$

- 52 ABAZOV 04D take into account all correlations to properly combine the CDF (ABE 95W) and DØ (ABBOTT 00B) measurements of the ratio R in the electron channel. The ratio R is defined as $[\sigma_W \cdot \mathrm{B}(W \to e \nu_e)] \ / \ [\sigma_Z \cdot \mathrm{B}(Z \to e e)]$. The combination gives $\mathrm{R}^{Tevatron} = 10.59 \pm 0.23. \ \sigma_W \ / \ \sigma_Z$ is calculated at next–to–next–to–leading order (3.360 \pm 0.051). The branching fraction $\mathrm{B}(Z \to e e)$ is taken from this Review as (3.363 \pm 0.004)%.
- 53 The first error was obtained by adding the statistical and systematic experimental uncertainties in quadrature. The second error reflects the dependence on theoretical prediction of total W cross section: $\sigma(546~{\rm GeV})=4.7^{+1.4}_{-0.7}$ nb and $\sigma(630~{\rm GeV})=5.8^{+1.8}_{-1.0}$ nb. See ALTARELLI 85B.
- 54 ABBOTT 00B measure $R \equiv [\sigma_W B(W \to e \nu_e)]/[\sigma_Z B(Z \to e e)] = 10.43 \pm 0.27$ for the $W \to e \nu_e$ decay channel. They use the SM theoretical prediction for $\sigma(W)/\sigma(Z)$ and the world average for $B(Z \to e e)$.
- ⁵⁵ ABE 95W result is from a measurement of $\sigma B(W \to e \nu)/\sigma B(Z \to e^+ e^-) = 10.90 \pm 0.32 \pm 0.29$, the theoretical prediction for the cross section ratio, the experimental knowledge of $\Gamma(Z \to e^+ e^-) = 83.98 \pm 0.18$ MeV, and $\Gamma(Z) = 2.4969 \pm 0.0038$ GeV.

$\Gamma(\mu^+ u)/\Gamma_{ m total}$				Г ₃ /Г
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.1057±0.0015 OUR FIT				
$0.1065 \pm 0.0026 \pm 0.0008$	1998	ABDALLAH	04G DLPH	$E_{cm}^{ee} = 161209 \; GeV$
$0.1003 \pm 0.0029 \pm 0.0012$	1423	ACHARD	04J L3	$E_{cm}^{ee} = 161209 \; GeV$
$0.1087 \pm 0.0025 \pm 0.0008$	2216	SCHAEL	04A ALEP	$E_{\rm cm}^{\it ee} = 183 – 209 \; {\rm GeV}$
$0.1050 \pm 0.0041 \pm 0.0012$	803	ABBIENDI,G	00 OPAL	$E_{\rm cm}^{\rm ee} = 161 + 172 + 183$
				<u>+</u> 189 GeV
0.10 ± 0.01	1216	⁵⁶ ABE	92ı CDF	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
• • • We do not use the	following	data for averages,	fits, limits, e	tc. • • •
$0.1092 \pm 0.0048 \pm 0.0012$	649	ABREU	00k DLPH	Repl. by ABDAL- LAH 04G
$0.0990 \pm 0.0046 \pm 0.0015$	617	ACCIARRI	00V L3	Repl. by ACHARD 04J
$0.1110 \pm 0.0044 \pm 0.0016$	710	BARATE	00J ALEP	Repl. by SCHAEL 04A

 $^{^{56}}$ ABE 92I quote the inverse quantity as 9.9 \pm 1.2 which we have inverted.

193

 $0.102 \pm 0.008 \pm 0.002$

ABBIENDI

99D OPAL

Repl. by ABBI-ENDI,G 00

$\Gamma(au^+ u)/\Gamma_{ ext{total}}$				Γ_4/Γ
VALUE	<u>EVTS</u>	DOCUMENT ID	TECN	COMMENT
0.1125 ± 0.0020 OUR FIT				
$0.1146 \pm 0.0039 \pm 0.0019$	2034	ABDALLAH	04G DLPH	$E_{cm}^{\mathit{ee}} = 161209 \; GeV$
$0.1189 \pm 0.0040 \pm 0.0020$	1375	ACHARD	04J L3	$E_{cm}^{\mathit{ee}} = 161209 \; GeV$
$0.1125 \pm 0.0032 \pm 0.0020$	2070	SCHAEL	04A ALEP	$E_{\rm cm}^{\it ee} = 183 – 209 \; {\rm GeV}$
$0.1075 \pm 0.0052 \pm 0.0021$	794	ABBIENDI,G	00 OPAL	$E_{\rm cm}^{\rm ee} = 161 + 172 + 183$
				+189GeV
• • • We do not use the	following da	ta for averages,	fits, limits, e	tc. • • •
$0.1105 \pm 0.0075 \pm 0.0032$	579	ABREU	00k DLPH	Repl. by ABDAL-
0.1104 0.0000 0.0000	F06	A C C LA D D L	00, 10	LAH 04G
$0.1124 \pm 0.0062 \pm 0.0022$	536	ACCIARRI	00v L3	Repl. by ACHARD 04J
$0.1051 \pm 0.0055 \pm 0.0022$	607	BARATE	00J ALEP	Repl. by SCHAEL 04A
$0.101 \pm 0.010 \pm 0.003$	183	ABBIENDI	99D OPAL	Repl. by ABBI-
				ENDI,G 00

 $\Gamma(\text{hadrons})/\Gamma_{\text{total}}$

 Γ_5/Γ

OUR FIT value is obtained by a fit to the lepton branching ratio data assuming lepton

VALUE	<u>EVTS</u>	DOCUMENT ID	TECN	COMMENT
0.6760 ± 0.0027 OUR FIT	•			
$0.6745 \pm 0.0041 \pm 0.0024$	13600	ABDALLAH	04G DLPH	$E_{cm}^{ee} = 161209 \; GeV$
$0.6750 \pm 0.0042 \pm 0.0030$	11246	ACHARD	04J L3	$E_{cm}^{ee} = 161209 \; GeV$
$0.6713 \pm 0.0037 \pm 0.0015$	16116	SCHAEL	04A ALEP	E ^{ee} _{cm} = 183–209 GeV
$0.6832 \!\pm\! 0.0061 \!\pm\! 0.0028$	5778	ABBIENDI,G	00 OPAL	$E_{cm}^{ee} = 161 + 172 + 183$
1 47 1				$+189~{\sf GeV}$

• • We do not use the following data for averages, fits, limits, etc.

