

$\rho(1700)$

$I^G(J^{PC}) = 1^+(1^{--})$

## THE $\rho(1450)$ AND THE $\rho(1700)$

Updated March 2002 by S. Eidelman (Novosibirsk) and J.J. Hernández-Rey (IFIC, Valencia).

In our 1988 edition, we replaced the  $\rho(1600)$  entry with two new ones, the  $\rho(1450)$  and the  $\rho(1700)$ , because there was emerging evidence that the 1600-MeV region actually contains two  $\rho$ -like resonances. ERKAL 86 had pointed out this possibility with a theoretical analysis on the consistency of  $2\pi$  and  $4\pi$  electromagnetic form factors and the  $\pi\pi$  scattering length. DONNACHIE 87, with a full analysis of data on the  $2\pi$  and  $4\pi$  final states in  $e^+e^-$  annihilation and photoproduction reactions, had also argued that in order to obtain a consistent picture, two resonances were necessary. The existence of  $\rho(1450)$  was supported by the analysis of  $\eta\rho^0$  mass spectra obtained in photoproduction and  $e^+e^-$  annihilation (DONNACHIE 87B), as well as that of  $e^+e^- \rightarrow \omega\pi$  (DONNACHIE 91).

The analysis of DONNACHIE 87 was further extended by CLEGG 88,94 to include new data on  $4\pi$  systems produced in  $e^+e^-$  annihilation, and in  $\tau$  decays ( $\tau$  decays to  $4\pi$  and  $e^+e^-$  annihilation to  $4\pi$  can be related by the Conserved Vector Current assumption). These systems were successfully analyzed using interfering contributions from two  $\rho$ -like states, and from the tail of the  $\rho(770)$  decaying into two-body states. While specific conclusions on  $\rho(1450) \rightarrow 4\pi$  were obtained, little could be said about the  $\rho(1700)$ .

An analysis by CLEGG 90 of  $6\pi$  mass spectra from  $e^+e^-$  annihilation and diffractive photoproduction provides evidence for two  $\rho$  mesons at about 2.1 and 1.8 GeV that decay strongly into  $6\pi$  states. While the former is a candidate for a new

resonance ( $\rho(2150)$ )), the latter could be a manifestation of the  $\rho(1700)$  distorted by threshold effects.

Independent evidence for two  $1^-$  states is provided by **KILLIAN** 80 in  $4\pi$  electroproduction at  $\langle Q^2 \rangle = 1$   $(\text{GeV}/c)^2$ , and by **FUKUI** 88 in a high-statistics sample of the  $\eta\pi\pi$  system in  $\pi^-p$  charge exchange.

This scenario with two overlapping resonances is supported by other data. **BISELLO** 89 measured the pion form factor in the interval 1.35–2.4 GeV and observed a deep minimum around 1.6 GeV. The best fit was obtained with the hypothesis of  $\rho$ -like resonances at 1420 and 1770 MeV, with widths of about 250 MeV. **ANTONELLI** 88 found that the  $e^+e^- \rightarrow \eta\pi^+\pi^-$  cross section is better fitted with two fully interfering Breit-Wigners, with parameters in fair agreement with those of **DONNACHIE** 87 and **BISELLO** 89. These results can be considered as a confirmation of the  $\rho(1450)$ .

Decisive evidence for the  $\pi\pi$  decay mode of both  $\rho(1450)$  and  $\rho(1700)$  came from recent results in  $\bar{p}p$  annihilation at rest (**ABELE** 97). It was shown that these resonances also possess a  $K\bar{K}$  decay mode (**ABELE** 98, **BERTIN** 98B, **ABELE** 99D). High statistics studies of the decays  $\tau \rightarrow \pi\pi\nu_\tau$  (**BARATE** 97M, **URHEIM** 97), and  $\tau \rightarrow 4\pi\nu_\tau$  (**EDWARDS** 00), also require the  $\rho(1450)$ , but are not sensitive to the  $\rho(1700)$ , because it is too close to the  $\tau$  mass.

The structure of these  $\rho$  states is not yet completely clear. **BARNES** 97 and **CLOSE** 97C claim that  $\rho(1450)$  has a mass consistent with radial  $2S$ , but its decays show characteristics of hybrids, and suggest that this state may be a  $2S$ -hybrid mixture. **DONNACHIE** 99 argues that hybrid states could have a  $4\pi$  decay mode dominated by the  $a_1\pi$ . Such behavior has recently been observed by **AKHMETSHIN** 99E in  $e^+e^- \rightarrow 4\pi$  in the energy range 1.05–1.38 GeV, and by **EDWARDS** 00 in  $\tau \rightarrow 4\pi$ .

decays. More data should be collected to clarify the nature of the  $\rho$  states, particularly in the energy range above 1.6 GeV.

We also list the  $\phi\pi$  state with  $J^{PC} = 1^{--}$  or  $C(1480)$  observed by **BITYUKOV** 87 under the  $\rho(1450)$ . While **ACHASOV** 96B shows that it may be a threshold effect, **CLEGG** 88 and **LANDSBERG** 92 suggest two independent vector states with this decay mode. Note, however, that  $C(1480)$  in its  $\phi\pi$  decay mode was not confirmed by  $e^+e^-$  (**DOLINSKY** 91, **BISELLO** 91C) and  $\bar{p}p$  (**ABELE** 97H) experiments.

Several observations on the  $\omega\pi$  system in the 1200-MeV region (**FRENKIEL** 72, **COSME** 76, **BARBER** 80C, **ASTON** 80C, **ATKINSON** 84C, **BRAU** 88, **AMSLER** 93B) may be interpreted in terms of either  $J^P = 1^- \rho(770) \rightarrow \omega\pi$  production (**LAYSSAC** 71), or  $J^P = 1^+ b_1(1235)$  production (**BRAU** 88, **AMSLER** 93B). We argue that no special entry for a  $\rho(1250)$  is needed. The LASS amplitude analysis (**ASTON** 91B) showing evidence for  $\rho(1270)$  is preliminary and needs confirmation. For completeness, the relevant observations are listed under the  $\rho(1450)$ .

## References

References may be found at the end of the  $\rho(1770)$  Listing.

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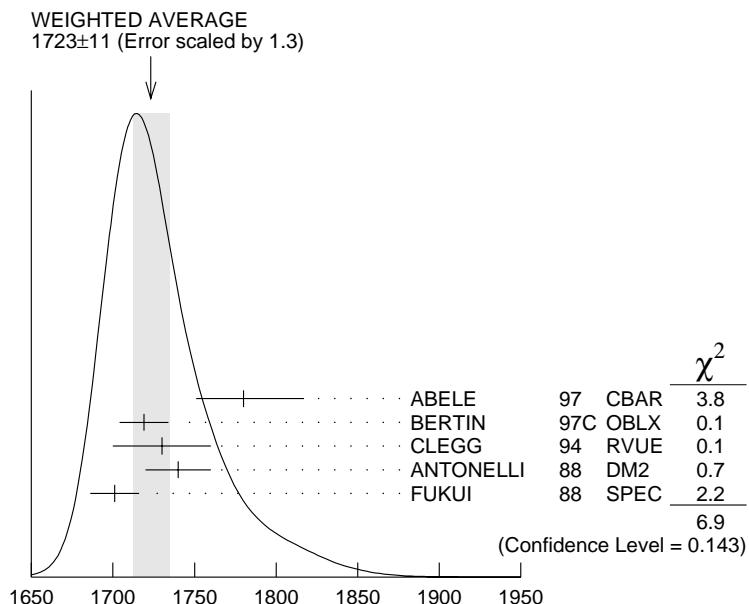
## $\rho(1700)$ MASS

### $\eta\rho^0$ AND $\pi^+\pi^-$ MODES

VALUE (MeV) DOCUMENT ID

**1700±20 OUR ESTIMATE**

**1723±11 OUR AVERAGE** Includes data from the 2 datablocks that follow this one. Error includes scale factor of 1.3. See the ideogram below.



$\rho(1700)$  mass,  $\eta\rho^0$  and  $\pi^+\pi^-$  modes (MeV)

### $\eta\rho^0$ MODE

VALUE (MeV) DOCUMENT ID TECN COMMENT

The data in this block is included in the average printed for a previous datablock.

1740±20	ANTONELLI	88	DM2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
1701±15	<sup>2</sup> FUKUI	88	SPEC	$8.95\pi^-p \rightarrow \eta\pi^+\pi^-n$

### $\pi\pi$ MODE

VALUE (MeV) DOCUMENT ID TECN COMMENT

The data in this block is included in the average printed for a previous datablock.

1780 $\pm^{37}_{-29}$	<sup>3</sup> ABELE	97	CBAR	$\bar{p}n \rightarrow \pi^-\pi^0\pi^0$
1719 $\pm 15$	<sup>3</sup> BERTIN	97C	OBLEX	$0.0\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
1730 $\pm 30$	CLEGG	94	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1768 ± 21	BISELLLO	89	DM2	$e^+ e^- \rightarrow \pi^+ \pi^-$
1745.7 ± 91.9	DUBNICKA	89	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
1546 ± 26	GESHKEN...	89	RVUE	
1650	<sup>4</sup> ERKAL	85	RVUE	20–70 $\gamma p \rightarrow \gamma \pi$
1550 ± 70	ABE	84B	HYBR	20 $\gamma p \rightarrow \pi^+ \pi^- p$
1590 ± 20	<sup>5</sup> ASTON	80	OMEG	20–70 $\gamma p \rightarrow p 2\pi$
1600 ± 10	<sup>6</sup> ATIYA	79B	SPEC	50 $\gamma C \rightarrow C 2\pi$
1598 $\pm^{+24}_{-22}$	BECKER	79	ASPK	17 $\pi^- p$ polarized
1659 ± 25	<sup>4</sup> LANG	79	RVUE	
1575	<sup>4</sup> MARTIN	78C	RVUE	17 $\pi^- p \rightarrow \pi^+ \pi^- n$
1610 ± 30	<sup>4</sup> FROGGATT	77	RVUE	17 $\pi^- p \rightarrow \pi^+ \pi^- n$
1590 ± 20	<sup>7</sup> HYAMS	73	ASPK	17 $\pi^- p \rightarrow \pi^+ \pi^- n$