$0.6789 \pm 0.0073 \pm 0.0043$	4843	ABREU	00k DLPH	Repl. by ABDAL-
				LAH 04G
$0.6820 \pm 0.0068 \pm 0.0033$	5328	ACCIARRI	00∨ L3	Repl. by ACHARD 04J
$0.6697 \pm 0.0065 \pm 0.0032$	5258	BARATE	00J ALEP	Repl. by SCHAEL 04A
$0.679 \pm 0.012 \pm 0.005$	1440	ABBIENDI	99D OPAL	Repl. by ABBI-
				ENDI,G 00

$\Gamma(\mu^+ u)/\Gamma(e^+ u)$				Γ_3/Γ_2
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.983 ± 0.018 OUR FIT				
$0.89\ \pm0.10$	13k	⁵⁷ ABACHI	95D D0	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
1.02 ± 0.08	1216	⁵⁸ ABE	92ı CDF	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
$1.00 \pm 0.14 \pm 0.08$	67	ALBAJAR	89 UA1	$E_{\rm cm}^{p\overline{p}}$ = 546,630 GeV

^{• • •} We do not use the following data for averages, fits, limits, etc. • • •

106				
$1.24 \begin{array}{c} +0.6 \\ -0.4 \end{array}$	14	ARNISON	84D UA1	Repl. by ALBAJAR 89

 $^{^{57}}$ ABACHI 95D obtain this result from the measured σ_W B($W
ightarrow ~\mu
u$)= 2.09 \pm 0.23 \pm 0.11 nb and $\sigma_W B(W \to e \nu) = 2.36 \pm 0.07 \pm 0.13$ nb in which the first error is the combined statistical and systematic uncertainty, the second reflects the uncertainty in the luminosity.

⁵⁸ ABE 92I obtain σ_W B($W \to \mu \nu$)= 2.21 \pm 0.07 \pm 0.21 and combine with ABE 91C σ_W B($W \to e \nu$)) to give a ratio of the couplings from which we derive this measurement.

$\Gamma(au^+ u)/\Gamma(e^+ u)$				Γ_4/Γ_2
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
1.046 ± 0.023 OUR FIT				
$0.961\!\pm\!0.061$	980	⁵⁹ ABBOTT		$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
0.94 ± 0.14	179	⁶⁰ ABE	92E CDF	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
$1.04 \pm 0.08 \pm 0.08$	754	⁶¹ ALITTI		$E_{ m cm}^{p\overline{p}}=$ 630 GeV
$1.02\ \pm0.20\ \pm0.12$	32	ALBAJAR	89 UA1	$E_{ m cm}^{p\overline{p}} =$ 546,630 GeV
\bullet \bullet We do not use the	following	data for averages	, fits, limits, e	tc. • • •
$0.995\!\pm\!0.112\!\pm\!0.083$	198	ALITTI	91c UA2	Repl. by ALITTI 92F
$1.02 \pm 0.20 \pm 0.10$	32	ALBAJAR	87 UA1	Repl. by ALBAJAR 89

 $^{^{59}}$ ABBOTT 00D measure $\sigma_W \times$ B($W \to ~\tau \, \nu_{\tau}) = 2.22 \pm 0.09 \pm 0.10 \pm 0.10$ nb. Using the ABBOTT 00B result $\sigma_W \times {\sf B}(W \to e \nu_e) = 2.31 \pm 0.01 \pm 0.05 \pm 0.10$ nb, they quote the ratio of the couplings from which we derive this measurement.

 $^{^{61}\,\}mathrm{This}$ measurement is derived by us from the ratio of the couplings of ALITTI 92F.

$\Gamma(\pi^+\gamma)/\Gamma(e^+ u)$				Γ_6/Γ_2
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 7 \times 10^{-4}$	95	ABE	98н CDF	$E_{cm}^{ar{p}} = 1.8 \; TeV$
$< 4.9 \times 10^{-3}$	95	⁶² ALITTI	92D UA2	$E_{ m cm}^{p \overline{p}} =$ 630 GeV
$< 58 \times 10^{-3}$	95	⁶³ ALBAJAR	90 UA1	$E_{ m cm}^{p\overline{p}}=$ 546, 630 GeV
62 ALITTI OZD limit i	c 3 8 v 10	1-3 at 00%CI		

ALITTI 92D limit is 3.8×10^{-3} at 90%CL.

 $^{^{63}}$ ALBAJAR 90 obtain < 0.048 at 90%CL.

$\Gamma(D_s^+\gamma)/\Gamma(e^+\nu)$					Γ_7/Γ_2		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT			
$<1.2 \times 10^{-2}$	95	ABE	98P CDF	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$			
$\Gamma(cX)/\Gamma(hadrons)$ Γ_8/Γ_8							
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT			
0.49 ±0.04 OUR AVERAGE							
$0.481\!\pm\!0.042\!\pm\!0.032$	3005	⁶⁴ ABBIENDI	00V OPAL	$E_{cm}^{ee} = 183 + 18$	9 GeV		
$0.51 \ \pm 0.05 \ \pm 0.03$	746	⁶⁵ BARATE	99м ALEP	$E_{cm}^{ee} = 172 + 18$	3 GeV		

 $^{^{64}}$ ABBIENDI 00V tag $W
ightarrow c\,\mathrm{X}$ decays using measured jet properties, lifetime information, and leptons produced in charm decays. From this result, and using the additional measurements of $\Gamma(W)$ and $\mathsf{B}(W \to \mathsf{hadrons}), |V_{cs}|$ is determined to be $0.969 \pm 0.045 \pm 0.036$.

 $^{^{60}}$ ABE 92E use two procedures for selecting $W
ightarrow ~ au
u_{ au}$ events. The missing E $_{ au}$ trigger leads to $132\pm14\pm8$ events and the au trigger to $47\pm9\pm4$ events. Proper statistical and systematic correlations are taken into account to arrive at $\sigma B(W \to \tau \nu) = 2.05 \pm 0.27$ nb. Combined with ABE 91C result on $\sigma B(W \rightarrow e \nu)$, ABE 92E quote a ratio of the couplings from which we derive this measurement.

 $^{^{65}}$ BARATE 99M tag c jets using a neural network algorithm. From this measurement $\left|V_{cs}\right|$ is determined to be 1.00 \pm 0.11 \pm 0.07.