## $\pi\omega$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1550 to 1620	<sup>8</sup> ACHASOV	00I	SND	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$
1580 to 1710	<sup>9</sup> ACHASOV	00I	SND	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$
1710 ± 90	ACHASOV	97	RVUE	$e^+ e^- \rightarrow \omega \pi^0$

## $K\bar{K}$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1740.8 ± 22.2	27k	<sup>1</sup> ABELE	99D	CBAR	$\pm$ 0.0 $\bar{p}p \rightarrow K^+ K^- \pi^0$
1582 ± 36	1600	CLELAND	82B	SPEC	$\pm$ 50 $\pi p \rightarrow K_S^0 K^\pm p$

<sup>1</sup> K-matrix pole. Isospin not determined, could be  $\omega(1650)$  or  $\phi(1680)$ .

## $2(\pi^+ \pi^-)$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1851 $\pm^{27}_{24}$		ACHASOV	97	RVUE	$e^+ e^- \rightarrow 2(\pi^+ \pi^-)$
1570 ± 20		<sup>10</sup> CORDIER	82	DM1	$e^+ e^- \rightarrow 2(\pi^+ \pi^-)$
1520 ± 30		<sup>5</sup> ASTON	81E	OMEG	20–70 $\gamma p \rightarrow p 4\pi$
1654 ± 25		<sup>11</sup> DIBIANCA	81	DBC	$\pi^+ d \rightarrow p p 2(\pi^+ \pi^-)$
1666 ± 39		<sup>10</sup> BACCI	80	FRAG	$e^+ e^- \rightarrow 2(\pi^+ \pi^-)$
1780	34	KILLIAN	80	SPEC	11 $e^- p \rightarrow 2(\pi^+ \pi^-)$
1500		<sup>12</sup> ATIYA	79B	SPEC	50 $\gamma C \rightarrow C 4\pi^\pm$
1570 ± 60	65	<sup>13</sup> ALEXANDER	75	HBC	7.5 $\gamma p \rightarrow p 4\pi$
1550 ± 60		<sup>5</sup> CONVERSI	74	OSPK	$e^+ e^- \rightarrow 2(\pi^+ \pi^-)$
1550 ± 50	160	SCHACHT	74	STRC	5.5–9 $\gamma p \rightarrow p 4\pi$
1450 ± 100	340	SCHACHT	74	STRC	9–18 $\gamma p \rightarrow p 4\pi$
1430 ± 50	400	BINGHAM	72B	HBC	9.3 $\gamma p \rightarrow p 4\pi$

## $\pi^+\pi^-\pi^0\pi^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>			
1660 $\pm$ 30	ATKINSON	85B OMEG	20–70 $\gamma p$

## 3( $\pi^+\pi^-$ ) AND 2( $\pi^+\pi^-\pi^0$ ) MODES

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>			
1783 $\pm$ 15	CLEGG	90 RVUE	$e^+e^- \rightarrow 3(\pi^+\pi^-)2(\pi^+\pi^-\pi^0)$

<sup>2</sup> Assuming  $\rho^+ f_0(1370)$  decay mode interferes with  $a_1(1260)^+\pi$  background. From a two Breit-Wigner fit.

<sup>3</sup> T-matrix pole.

<sup>4</sup> From phase shift analysis of HYAMS 73 data.

<sup>5</sup> Simple relativistic Breit-Wigner fit with constant width.

<sup>6</sup> An additional 40 MeV uncertainty in both the mass and width is present due to the choice of the background shape.

<sup>7</sup> Included in BECKER 79 analysis.

<sup>8</sup> Taking into account both  $\rho(1450)$  and  $\rho(1700)$  contributions. Using the data of ACHASOV 00I on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .  $\rho(1450)$  mass and width fixed at 1400 MeV and 500 MeV respectively.

<sup>9</sup> Taking into account the  $\rho(1700)$  contribution only. Using the data of ACHASOV 00I on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .

<sup>10</sup> Simple relativistic Breit-Wigner fit with model dependent width.

<sup>11</sup> One peak fit result.

<sup>12</sup> Parameters roughly estimated, not from a fit.

<sup>13</sup> Skew mass distribution compensated by Ross-Stodolsky factor.

## $\rho(1700)$ WIDTH

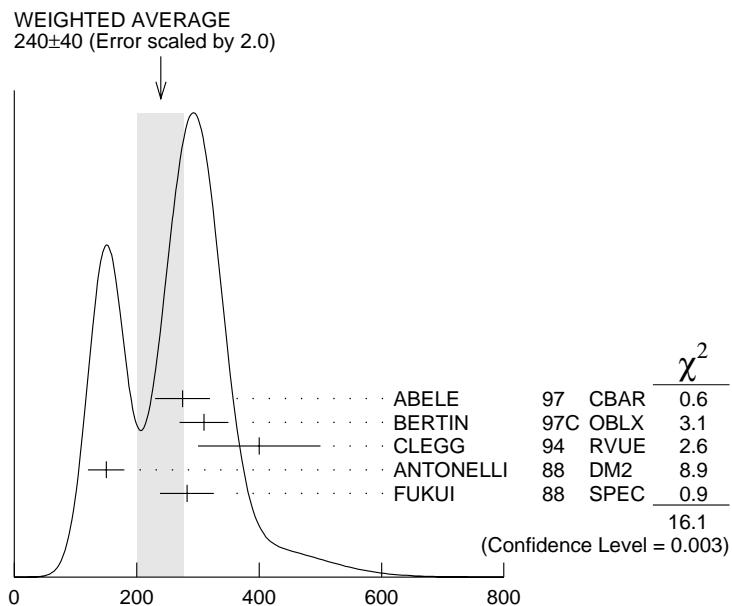
### $\eta\rho^0$ AND $\pi^+\pi^-$ MODES

VALUE (MeV)

DOCUMENT ID

#### 240 $\pm$ 60 OUR ESTIMATE

240 $\pm$ 40 OUR AVERAGE Includes data from the 2 datablocks that follow this one. Error includes scale factor of 2.0. See the ideogram below.



$\rho(1700)$  width,  $\eta\rho^0$  and  $\pi^+\pi^-$  modes (MeV)

### $\eta\rho^0$ MODE

VALUE (MeV)

DOCUMENT ID

TECN

COMMENT

The data in this block is included in the average printed for a previous datablock.

150 $\pm$ 30

ANTONELLI 88 DM2  $e^+e^- \rightarrow \eta\pi^+\pi^-$   
15 FUKUI 88 SPEC  $8.95\pi^-p \rightarrow \eta\pi^+\pi^-n$

### $\pi\pi$ MODE

VALUE (MeV)

DOCUMENT ID

TECN

COMMENT

The data in this block is included in the average printed for a previous datablock.

275  $\pm$  45

16 ABELE 97 CBAR  $\bar{p}n \rightarrow \pi^-\pi^0\pi^0$   
16 BERTIN 97C OBLX  $0.0\bar{p}p \rightarrow \pi^+\pi^-\pi^0$   
400  $\pm$  100 CLEGG 94 RVUE  $e^+e^- \rightarrow \pi^+\pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

224 ± 22	BISELLLO	89	DM2	$e^+ e^- \rightarrow \pi^+ \pi^-$
242.5 ± 163.0	DUBNICKA	89	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
620 ± 60	GESHKEN...	89	RVUE	
<315	17 ERKAL	85	RVUE	20–70 $\gamma p \rightarrow \gamma \pi$
280 ± 30 — 80	ABE	84B	HYBR	20 $\gamma p \rightarrow \pi^+ \pi^- p$
230 ± 80	18 ASTON	80	OMEG	20–70 $\gamma p \rightarrow p 2\pi$
283 ± 14	19 ATIYA	79B	SPEC	50 $\gamma C \rightarrow C 2\pi$
175 ± 98 — 53	BECKER	79	ASPK	17 $\pi^- p$ polarized
232 ± 34	17 LANG	79	RVUE	
340	17 MARTIN	78C	RVUE	17 $\pi^- p \rightarrow \pi^+ \pi^- n$
300 ± 100	17 FROGGATT	77	RVUE	17 $\pi^- p \rightarrow \pi^+ \pi^- n$
180 ± 50	20 HYAMS	73	ASPK	17 $\pi^- p \rightarrow \pi^+ \pi^- n$

## K $\bar{K}$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
187.2 ± 26.7	27k	14 ABELE	99D	CBAR	± 0.0 $\bar{p} p \rightarrow K^+ K^- \pi^0$
265 ± 120	1600	CLELAND	82B	SPEC	± 50 $\pi p \rightarrow K_S^0 K^\pm p$