AVERAGE PARTICLE MULTIPLICITIES IN HADRONIC W DECAY

Summed over particle and antiparticle, when appropriate.

 $\langle N_{\pi^{\pm}} \rangle$

 $\frac{DOCUMENT\ ID}{67}$ $\frac{TECN}{ABREU,P}$ $\frac{COMMENT}{DLPH}$ $\frac{E_{cm}^{ee}}{E_{cm}^{ee}}$ 189 GeV

 67 ABREU,P 00F measure $\left\langle N_{\pi^{\pm}} \right\rangle = 31.65 \pm 0.48 \pm 0.76$ and $15.51 \pm 0.38 \pm 0.40$ in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.

 $\langle N_{K^{\pm}} \rangle$

 15.70 ± 0.35

VALUEDOCUMENT IDTECNCOMMENT2.20 \pm 0.1968 ABREU,P00F DLPH $E_{cm}^{ee} = 189 \text{ GeV}$

 $\langle N_p \rangle$

VALUEDOCUMENT IDTECNCOMMENT 0.92 ± 0.14 69 ABREU,P00F DLPH $E_{cm}^{ee} = 189 \text{ GeV}$

$\langle N_{\rm charged} \rangle$

 VALUE
 DOCUMENT ID
 TECN
 COMMENT

 19.41 \pm 0.15 OUR AVERAGE
 70 ABREU,P
 00F DLPH
 $E_{cm}^{ee} = 183 + 189 \text{ GeV}$

 19.3 \pm 0.3 \pm 0.3
 71 ABBIENDI
 99N OPAL
 $E_{cm}^{ee} = 183 \text{ GeV}$

 19.23 \pm 0.74
 72 ABREU
 98C DLPH
 $E_{cm}^{ee} = 172 \text{ GeV}$

 $^{^{66}}$ ABREU 98N tag c and s jets by identifying a charged kaon as the highest momentum particle in a hadronic jet. They also use a lifetime tag to independently identify a c jet, based on the impact parameter distribution of charged particles in a jet. From this measurement $\left|V_{c\,s}\right|$ is determined to be $0.94^{\,+0.32}_{\,-0.26}\pm0.13$.

 $^{^{68}}$ ABREU,P 00F measure $\langle \textit{N}_{\textit{K}^{\pm}} \rangle = 4.38 \pm 0.42 \pm 0.12$ and $2.23 \pm 0.32 \pm 0.17$ in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.

 $^{^{69}}$ ABREU,P 00F measure $\langle N_p \rangle = 1.82 \pm 0.29 \pm 0.16$ and 0.94 \pm 0.23 \pm 0.06 in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.

 $^{^{70}\,\}text{ABREU,P}$ 00F measure $\left\langle \textit{N}_{\text{charged}}\right\rangle = 39.12 \pm 0.33 \pm 0.36$ and $38.11 \pm 0.57 \pm 0.44$ in the fully hadronic final states at 189 and 183 GeV respectively, and $\left\langle \textit{N}_{\text{charged}}\right\rangle = 19.49 \pm 0.31 \pm 0.27$ and $19.78 \pm 0.49 \pm 0.43$ in the semileptonic final states. The value quoted is a weighted average without assuming any correlations.

 $^{^{71}}$ ABBIENDI 99N use the final states $W^+W^- o q \overline{q} \ell \overline{
u}_\ell$ to derive this value.

 $^{^{72}}$ ABREU 98C combine results from both the fully hadronic as well semileptonic WW final states after demonstrating that the W decay charged multiplicity is independent of the topology within errors.

TRIPLE GAUGE COUPLINGS (TGC'S)

A REVIEW GOES HERE - Check our WWW List of Reviews

Δg_1^Z

OUR FIT below is obtained by combining the measurements taking into account properly the common systematic errors (see LEPEWWG/TGC/2005-01 at http://lepewwg.web.cern.ch/LEPEWWG/lepww/tgc).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT				
$-0.016^{+0.022}_{-0.019}$ OUR FIT								
$0.001\!\pm\!0.027\!\pm\!0.013$	9310	⁷³ SCHAEL	05A ALEP	E ^{ee} _{cm} = 183–209 GeV				
$-0.013 ^{+ 0.034}_{- 0.033}$	9800	⁷⁴ ABBIENDI	04D OPAL	E ^{ee} _{cm} = 183–209 GeV				
$-0.034^{+0.034}_{-0.032}\!\pm\!0.015$	8325	⁷⁵ ACHARD	04D L3	$E_{cm}^{ee} = 161209 \; GeV$				
• • • We do not use the following data for averages, fits, limits, etc. • •								
$-0.02\ \pm0.07\ \pm0.01$	2114	⁷⁶ ABREU	01ı DLPH	$E_{cm}^{ee} = 183 + 189 \; GeV$				
	331	⁷⁷ ABBOTT	99ı D0	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$				

 73 SCHAEL 05A study single–photon, single–W, and WW–pair production from 183 to 209 GeV. The result quoted here is derived from the WW–pair production sample. Each parameter is determined from a single–parameter fit in which the other parameters assume their Standard Model values.

ABBIENDI 04D combine results from W^+W^- in all decay channels. Only *CP*-conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is $-0.077 < \Delta g_1^Z < 0.054$.

 $^{75}\,\mathrm{ACHARD}$ 04D study WW—pair production, single—W production and single—photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained from the WW—pair production sample including data from 161 to 183 GeV, ACCIARRI 99Q. Each parameter is determined from a single—parameter fit in which the other parameters assume their Standard Model values.

 76 ABREU 011 combine results from $e^+\,e^-$ interactions at 189 GeV leading to $W^+\,W^-$ and $W\,e\,\nu_e$ final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is $-0.16 < \Delta g_1^Z < 0.13$.

 77 ABBOTT 99I perform a simultaneous fit to the $W\gamma,~WW\to~$ dilepton, $WW/WZ\to e\nu jj,~WW/WZ\to \mu\nu jj,~$ and $WZ\to~$ trilepton data samples. For $\Lambda=2.0$ TeV, the 95%CL limits are $-0.37<\Delta g_1^Z<0.57,~$ fixing $\lambda_Z{=}\Delta\kappa_Z=0$ and assuming Standard Model values for the $WW\gamma$ couplings.