14 K-matrix pole. Isospin not determined, could be  $\omega(1650)$  or  $\phi(1680)$ .

## 2( $\pi^+ \pi^-$ ) MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
510 ± 40	21 CORDIER	82	DM1	$e^+ e^- \rightarrow 2(\pi^+ \pi^-)$
400 ± 50	18 ASTON	81E	OMEG	20–70 $\gamma p \rightarrow p 4\pi$
400 ± 146	22 DIBIANCA	81	DBC	$\pi^+ d \rightarrow p p 2(\pi^+ \pi^-)$
700 ± 160	21 BACCI	80	FRAG	$e^+ e^- \rightarrow 2(\pi^+ \pi^-)$
100	34 KILLIAN	80	SPEC	11 $e^- p \rightarrow 2(\pi^+ \pi^-)$
600	23 ATIYA	79B	SPEC	50 $\gamma C \rightarrow C 4\pi^\pm$
340 ± 160	65 ALEXANDER	75	HBC	7.5 $\gamma p \rightarrow p 4\pi$
360 ± 100	18 CONVERSI	74	OSPK	$e^+ e^- \rightarrow 2(\pi^+ \pi^-)$
400 ± 120	160 SCHACHT	74	STRC	5.5–9 $\gamma p \rightarrow p 4\pi$
850 ± 200	340 SCHACHT	74	STRC	9–18 $\gamma p \rightarrow p 4\pi$
650 ± 100	400 BINGHAM	72B	HBC	9.3 $\gamma p \rightarrow p 4\pi$

## $\pi^+ \pi^- \pi^0 \pi^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
300 ± 50	ATKINSON	85B	OMEG 20–70 $\gamma p$

## $\omega \pi^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
350 to 580	26 ACHASOV	00I	SND $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$
490 to 1040	27 ACHASOV	00I	SND $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$

**3( $\pi^+\pi^-$ ) AND 2( $\pi^+\pi^-\pi^0$ ) MODES**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
285±20	CLEGG	90 RVUE	$e^+e^- \rightarrow 3(\pi^+\pi^-)2(\pi^+\pi^-\pi^0)$
15 Assuming $\rho^+ f_0(1370)$ decay mode interferes with $a_1(1260)^+\pi^-$ background. From a two Breit-Wigner fit.			
16 T-matrix pole.			
17 From phase shift analysis of HYAMS 73 data.			
18 Simple relativistic Breit-Wigner fit with constant width.			
19 An additional 40 MeV uncertainty in both the mass and width is present due to the choice of the background shape.			
20 Included in BECKER 79 analysis.			
21 Simple relativistic Breit-Wigner fit with model-dependent width.			
22 One peak fit result.			
23 Parameters roughly estimated, not from a fit.			
24 Skew mass distribution compensated by Ross-Stodolsky factor.			
25 Width errors enlarged by us to $4\Gamma/\sqrt{N}$ ; see the note with the $K^*(892)$ mass.			
26 Taking into account both $\rho(1450)$ and $\rho(1700)$ contributions. Using the data of ACHASOV 00I on $e^+e^- \rightarrow \omega\pi^0$ and of EDWARDS 00A on $\tau^- \rightarrow \omega\pi^-\nu_\tau$ . $\rho(1450)$ mass and width fixed at 1400 MeV and 500 MeV respectively.			
27 Taking into account the $\rho(1700)$ contribution only. Using the data of ACHASOV 00I on $e^+e^- \rightarrow \omega\pi^0$ and of EDWARDS 00A on $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .			

 **$\rho(1700)$  DECAY MODES**

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $4\pi$	
$\Gamma_2$ $2(\pi^+\pi^-)$	large
$\Gamma_3$ $\rho\pi\pi$	dominant
$\Gamma_4$ $\rho^0\pi^+\pi^-$	large
$\Gamma_5$ $\rho^0\pi^0\pi^0$	
$\Gamma_6$ $\rho^\pm\pi^\mp\pi^0$	large
$\Gamma_7$ $a_1(1260)\pi$	
$\Gamma_8$ $h_1(1170)\pi$	
$\Gamma_9$ $\pi(1300)\pi$	
$\Gamma_{10}$ $\rho\rho$	
$\Gamma_{11}$ $\pi^+\pi^-$	seen
$\Gamma_{12}$ $\pi\pi$	seen
$\Gamma_{13}$ $K\bar{K}^*(892)+\text{c.c.}$	seen
$\Gamma_{14}$ $\eta\rho$	seen
$\Gamma_{15}$ $a_2(1320)\pi$	not seen
$\Gamma_{16}$ $K\bar{K}$	seen
$\Gamma_{17}$ $e^+e^-$	seen
$\Gamma_{18}$ $\pi^0\omega$	seen

**$\rho(1700) \Gamma(i)\Gamma(e^+e^-)/\Gamma_{\text{total}}$**

This combination of a partial width with the partial width into  $e^+e^-$  and with the total width is obtained from the cross-section into channel  $i$  in  $e^+e^-$  annihilation.

**$\Gamma(2(\pi^+\pi^-)) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$**

**$\Gamma_2\Gamma_{17}/\Gamma$**

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>2.83±0.42</b>	BACCI	80	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
2.6 ±0.2	DELCOURT	81B DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$

**$\Gamma(\pi^+\pi^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$**

**$\Gamma_{11}\Gamma_{17}/\Gamma$**

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
0.13	28 DIEKMAN	88 RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
$0.029^{+0.016}_{-0.012}$	KURDADZE	83 OLYA	$0.64-1.4 e^+e^- \rightarrow \pi^+\pi^-$

28 Using total width = 220 MeV.

**$\Gamma(K\bar{K}^*(892)+\text{c.c.}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$**

**$\Gamma_{13}\Gamma_{17}/\Gamma$**

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
0.305±0.071	29 BIZOT	80 DM1	$e^+e^-$

**$\Gamma(\eta\rho) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$**

**$\Gamma_{14}\Gamma_{17}/\Gamma$**

VALUE (eV)	DOCUMENT ID	TECN	COMMENT
<b>7 ±3</b>	ANTONELLI	88 DM2	$e^+e^- \rightarrow \eta\pi^+\pi^-$

**$\Gamma(K\bar{K}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$**

**$\Gamma_{16}\Gamma_{17}/\Gamma$**

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
0.035±0.029	29 BIZOT	80 DM1	$e^+e^-$

**$\Gamma(\rho\pi\pi) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$**

**$\Gamma_3\Gamma_{17}/\Gamma$**

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
3.510±0.090	29 BIZOT	80 DM1	$e^+e^-$
<b>29 Model dependent.</b>			

## $\rho(1700)$ BRANCHING RATIOS

### $\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{11}/\Gamma$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$0.287^{+0.043}_{-0.042}$	BECKER 79	ASPK	$17 \pi^- p$ polarized	
$0.15$ to $0.30$	MARTIN 30	RVUE	$17 \pi^- p \rightarrow \pi^+ \pi^- n$	
$<0.20$	COSTA... 31	RVUE	$e^+ e^- \rightarrow 2\pi, 4\pi$	
$0.30 \pm 0.05$	FROGGATT 30	RVUE	$17 \pi^- p \rightarrow \pi^+ \pi^- n$	
$<0.15$	EISENBERG 32	HBC	$5 \pi^+ p \rightarrow \Delta^{++} 2\pi$	
$0.25 \pm 0.05$	HYAMS 33	ASPK	$17 \pi^- p \rightarrow \pi^+ \pi^- n$	

<sup>30</sup> From phase shift analysis of HYAMS 73 data.

<sup>31</sup> Estimate using unitarity, time reversal invariance, Breit-Wigner.

<sup>32</sup> Estimated using one-pion-exchange model.

<sup>33</sup> Included in BECKER 79 analysis.

### $\Gamma(\pi^+\pi^-)/\Gamma(2(\pi^+\pi^-))$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{11}/\Gamma_2$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$0.13 \pm 0.05$	ASTON 80	OMEG	$20-70 \gamma p \rightarrow p 2\pi$	
$<0.14$	DAVIER 34	STRC	$6-18 \gamma p \rightarrow p 4\pi$	
$<0.2$	BINGHAM 35	HBC	$9.3 \gamma p \rightarrow p 2\pi$	

<sup>34</sup> Upper limit is estimate.

<sup>35</sup>  $2\sigma$  upper limit.

### $\Gamma(\pi\pi)/\Gamma(4\pi)$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{12}/\Gamma_1$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$0.16 \pm 0.04$	ABELE 40,41	CBAR	$0.0 \bar{p}n \rightarrow 5\pi$	

### $\Gamma(K\bar{K}^*(892)+\text{c.c.})/\Gamma(2(\pi^+\pi^-))$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{13}/\Gamma_2$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$0.15 \pm 0.03$	DEL COURT 36	DM1	$e^+ e^- \rightarrow \bar{K}K\pi$	

<sup>36</sup> Assuming  $\rho(1700)$  and  $\omega$  radial excitations to be degenerate in mass.