$\Delta \kappa_{\gamma}$

OUR FIT below is obtained by combining the measurements taking into account properly the common systematic errors (see LEPEWWG/TGC/2005-01 at http://lepewwg.web.cern.ch/LEPEWWG/lepww/tgc).

<u>V</u> ALUE	<u>EVTS</u>	DOCUMENT ID	TECN	COMMENT
$-0.027^{+0.044}_{-0.045}$ OUR FI	Т			
$-0.029\!\pm\!0.055\!\pm\!0.030$	10689	⁷⁸ SCHAEL	05A ALEP	Eee = 183–209 GeV
$-0.12 \begin{array}{l} +0.09 \\ -0.08 \end{array}$	9800	⁷⁹ ABBIENDI	04D OPAL	E ^{ee} _{cm} = 183–209 GeV
$0.013^{\displaystyle +0.067}_{\displaystyle -0.064}\!\pm\!0.026$	10575	⁸⁰ ACHARD	04D L3	$E_{cm}^{ee} = 161209 \; GeV$

HTTP://PDG.LBL.GOV

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• • • We do not use the following data for averages, fits, limits, etc. • •

- 78 SCHAEL 05A study single–photon, single–W, and WW–pair production from 183 to 209 GeV. Each parameter is determined from a single–parameter fit in which the other parameters assume their Standard Model values.
- 79 ABBIENDI 04D combine results from $W^+\,W^-$ in all decay channels. Only $\it CP$ -conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is $-0.27 < \Delta \kappa_{\gamma} < 0.07$.
- 80 ACHARD 04D study WW-pair production, single-W production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained including data from 161 to 183 GeV, ACCIARRI 99Q. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 81 ABREU 011 combine results from e^+e^- interactions at 189 GeV leading to W^+W^- , $We\nu_e$, and $\nu\overline{\nu}\gamma$ final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is $-0.13 < \Delta\kappa_\gamma < 0.68$.
- ⁸² BREITWEG 00 search for W production in events with large hadronic p_T . For $p_T >$ 20 GeV, the upper limit on the cross section gives the 95%CL limit $-4.7 < \Delta \kappa_{\gamma} < 1.5$ (for $\lambda_{\gamma} =$ 0).
- ⁸³ ABBOTT 99I perform a simultaneous fit to the $W\gamma$, $WW\to dilepton$, $WW/WZ\to e\nu jj$, $WW/WZ\to \mu\nu jj$, and $WZ\to trilepton$ data samples. For $\Lambda=2.0$ TeV, the 95%CL limits are $-0.25<\Delta\kappa_{\gamma}<0.39$.

 λ_{γ}

OUR FIT below is obtained by combining the measurements taking into account properly the common systematic errors (see LEPEWWG/TGC/2005-01 at http://lepewwg.web.cern.ch/LEPEWWG/lepww/tgc).

VALUE	<u>EVTS</u>	DOCUMENT ID	TECN	COMMENT			
$-0.028^{+0.020}_{-0.021}$ OUR FIT							
$-0.012\!\pm\!0.027\!\pm\!0.011$	10689	⁸⁴ SCHAEL	05A ALEP	E ^{ee} _{cm} = 183–209 GeV			
$-0.060 {}^{+ 0.034}_{- 0.033}$	9800	⁸⁵ ABBIENDI	04D OPAL	E ^{ee} _{cm} = 183–209 GeV			
$-0.021^{+0.035}_{-0.034}{\pm}0.017$	10575	⁸⁶ ACHARD	04D L3	$E_{cm}^{\mathit{ee}} = 161209 \; GeV$			
 • • • We do not use the following data for averages, fits, limits, etc. 							

0.05
$$\pm$$
0.09 \pm 0.01 2298 87 ABREU 011 DLPH $E_{\rm cm}^{\rm ee}=183+189~{\rm GeV}$ 88 BREITWEG 00 ZEUS $e^+p_-\to e^+W^\pm X$, $\sqrt{s}\approx 300~{\rm GeV}$ 0.00 $^{+0.10}_{-0.09}$ 331 89 ABBOTT 991 D0 $E_{\rm cm}^{p\overline{p}}=1.8~{\rm TeV}$

 84 SCHAEL 05A study single–photon, single–W, and WW–pair production from 183 to 209 GeV. Each parameter is determined from a single–parameter fit in which the other

parameters assume their Standard Model values.

ABBIENDI 04D combine results from W^+W^- in all decay channels. Only *CP*-conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is $-0.13 < \lambda_{\gamma} < 0.01$.

- 86 ACHARD 04D study WW-pair production, single-W production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained including data from 161 to 183 GeV, ACCIARRI 99Q. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- ⁸⁷ ABREU 011 combine results from e^+e^- interactions at 189 GeV leading to W^+W^- , $We\nu_e$, and $\nu\overline{\nu}\gamma$ final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is $-0.11 < \lambda_\gamma < 0.23$.
- ⁸⁸ BREITWEG 00 search for W production in events with large hadronic p_T . For $p_T > 20$ GeV, the upper limit on the cross section gives the 95%CL limit $-3.2 < \lambda_{\gamma} < 3.2$ (for $\Delta \kappa_{\gamma} = 0$).
- ⁸⁹ ABBOTT 99I perform a simultaneous fit to the $W\gamma$, $WW\to dilepton$, $WW/WZ\to e\nu jj$, $WW/WZ\to \mu\nu jj$, and $WZ\to trilepton$ data samples. For $\Lambda=2.0$ TeV, the 95%CL limits are $-0.18<\lambda_{\gamma}<0.19$.

$\Delta \kappa_Z$

This coupling is *CP*-conserving (*C*- and *P*- separately conserving).

 VALUE
 EVTS
 DOCUMENT ID
 TECN
 COMMENT

 $-0.076^{+0.059}_{-0.056} \pm 0.024$ 7171
 90 ACHARD
 04D L3
 $E^{ee}_{cm} = 189-209$ GeV

 $^{90}\,\mathrm{ACHARD}$ 04D study WW—pair production, single—W production and single—photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the WW—pair production sample. Each parameter is determined from a single—parameter fit in which the other parameters assume their Standard Model values.

λ_{Z}

This coupling is CP-conserving (C- and P- separately conserving).

 VALUE
 EVTS
 DOCUMENT ID
 TECN
 COMMENT

 $-0.088^{+0.060}_{-0.057} \pm 0.023$ 7171
 91 ACHARD
 04D L3
 $E^{ee}_{cm} = 189-209$ GeV

Δg_5^Z

This coupling is *CP*-conserving but *C*- and *P*-violating.

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VALUE	<u>EVTS</u>	DOCUMENT ID	TECN	COMMENT		
-0.07 ± 0.09 OUR AV	ERAGE	Error includes scale	factor of 1.1.			
$-0.04^{igoplus 0.13}_{igoplus 0.12}$	9800	⁹² ABBIENDI	04D OPAL	E ^{ee} _{cm} = 183–209 GeV		
$0.00\!\pm\!0.13\!\pm\!0.05$	7171	⁹³ ACHARD	04D L3	$E_{cm}^{ee} = 189209 \; GeV$		
$-0.44^{\color{red}+0.23}_{-0.22}\!\pm\!0.12$	1154	⁹⁴ ACCIARRI	99Q L3	Eee = 161+172+ 183		

• • • We do not use the following data for averages, fits, limits, etc. • • •

 -0.16 ± 0.23 95 EBOLI 00 THEO LEP1, SLC+ Tevatron

- 92 ABBIENDI 04D combine results from $W^+\,W^-$ in all decay channels. Only *CP*-conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is $-0.28 < \Delta g_5^Z < 0.21$.
- 93 ACHARD 04D study WW-pair production, single–W production and single–photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained

 $^{^{91}}$ ACHARD 04D study WW-pair production, single–W production and single–photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the WW-pair production sample. Each parameter is determined from a single–parameter fit in which the other parameters assume their Standard Model values.

using the WW-pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.

g_{Δ}^{Z}

This coupling is *CP*-violating (*C*-violating and *P*-conserving).

VALUE	<u>EVTS</u>	DOCUMENT ID	_	<u>COMMENT</u>	
$-0.02_{-0.33}^{+0.32}$	1065	⁹⁶ ABBIENDI	01н OPAL	<i>E</i> ^{ee} _{cm} = 189 GeV	

 $^{^{96}}$ ABBIENDI 01H study W-pair events, with one leptonically and one hadronically decaying W. The coupling is extracted using information from the W production angle together with decay angles from the leptonically decaying W.

$\tilde{\kappa}_{Z}$

This coupling is *CP*-violating (*C*-conserving and *P*-violating).

VALUE	EVTS	DOCUMENT ID	_	COMMENT	_
$-0.20^{+0.10}_{-0.07}$	1065	⁹⁷ ABBIENDI	01н OPAL	<i>E</i> ^{ee} _{cm} = 189 GeV	

 $^{^{97}}$ ABBIENDI 01H study W-pair events, with one leptonically and one hadronically decaying W. The coupling is extracted using information from the W production angle together with decay angles from the leptonically decaying W.

$\widetilde{\lambda}_{\pmb{Z}}$

This coupling is *CP*-violating (*C*-conserving and *P*-violating).

VALUE	<u>EVTS</u>	DOCUMENT ID	•	COMMENT	
$-0.18^{+0.24}_{-0.16}$	1065	⁹⁸ ABBIENDI	01н OPAL	<i>E</i> ee = 189 GeV	

 $^{^{98}}$ ABBIENDI 01H study W-pair events, with one leptonically and one hadronically decaying W. The coupling is extracted using information from the W production angle together with decay angles from the leptonically decaying W.

W ANOMALOUS MAGNETIC MOMENT

The full magnetic moment is given by $\mu_W=e(1+\kappa+\lambda)/2m_W$. In the Standard Model, at tree level, $\kappa=1$ and $\lambda=0$. Some papers have defined $\Delta\kappa=1-\kappa$ and assume that $\lambda=0$. Note that the electric quadrupole moment is given by $-e(\kappa-\lambda)/m_W^2$. A description of the parameterization of these moments and additional references can be found in HAGIWARA 87 and BAUR 88. The parameter Λ appearing in the theoretical limits below is a regularization cutoff which roughly corresponds to the energy scale where the structure of the W boson becomes manifest.

VALUE (e/2m _W)	EVTS	DOCUMENT ID		TECN	COMMENT
$2.22^{+0.20}_{-0.19}$	2298	⁹⁹ ABREU	011	DLPH	E ^{ee} _{cm} = 183+189 GeV

 $^{^{94}}$ ACCIARRI 99Q study $W\text{-pair}, \, \text{single-}W, \, \text{and single photon events.}$

⁹⁵ EBOLI 00 extract this indirect value of the coupling studying the non-universal one-loop contributions to the experimental value of the $Z \rightarrow b \bar{b}$ width (Λ =1 TeV is assumed).

• • • We do not use the following data for averages, fits, limits, etc. • • •

100	ABE	95G	CDF
101	ALITTI		UA2
102	SAMUEL	92	THEO
103	SAMUEL	91	THEO
104	GRIFOLS	88	THEO
105	GROTCH	87	THEO
106	VANDERBIJ	87	THEO
107	GRAU	85	THEO
108	SUZUKI	85	THEO
109	HERZOG	84	THEO