### $\Gamma(\eta\rho)/\Gamma_{\text{total}}$

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{14}/\Gamma$
$<0.04$		DONNACHIE 87B	RVUE		
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
possibly seen		AKHMETSHIN 00D	CMD2	$e^+ e^- \rightarrow \eta\pi^+\pi^-$	
$<0.02$	58	ATKINSON 86B	OMEG	$20-70 \gamma p$	

### $\Gamma(a_2(1320)\pi)/\Gamma_{\text{total}}$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{15}/\Gamma$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
not seen	AMELIN 00	VES	$37 \pi^- p \rightarrow \eta\pi^+\pi^- n$	

$\Gamma(\eta\rho)/\Gamma(2(\pi^+\pi^-))$

$\Gamma_{14}/\Gamma_2$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. • • •			
0.123±0.027	DELCOURT 82	DM1	$e^+e^- \rightarrow \pi^+\pi^-$ MM
~0.1	ASTON 80	OMEG	20–70 $\gamma p$

$\Gamma(\pi^+\pi^- \text{ neutrals})/\Gamma(2(\pi^+\pi^-))$

$(\Gamma_5 + \Gamma_6 + 0.714\Gamma_{14})/\Gamma_2$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. • • •			
2.6±0.4	37 BALLAM	74	HBC 9.3 $\gamma p$
37 Upper limit. Background not subtracted.			

$\Gamma(\pi^0\omega)/\Gamma_{\text{total}}$

$\Gamma_{18}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. • • •			
seen	ACHASOV 97	RVUE	$e^+e^- \rightarrow \omega\pi^0$

$\Gamma(a_1(1260)\pi)/\Gamma(4\pi)$

$\Gamma_7/\Gamma_1$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. • • •			
0.16±0.05	40 ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$

$\Gamma(h_1(1170)\pi)/\Gamma(4\pi)$

$\Gamma_8/\Gamma_1$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. • • •			
0.17±0.06	40 ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$

$\Gamma(\pi(1300)\pi)/\Gamma(4\pi)$

$\Gamma_9/\Gamma_1$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. • • •			
0.30±0.10	40 ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$

$\Gamma(\rho\rho)/\Gamma(4\pi)$

$\Gamma_{10}/\Gamma_1$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. • • •			
0.09±0.03	40 ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$

$\Gamma(\rho\pi\pi)/\Gamma(4\pi)$

$\Gamma_3/\Gamma_1$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. • • •			
0.28±0.06	40 ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$

$\Gamma(K\bar{K})/\Gamma(2(\pi^+\pi^-))$

$\Gamma_{16}/\Gamma_2$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. • • •					
0.015±0.010	38 DELCOURT	81B DM1			$e^+e^- \rightarrow K\bar{K}$
<0.04	95 BINGHAM	72B HBC	0		9.3 $\gamma p$

38 Assuming  $\rho(1700)$  and  $\omega$  radial excitations to be degenerate in mass.

$\Gamma(K\bar{K})/\Gamma(K\bar{K}^*(892)+c.c.)$  $\Gamma_{16}/\Gamma_{13}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>			
0.052 ± 0.026	BUON	82 DM1	$e^+e^- \rightarrow \text{hadrons}$

 $\Gamma(\rho^0\pi^+\pi^-)/\Gamma(2(\pi^+\pi^-))$  $\Gamma_4/\Gamma_2$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
~ 1.0		DELCOURT	81B DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
0.7 ± 0.1	500	SCHACHT	74 STRC	$5.5-18 \gamma p \rightarrow p4\pi$
0.80	39	BINGHAM	72B HBC	$9.3 \gamma p \rightarrow p4\pi$

<sup>39</sup> The  $\pi\pi$  system is in *S*-wave. $\Gamma(\rho^0\pi^0\pi^0)/\Gamma(\rho^\pm\pi^\mp\pi^0)$  $\Gamma_5/\Gamma_6$ 

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
<0.10	ATKINSON	85B OMEG		20-70 $\gamma p$
<0.15	ATKINSON	82 OMEG 0		$20-70 \gamma p \rightarrow p4\pi$

<sup>40</sup>  $\omega\pi$  not included.<sup>41</sup> Using ABELE 97.**p(1700) REFERENCES**