- ABREU 011 combine results from e^+e^- interactions at 189 GeV leading to W^+W^- , $We\nu_e$, and $\nu\overline{\nu}\gamma$ final states with results from ABREU 99L at 183 GeV to determine Δg_1^Z , $\Delta\kappa_\gamma$, and λ_γ . $\Delta\kappa_\gamma$ and λ_γ are simultaneously floated in the fit to determine μ_W .
- $^{\mu}W$. 100 ABE 95G report $-1.3<\kappa<3.2$ for $\lambda=$ 0 and $-0.7<\lambda<0.7$ for $\kappa=$ 1 in $p\overline{p}\rightarrow~e\nu_{e}\gamma$ X and $\mu\nu_{tt}\gamma$ X at $\sqrt{s}=1.8$ TeV.
- 101 ALITTI 92C measure $\kappa=1^{+2.6}_{-2.2}$ and $\lambda=0^{+1.7}_{-1.8}$ in $p\overline{p}\to e\nu\gamma+$ X at $\sqrt{s}=630$ GeV. At 95%CL they report $-3.5<\kappa<5.9$ and $-3.6<\lambda<3.5$.
- $^{102}\,\text{SAMUEL}$ 92 use preliminary CDF and UA2 data and find $-2.4 < \kappa < 3.7$ at 96%CL and $-3.1 < \kappa < 4.2$ at 95%CL respectively. They use data for $W\,\gamma$ production and radiative W decay.
- ¹⁰³ SAMUEL 91 use preliminary CDF data for $p\overline{p} \to W\gamma X$ to obtain $-11.3 \le \Delta \kappa \le 10.9$. Note that their $\kappa = 1 \Delta \kappa$.
- 104 GRIFOLS 88 uses deviation from ho parameter to set limit $\Delta\kappa \lesssim 65~(M_W^2/\Lambda^2)$.
- 105 GROTCH 87 finds the limit $-37 < \Delta \kappa < 73.5$ (90% CL) from the experimental limits on $e^+e^- \to \nu \overline{\nu} \gamma$ assuming three neutrino generations and $-19.5 < \Delta \kappa < 56$ for four generations. Note their $\Delta \kappa$ has the opposite sign as our definition.
- 106 VANDERBIJ 87 uses existing limits to the photon structure to obtain $|\Delta\kappa|<33$ (m_W/Λ) . In addition VANDERBIJ 87 discusses problems with using the ρ parameter of the Standard Model to determine $\Delta\kappa$.
- GRAU 85 uses the muon anomaly to derive a coupled limit on the anomalous magnetic dipole and electric quadrupole (λ) moments $1.05 > \Delta \kappa \ln(\Lambda/m_W) + \lambda/2 > -2.77$. In the Standard Model $\lambda = 0$.
- SUZUKI 85 uses partial-wave unitarity at high energies to obtain $|\Delta\kappa|\lesssim 190~(m_W/\Lambda)^2$. From the anomalous magnetic moment of the muon, SUZUKI 85 obtains $|\Delta\kappa|\lesssim 2.2/\ln(\Lambda/m_W)$. Finally SUZUKI 85 uses deviations from the ρ parameter and obtains a very qualitative, order-of-magnitude limit $|\Delta\kappa|\lesssim 150~(m_W/\Lambda)^4$ if $|\Delta\kappa|\ll 1$.
- ¹⁰⁹ HERZOG 84 consider the contribution of W-boson to muon magnetic moment including anomalous coupling of $WW\gamma$. Obtain a limit $-1 < \Delta\kappa < 3$ for $\Lambda \gtrsim 1$ TeV.

ANOMALOUS W/Z QUARTIC COUPLINGS

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a_0/Λ^2 , a_c/Λ^2 , a_n/Λ^2

Using the $WW\gamma$ final state, the LEP combined 95% CL limits on the anomalous contributions to the $WW\gamma\gamma$ and $WWZ\gamma$ vertices (as of summer 2003) are given below:

(See P. Wells, "Experimental Tests of the Standard Model," Int. Europhysics Conference on High-Energy Physics, Aachen, Germany, 17–23 July 2003)

VALUE

DOCUMENT ID TECN

● We do not use the following data for averages, fits, limits, etc.

110 ABBIENDI 04B OPAL 111 ABBIENDI 04L OPAL 112 HEISTER 04A ALEP 113 ABDALLAH 03I DLPH 114 ACHARD 02F L3

- ABBIENDI 04B select 187 $e^+e^- o W^+W^-\gamma$ events in the C.M. energy range 180–209 GeV, where $E_{\gamma} >$ 2.5 GeV, the photon has a polar angle $|\cos\theta_{\gamma}| <$ 0.975 and is well isolated from the nearest jet and charged lepton, and the effective masses of both fermion-antifermion systems agree with the W mass within 3 Γ_W . The measured differential cross section as a function of the photon energy and photon polar angle is used to extract the 95% CL limits: $-0.020~{\rm GeV}^{-2} < a_0/\Lambda^2 < 0.020~{\rm GeV}^{-2}, -0.053~{\rm GeV}^{-2} < a_c/\Lambda^2 < 0.037~{\rm GeV}^{-2}$ and $-0.16~{\rm GeV}^{-2} < a_n/\Lambda^2 < 0.15~{\rm GeV}^{-2}$.
- ABBIENDI 04L select 20 $e^+e^- \rightarrow \nu \overline{\nu} \gamma \gamma$ acoplanar events in the energy range 180–209 GeV and 176 $e^+e^- \rightarrow q \overline{q} \gamma \gamma$ events in the energy range 130–209 GeV. These samples are used to constrain possible anomalous $W^+W^-\gamma \gamma$ and $ZZ\gamma\gamma$ quartic couplings. Further combining with the $W^+W^-\gamma$ sample of ABBIENDI 04B the following one-parameter 95% CL limits are obtained: -0.007 $< a_0^Z/\Lambda^2 < 0.023$ GeV⁻², -0.029 $< a_c^Z/\Lambda^2 < 0.029$ GeV⁻², -0.020 $< a_0^W/\Lambda^2 < 0.020$ GeV⁻², -0.052 $< a_c^W/\Lambda^2 < 0.037$ GeV⁻².
- In the CM energy range 183 to 209 GeV HEISTER 04A select 30 $e^+\,e^- \to \nu \overline{\nu} \gamma \gamma$ events with two acoplanar, high energy and high transverse momentum photons. The photon–photon acoplanarity is required to be $>5^\circ$, ${\rm E}_\gamma/\sqrt{s}>0.025$ (the more energetic photon having energy $>0.2~\sqrt{s}$), ${\rm p}_{T\gamma}/{\rm E}_{beam}>0.05$ and $|\cos\theta\gamma|<0.94$. A likelihood fit to the photon energy and recoil missing mass yields the following one–parameter 95% CL limits: $-0.012 < a_0^Z/\Lambda^2 < 0.019~{\rm GeV}^{-2}$, $-0.041 < a_c^Z/\Lambda^2 < 0.044~{\rm GeV}^{-2}$, $-0.060 < a_0^W/\Lambda^2 < 0.055~{\rm GeV}^{-2}$, $-0.099 < a_c^W/\Lambda^2 < 0.093~{\rm GeV}^{-2}$.
- 113 ABDALLAH 03I select 122 $e^+e^- \to W^+W^-\gamma$ events in the C.M. energy range 189–209 GeV, where $E_{\gamma} >$ 5 GeV, the photon has a polar angle $|\cos\theta_{\gamma}| <$ 0.95 and is well isolated from the nearest charged fermion. A fit to the photon energy spectra yields $a_c/\Lambda^2 = 0.000^{+0.019}_{-0.040}~{\rm GeV}^{-2}$, $a_0/\Lambda^2 = -0.004^{+0.018}_{-0.010}~{\rm GeV}^{-2}$, $\tilde{a}_0/\Lambda^2 = -0.007^{+0.019}_{-0.008}~{\rm GeV}^{-2}$, $a_n/\Lambda^2 = -0.09^{+0.16}_{-0.05}~{\rm GeV}^{-2}$, and $\tilde{a}_n/\Lambda^2 = +0.05^{+0.07}_{-0.15}~{\rm GeV}^{-2}$, keeping the other parameters fixed to their Standard Model values (0). The 95% CL limits are: $-0.063~{\rm GeV}^{-2} < a_c/\Lambda^2 < +0.032~{\rm GeV}^{-2}$, $-0.020~{\rm GeV}^{-2} < a_0/\Lambda^2 < +0.020~{\rm GeV}^{-2}$, $-0.020~{\rm GeV}^{-2} < \tilde{a}_0/\Lambda^2 < +0.020~{\rm GeV}^{-2}$, $-0.18~{\rm GeV}^{-2} < a_n/\Lambda^2 < +0.14~{\rm GeV}^{-2}$, $-0.16~{\rm GeV}^{-2} < \tilde{a}_n/\Lambda^2 < +0.17~{\rm GeV}^{-2}$. 114 ACHARD 02F select 86 $e^+e^- \to W^+W^-\gamma$ events at 192–207 GeV, where $E_{\gamma} >$ 5
- GeV and the photon is well isolated. They also select 43 acoplanar $e^+e^- \to \nu \overline{\nu} \gamma \gamma$ events in this energy range, where the photon energies are >5 GeV and >1 GeV and the photon polar angles are between 14° and 166°. All these 43 events are in the recoil mass region corresponding to the Z (75–110 GeV). Using the shape and normalization of the photon spectra in the $W^+W^-\gamma$ events, and combining with the 42 event sample from 189 GeV data (ACCIARRI 00T), they obtain: $a_0/\Lambda^2 = 0.000 \pm 0.010$ GeV $^{-2}$, $a_c/\Lambda^2 = 0.000 \pm 0.010$ GeV $^{-2}$, $a_c/\Lambda^2 = 0.000 \pm 0.010$ GeV $^{-2}$, $a_c/\Lambda^2 = 0.000$

 $-0.013\pm0.023~{\rm GeV}^{-2}$, and $a_n/\Lambda^2 = -0.002\pm0.076~{\rm GeV}^{-2}$. Further combining the analyses of $W^+W^-\gamma$ events with the low recoil mass region of $\nu\overline{\nu}\gamma\gamma$ events (including samples collected at $183+189~{\rm GeV}$), they obtain the following one-parameter 95% CL limits: $-0.015~{\rm GeV}^{-2} < a_0/\Lambda^2 < 0.015~{\rm GeV}^{-2}$, $-0.048~{\rm GeV}^{-2} < a_c/\Lambda^2 < 0.026~{\rm GeV}^{-2}$, and $-0.14~{\rm GeV}^{-2} < a_n/\Lambda^2 < 0.13~{\rm GeV}^{-2}$.

W REFERENCES

SCHAEL	05A	PL B614 7	S. Schael et al.	(ALEPH Collab.)
ABAZOV	04D 04B	PR D70 092008 PL B580 17	V.M. Abazov <i>et al.</i> G. Abbiendi <i>et al.</i>	(CDF, D0 Collab.)
ABBIENDI ABBIENDI	04D	EPJ C33 463	G. Abbiendi <i>et al.</i>	(OPAL Collab.) (OPAL Collab.)
ABBIENDI	04D	PR D70 032005	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABDALLAH	04C	EPJ C34 127	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ACHARD	04G 04D	PL B586 151	P. Achard <i>et al.</i>	(L3 Collab.)
ACHARD	04D	PL B600 22	P. Achard <i>et al.</i>	(L3 Collab.)
HEISTER	04A	PL B602 31	A. Heister <i>et al.</i>	(ALEPH Collab.)
SCHAEL	04A	EPJ C38 147	S. Schael <i>et al.</i>	(ALEPH Collab.)
ABBIENDI	03C	EPJ C26 321	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABDALLAH	031	EPJ C31 139	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ABAZOV	02D	PR D66 012001	V.M. Abazov et al.	` (D0 Collab.)
ABAZOV	02E	PR D66 032008	V.M. Abazov et al.	(D0 Collab.)
ACHARD	02F	PL B527 29	P. Achard et al.	(L3 Collab.)
CHEKANOV	02C	PL B539 197	S. Chekanov et al.	(ZEÙS Collab.)
ABBIENDI	01F	PL B507 29	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	01H	EPJ C19 229	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	011	PL B502 9	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	01K	PL B511 159	P. Abreu et al.	(DELPHI Collab.)
ADLOFF	01A	EPJ C19 269	C. Adloff et al.	(H1 Collab.)
AFFOLDER	01E	PR D64 052001	T. Affolder et al.	(CDF Collab.)
ABBIENDI	00V	PL B490 71	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI,G	00	PL B493 249	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBOTT	00 00D	PRL 84 222	B. Abbott <i>et al.</i> B. Abbott <i>et al.</i>	(D0 Collab.)
ABBOTT ABBOTT	00B 00D	PR D61 072001 PRL 84 5710	B. Abbott <i>et al.</i>	(D0 Collab.)
ABREU	00D	PL B479 89	P. Abreu <i>et al.</i>	(D0 Collab.) (DELPHI Collab.)
ABREU,P	00F	EPJ C18 203	P. Abreu <i>et al.</i>	(DELPHI Collab.)
Also	02	EPJ C25 493 (erratum)		(DELPHI Collab.)
ACCIARRI	00T	PL B490 187	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACCIARRI	00V	PL B496 19	M. Acciarri <i>et al.</i>	(L3 Collab.)
ADLOFF	00B	EPJ C13 609	C. Adloff et al.	(H1 Collab.)
AFFOLDER	M00	PRL 85 3347	T. Affolder et al.	(CDF Collab.)
BARATE	00J	PL B484 205	R. Barate et al.	(ALEPH Collab.)
BARATE	00T	EPJ C17 241	R. Barate <i>et al.</i>	(ALEPH Collab.)
BREITWEG	00	PL B471 411	J. Breitweg <i>et al.</i>	(ZEUS Collab.)
BREITWEG	00D	EPJ C12 411	J. Breitweg <i>et al.</i>	(ZEUS Collab.)
EBOLI	00	MPL A15 1	O. Eboli, M. Gonzalez-Garcia,	
ABBIENDI	99C	PL B453 138	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	99D	EPJ C8 191	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	99N	PL B453 153	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBOTT ABBOTT	99H 99I	PR D60 052003 PR D60 072002	B. Abbott <i>et al.</i> B. Abbott <i>et al.</i>	(D0 Collab.)
ABREU	99L	PL B459 382	P. Abreu <i>et al.</i>	(D0 Collab.) (DELPHI Collab.)
ABREU	99T	PL B462 410	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	99	PL B454 386	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACCIARRI	99Q	PL B467 171	M. Acciarri <i>et al.</i>	(L3 Collab.)
BARATE	99	PL B453 121	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	991	PL B453 107	R. Barate et al.	(ALEPH Collab.)
BARATE	99L	PL B462 389	R. Barate et al.	(ALEPH Collab.)
BARATE	99M	PL B465 349	R. Barate <i>et al.</i>	(ALEPH Collab.)
ABBOTT	98N	PR D58 092003	B. Abbott <i>et al.</i>	(D0 Collab.)
ABBOTT	98P	PR D58 012002	B. Abbott <i>et al.</i>	(D0 Collab.)
ABE	98H	PR D58 031101	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	98P	PR D58 091101	F. Abe <i>et al.</i>	(CDF Collab.)
ABREU	98C	PL B416 233	P. Abreu et al.	(DELPHI Collab.)
ABREU ACCIARRI	98N 97	PL B439 209 PL B398 223	P. Abreu <i>et al.</i> M. Acciarri <i>et al.</i>	(DELPHI Collab.) (L3 Collab.)
ACCIAINI	51	1 L D390 223	ivi. Acciaili et al.	(L3 Collab.)