ABELE	01B	EPJ C21 261	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ACHASOV	00I	PL B486 29	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AKHMETSHIN	00D	PL B489 125	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AMELIN	00	NP A668 83	D. Amelin <i>et al.</i>	(VES Collab.)
EDWARDS	00A	PR D61 072003	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
ABELE	99D	PL B468 178	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE	97	PL B391 191	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ACHASOV	97	PR D55 2663	N.N. Achasov <i>et al.</i>	(NOVM)
BERTIN	97C	PL B408 476	A. Bertin <i>et al.</i>	(OBELIX Collab.)
CLEGG	94	ZPHY C62 455	A.B. Clegg, A. Donnachie	(LANC, MCHS)
CLEGG	90	ZPHY C45 677	A.B. Clegg, A. Donnachie	(LANC, MCHS)
BISELLO	89	PL B220 321	D. Bisello <i>et al.</i>	(DM2 Collab.)
DUBNICKA	89	JPG 15 1349	S. Dubnicka <i>et al.</i>	(JINR, SLOV)
GESHKEN...	89	ZPHY C45 351	B.V. Geshkenbein	(ITEP)
ANTONELLI	88	PL B212 133	A. Antonelli <i>et al.</i>	(DM2 Collab.)
DIEKMAN	88	PRPL 159 101	B. Diekmann	(BONN)
FUKUI	88	PL B202 441	S. Fukui <i>et al.</i>	(SUGI, NAGO, KEK, KYOT+)
DONNACHIE	87B	ZPHY C34 257	A. Donnachie, A.B. Clegg	(MCHS, LANC)
ATKINSON	86B	ZPHY C30 531	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
ATKINSON	85B	ZPHY C26 499	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
ERKAL	85	ZPHY C29 485	C. Erkal, M.G. Olsson	(WISC)
ABE	84B	PRL 53 751	K. Abe <i>et al.</i>	
KURDADZE	83	JETPL 37 733	L.M. Kurdadze <i>et al.</i>	(NOVO)
		Translated from ZETFP	37 613.	
ATKINSON	82	PL 108B 55	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
BUON	82	PL 118B 221	J. Buon <i>et al.</i>	(LALO, MONP)
CLELAND	82B	NP B208 228	W.E. Cleland <i>et al.</i>	(DURH, GEVA, LAUS+)
CORDIER	82	PL 109B 129	A. Cordier <i>et al.</i>	(LALO)
DEL COURT	82	PL 113B 93	B. Delcourt <i>et al.</i>	(LALO)
ASTON	81E	NP B189 15	D. Aston	(BONN, CERN, EPOL, GLAS, LANC+)
DEL COURT	81B	Bonn Conf. 205	B. Delcourt	(ORSAY)
Also	82	PL 109B 129	A. Cordier <i>et al.</i>	(LALO)
DIBIANCA	81	PR D23 595	F.A. di Bianca <i>et al.</i>	(CASE, CMU)
ASTON	80	PL 92B 215	D. Aston	(BONN, CERN, EPOL, GLAS, LANC+)

BACCI	80	PL 95B 139	C. Bacci <i>et al.</i>	(ROMA, FRAS)
BIZOT	80	Madison Conf. 546	J.C. Bizot <i>et al.</i>	(LAJO, MONP)
KILLIAN	80	PR D21 3005	T.J. Killian <i>et al.</i>	(CORN)
ATIYA	79B	PRL 43 1691	M.S. Atiya <i>et al.</i>	(COLU, ILL, FNAL)
BECKER	79	NP B151 46	H. Becker <i>et al.</i>	(MPIIM, CERN, ZEEM, CRAC)
LANG	79	PR D19 956	C.B. Lang, A. Mas-Parareda	(GRAZ)
MARTIN	78C	ANP 114 1	A.D. Martin, M.R. Pennington	(CERN)
COSTA...	77B	PL 71B 345	B. Costa de Beauregard, B. Pire, T.N. Truong	(EPOL)
FROGGATT	77	NP B129 89	C.D. Froggatt, J.L. Petersen	(GLAS, NORD)
ALEXANDER	75	PL 57B 487	G. Alexander <i>et al.</i>	(TELA)
BALLAM	74	NP B76 375	J. Ballam <i>et al.</i>	(SLAC, LBL, MPIM)
CONVERSI	74	PL 52B 493	M. Conversi <i>et al.</i>	(ROMA, FRAS)
SCHACHT	74	NP B81 205	P. Schacht <i>et al.</i>	(MPIM)
DAVIER	73	NP B58 31	M. Davier <i>et al.</i>	(SLAC)
EISENBERG	73	PL 43B 149	Y. Eisenberg <i>et al.</i>	(REHO)
HYAMS	73	NP B64 134	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
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		Translated from YAF 65 158.		
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FRAEBETTI	01	PL B514 240	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
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EDWARDS	00A	PR D61 072003	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
ABELE	99C	PL B450 275	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
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KULZINGER	99	EPJ C7 73	G. Kulzinger <i>et al.</i>	
ANTONELLI	98	NP B517 3	A. Antonelli <i>et al.</i>	(FENICE Collab.)
BELOZEROVA	98	PPN 29 63	T.S. Belozerova, V.K. Henner	
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BARNES	97	PR D55 4157	T. Barnes <i>et al.</i>	(ORNL, RAL, MCHS)
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		Translated from YAF 55 1896.		
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BRAU	88	PR D37 2379	J.E. Brau <i>et al.</i>	JP
CASTRO	88	Preprint LAL-88-58	A. Castro <i>et al.</i>	(DM2 Collab.)
CLEGG	88	ZPHY C40 313	A.B. Clegg, A. Donnachie	(MCHS, LANC)
ASTON	87	NP B292 693	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
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FRENKIEL	72	NP B47 61	P. Frenkiel <i>et al.</i>	(CDEF, CERN)
ALVENSLEB...	71	PRL 26 273	H. Alvensleben <i>et al.</i>	(DESY, MIT) G
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