ACCIARRI ACCIARRI ABACHI ABE ABE AIso ABE AIso ABE AIso ABE AIITTI ALITTI ALITTI ALITTI	97M 97S 95D 95C 95G 95P 95Q 95W 94B 92E 92I 92 92B 92C 92D 92F	PL B413 176 PRL 75 1456 PRL 74 341 PRL 74 1936 PRL 75 11 PR D52 4784	M. Acciarri et al. M. Acciarri et al. S. Abachi et al. F. Abe et al. J. Alitti et al.	(L3 Collab.) (L3 Collab.) (D0 Collab.) (CDF Collab.) (UA2 Collab.) (UA2 Collab.) (UA2 Collab.) (UA2 Collab.) (UA2 Collab.) (UA2 Collab.)
SAMUEL	92	PL B280 124	M.A. Samuel <i>et al.</i>	(OKSU, CARL)
ABE	91C	PR D44 29	F. Abe <i>et al.</i>	(CDF Collab.)
ALBAJAR	91	PL B253 503	C. Albajar <i>et al.</i>	(UA1 Collab.)
ALITTI	91C	ZPHY C52 209	J. Alitti <i>et al.</i>	(UA2 Collab.)
SAMUEL	91	PRL 67 9	M.A. Samuel et al.	(OKSU, CARL)
Also	91C	PRL 67 2920 erratum	M.A. Samuel <i>et al.</i>	(65 - 6)
ABE	90	PRL 64 152	F. Abe <i>et al.</i>	(CDF Collab.)
Also	91C	PR D44 29	F. Abe <i>et al.</i>	(CDF Collab.)
ABE Also	90G 91B	PRL 65 2243	F. Abe et al.	(CDF Collab.)
ALBAJAR	916	PR D43 2070 PL B241 283	F. Abe <i>et al.</i> C. Albajar <i>et al.</i>	(CDF Collab.)
ALITTI	90 90B	PL B241 203 PL B241 150	J. Alitti <i>et al.</i>	(UA1 Collab.) (UA2 Collab.)
ALITTI	90C	ZPHY C47 11	J. Alitti <i>et al.</i>	(UA2 Collab.)
ABE	89I	PRL 62 1005	F. Abe <i>et al.</i>	(CDF Collab.)
ALBAJAR	89	ZPHY C44 15	C. Albajar <i>et al.</i>	(UA1 Collab.)
BAUR	88	NP B308 127	U. Baur, D. Zeppenfeld	(FSU, WISC)
GRIFOLS	88	IJMP A3 225	J.A. Grifols, S. Peris, J. Sola	(BARC, DESY)
Also	87	PL B197 437	J.A. Grifols, S. Peris, J. Sola	(BARC, DESY)
ALBAJAR	87	PL B185 233	C. Albajar et al.	`(UA1 Collab.)
ANSARI	87	PL B186 440	R. Ansari <i>et al.</i>	(UA2 Collab.)
ANSARI	87C	PL B194 158	R. Ansari <i>et al.</i>	(UA2 Collab.)
GROTCH	87	PR D36 2153	H. Grotch, R.W. Robinett	(PSU)
HAGIWARA	87	NP B282 253	K. Hagiwara <i>et al.</i>	(KEK, UCLA, FSU)
VANDERBIJ	87	PR D35 1088	J.J. van der Bij	(FNAL)
APPEL	86	ZPHY C30 1	J.A. Appel <i>et al.</i>	(UA2 Collab.)
ARNISON	86	PL 166B 484	G.T.J. Arnison et al.	(UA1 Collab.) J
ALTARELLI	85B	ZPHY C27 617	G. Altarelli, R.K. Ellis, G. Martinelli	(CERN+)
GRAU	85	PL 154B 283	A. Grau, J.A. Grifols	(BARC)
SUZUKI	85	PL 153B 289	M. Suzuki	(LBL)
ARNISON	84D	PL 134B 469	G.T.J. Arnison <i>et al.</i>	(UA1 Collab.)
HERZOG Also	84 84B	PL 148B 355 PL 155B 468 erratum	F. Herzog F. Herzog	(WISC) (WISC)
ARNISON	83	PL 133B 406 erratum PL 122B 103	G.T.J. Arnison <i>et al.</i>	(WISC) (UA1 Collab.)
BANNER	83B	PL 122B 103 PL 122B 476	M. Banner <i>et al.</i>	(UA2 Collab.)
D/ MININE IN	000	1 1220 710	Danner et al.	(O/12 Collab.)